



Who likes de Sitter?

Ulf Danielsson
Uppsala University



Who likes de Sitter?

Ulf Danielsson
Uppsala University

History

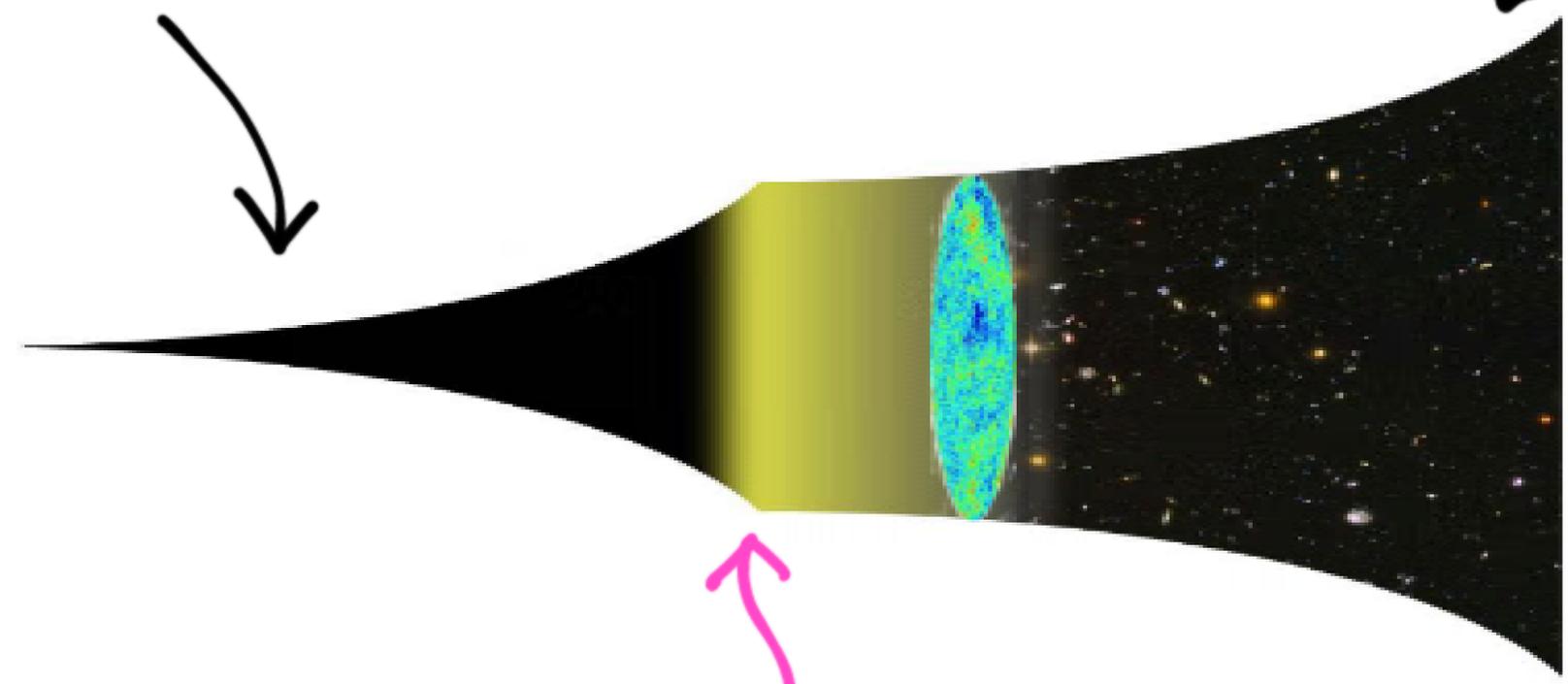
$\lesssim 1998$: String theory \Rightarrow ~~dS~~

1998 \sim 2003: ?

$\gtrsim 2003$: String theory \Rightarrow lots of dS !

≈ 60 e-folds

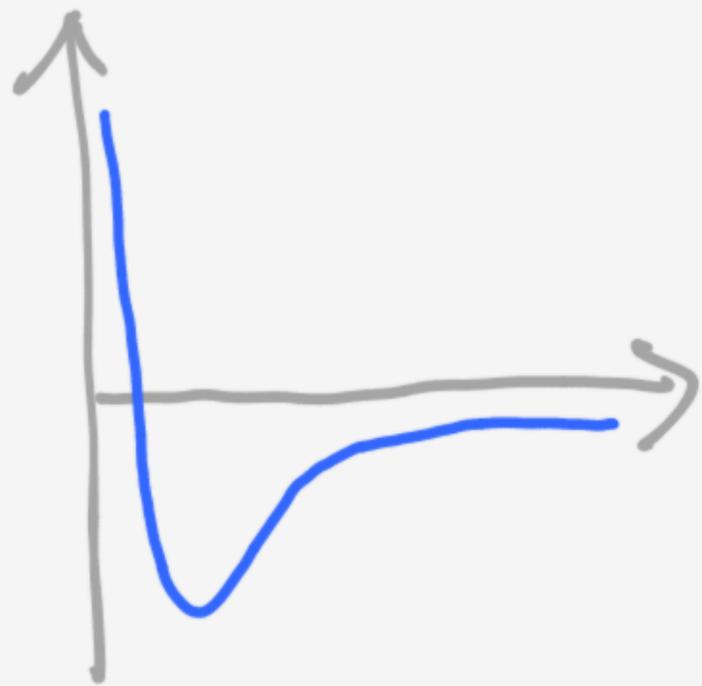
\approx few e-folds



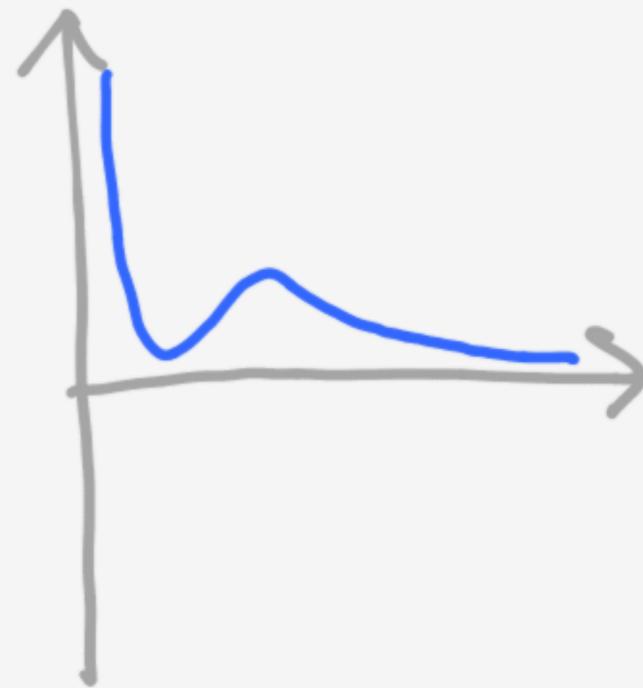
Big Bang!

Explicit ~~SUSY~~... (through uplift)

Add \bar{D} (actually $D + \bar{D}$) to stabilized SUSY AdS...

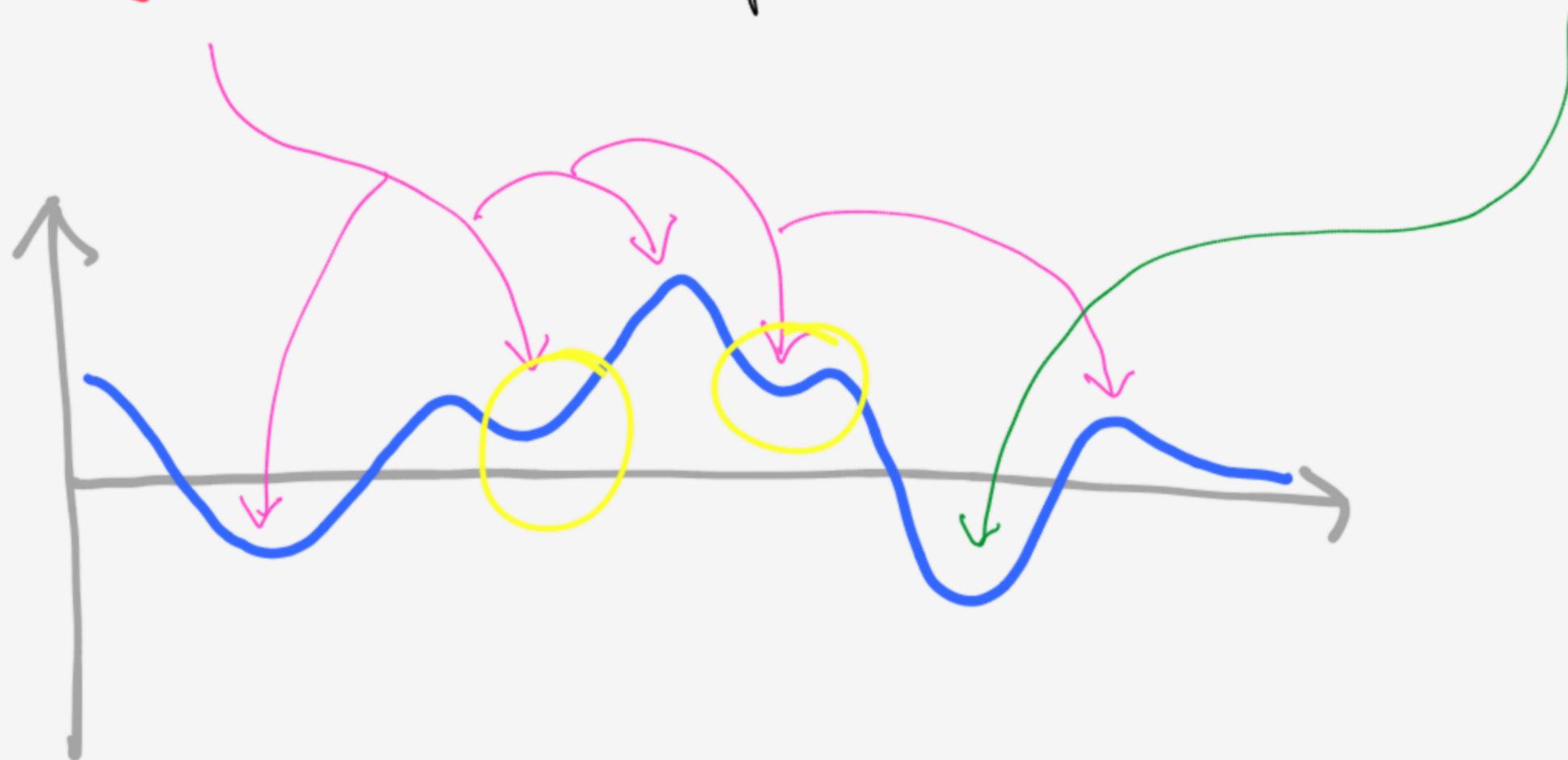


\longrightarrow
KKLT



Spontaneous ~~SUSY~~... (at tree level)

Consider ~~SUSY~~ vacua in potential with SUSY
vacua...



Rest of talk

* String theory on $T^6/\mathbb{Z}_2 \times \mathbb{Z}_2$

* Uplifting and a generic problem

* Conclusions...

String theory on $T^6/\mathbb{Z}_2 \times \mathbb{Z}_2$

Orbifolding $\rightarrow \mathcal{N} = 2$

in and the orientifold $\rightarrow \mathcal{N} = 1$

Can be interpreted as

IIA with O6 or IIB with O3/O7

We will focus on IIA ...

Orbifold:

$$\theta_1: (y^1, y^2, y^3, y^4, y^5, y^6) \rightarrow (-y^1, -y^2, y^3, -y^4, -y^5, y^6)$$

$$\theta_2: (y^1, y^2, y^3, y^4, y^5, y^6) \rightarrow (-y^1, y^2, -y^3, -y^4, y^5, -y^6)$$

$$\theta_1\theta_2: (y^1, y^2, y^3, y^4, y^5, y^6) \rightarrow (y^1, -y^2, -y^3, y^4, -y^5, -y^6)$$

... which splits;

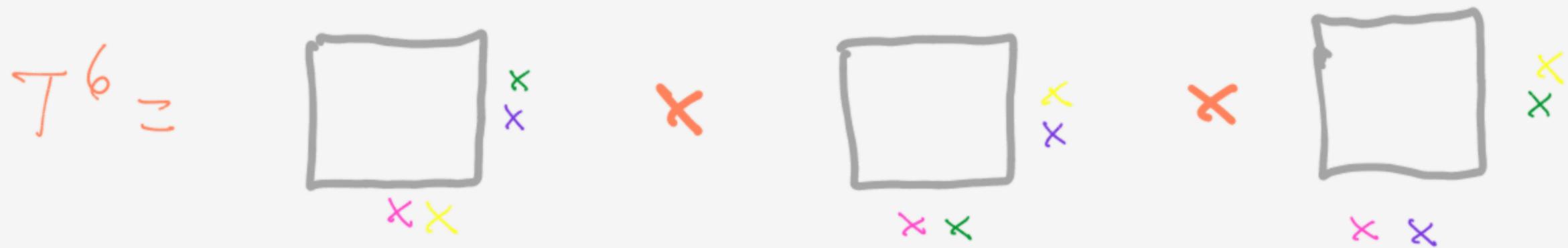
$$T^6 \cong \square_{y_1} \times \square_{y_2} \times \square_{y_3}$$

The diagram shows three squares representing tori. The first square is labeled y_1 at the bottom and y_4 at the top right. The second square is labeled y_2 at the bottom and y_5 at the top right. The third square is labeled y_3 at the bottom and y_6 at the top right. Orange 'x' symbols are placed between the squares to indicate a direct product.

Orientifold at fixed point of

$$\sigma: (y^1, y^2, y^3, y^4, y^5, y^6) \rightarrow (y^1, y^2, y^3, -y^4, -y^5, -y^6)$$

... and at $(\sigma\theta_1), (\sigma\theta_2), (\sigma\theta_1, \theta_2)$



Consider compactification on:

$$ds_{10}^2 = z^{-2} ds_4^2 + \rho ds_6^2$$

with $z = \rho^3 e^{-\phi}$ for Einstein frame...

We get $V = V_R + V_H + \sum V_q + V_{G/D}$ ($q = 0, 2, 4, 6$)

where $V_R \sim -R_6 \sim \rho^{-1} z^{-2}$ $V_H \sim \rho^{-3} z^{-2}$

$$V_q \sim F_q^2 \sim z^{-4} \rho^{3-q} \quad V_{G/D} \sim z^{-3}$$

Basic result:

dS requires $V_R > 0$ $V_{G/D} < 0$

Are there (meta)-stable solutions without tachyons?

Not enough to consider just ρ and z ...

For instance!

$$ds_6^2 = \rho \left(\sigma^3 ds_3^{\parallel 2} + \sigma^{-3} ds_3^{\perp 2} \right)$$

... with respect to G -plane.

Tachyons will in general turn up when σ involved.

To be more explicit one can make use of

$SU(3)$ -structures

Decompose: $SO(6) \supset SU(3)$

Globally defined (singlets under $SU(3)$):

Real 2-form: \downarrow

Holomorphic 3-form: Ω

For US:

$$dJ = \frac{3}{2} \operatorname{Im}(W_1 \Omega) + W_3$$

Torsion-
classes!

$$d\Omega = W_1 \wedge \nu + W_2 \wedge \nu$$

ν and Ω specify metric and:

$$R_6 = -\frac{15}{2} |W_1|^2 + \frac{1}{2} |W_2|^2 + \frac{1}{2} |W_3|^2$$

Ansatz for $10D$ (not completely general) can be written:

$$F_0 \sim f_0, \quad F_2 \sim f_1 \downarrow + f_2 W_2, \quad F_4 \sim f_3 \downarrow \downarrow + f_4 W_2 \uparrow \downarrow$$

$$F_6 \sim f_5 \text{vol}_6, \quad H \sim f_6 \Omega_R + f_7 W_3, \quad \textcircled{\gamma \sim f_8 \Omega_R + f_9 W_3}$$

source

U.D., S. Haque, G. Shiu, T.v. Riet 0907.2041

U.D., P. Koerber, T.v. Riet, 1003.3590

U.D., S. Haque, P. Koerber, G. Shiu, T.v. Riet, T. Wrase

Solve equations of motion and find...

Always at least

one tachyon!

To be more systematic...

Unified description for IIA/B in $D=4, N=1, SU(6)$

with S, T^I, U^I $I=1,2,3$, i.e., $2 \times 3 = 6$ moduli.

We have;

$$K = -\ln(-i(S-\bar{S})) - \sum_{I=1}^3 \ln(-i(T_I - \bar{T}_I)) - \sum_{I=1}^3 \ln(-i(U_I - \bar{U}_I))$$

$$W = W(S, T^I, U^I)$$

With isotropy, $T_1 = T_2 = T_3 = T$, $U_1 = U_2 = U_3 = U$

we have:

$$S = i \alpha \sigma^{-9/2} + c_1 \quad T = i \alpha \sigma^{9/2} + c_2$$
$$U = i \rho + b$$

axions



All contributions to potential V come

from particular terms in superpotential W , ...

In IIA: $1 \sim F_6, U \sim F_4, U^2 \sim F_2, U^3 \sim F_0$

$\Rightarrow 1 + 3 + 3 + 1 = 8$ fluxes

$S, T \sim H_3 \Rightarrow 1 + 3 = 4$ fluxes

$SU, TU \sim \omega$ (metric) $\Rightarrow 3 + 9 = 12$

$\omega^2 = 0$ (Jacobi constraint) to get twisted tori $\Rightarrow 12 \rightarrow 6$

$\omega^2 \neq 0$ corresponds to having KK-monopole

To appear (U.D., J.P. Derendinger, G. Dibitetto, A. Guarino)

$$V_{STU} = V_{SU(3)} \quad (\text{even when } \omega^2 \neq 0)$$

... using

$$W^{(\omega)} = b_I{}^I SU_I + c_I{}^{IJ} T_{IJ} U_J$$

3 + 9 = 12

Can be extended to G_2 -structures in M-theory,
which also includes some non-geometric fluxes.

Key question:

Are there metastable dS?

It is convenient to use SUSY-breaking variables as parameters...

$$D_x W = A_x + iB_x$$

with $x = 1, \dots, N/2$ and N as # of real fields

We have! $S, T, U \rightarrow N = 2 \times 3 = 6$

$$S, T^{\mathbb{I}}, U^{\mathbb{I}} \rightarrow N = 2 \times 7 = 14$$

Replace N of M fluxes...

N SUSY-breaking + $(M-N)$ fluxes

↑
quadratic
in eq. of motion

↑
linear
in eq. of motion

Schematically;

$$(DW)^2 + F DW = 0$$

$DW = 0 \Rightarrow$ solution

Move candidate solution to origin of moduli space ...

$$S_0 = T_0 = U_0 = \vec{0}$$

Adjust N parameters for solution ...

$M - N$ left to adjust to assure no tachyons ...

Must have $M - N - N = M - 2N \geq 0$

Isotropic; $N = 6, M = 8$

Non isotropic; $N = 14, M = 24 (18)$

Not enough!

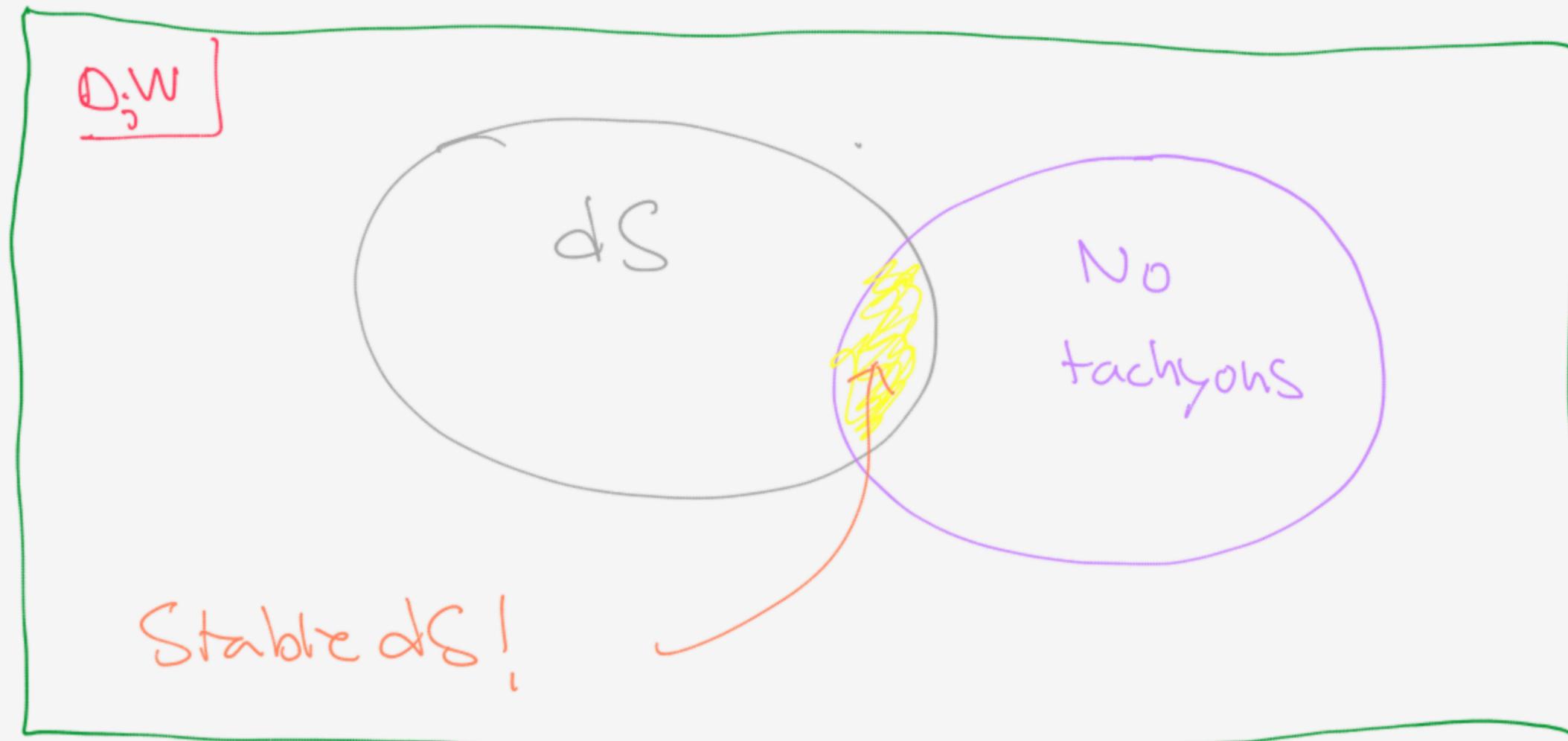
Need more fluxes!

	II B	II A	Iso	Non-iso
U^1	U^1 U^2 U^3	U^1 U^2 U^3	-	1
U^2			-	3
U^3			-	3
S	S SU SU^2 SU^3	S SU SU^2	-	1
SU			-	3
SU^2			-	3
SU^3	-	1	1	
T	T TU TU^2 TU^3	T TU TU^2	-	3
TU			-	9
TU^2			-	9
TU^3	-	3	3	

In 1212,4984 (G, Dibitetto, U.D.)

IIB with $Q = IIA$ with Q and R

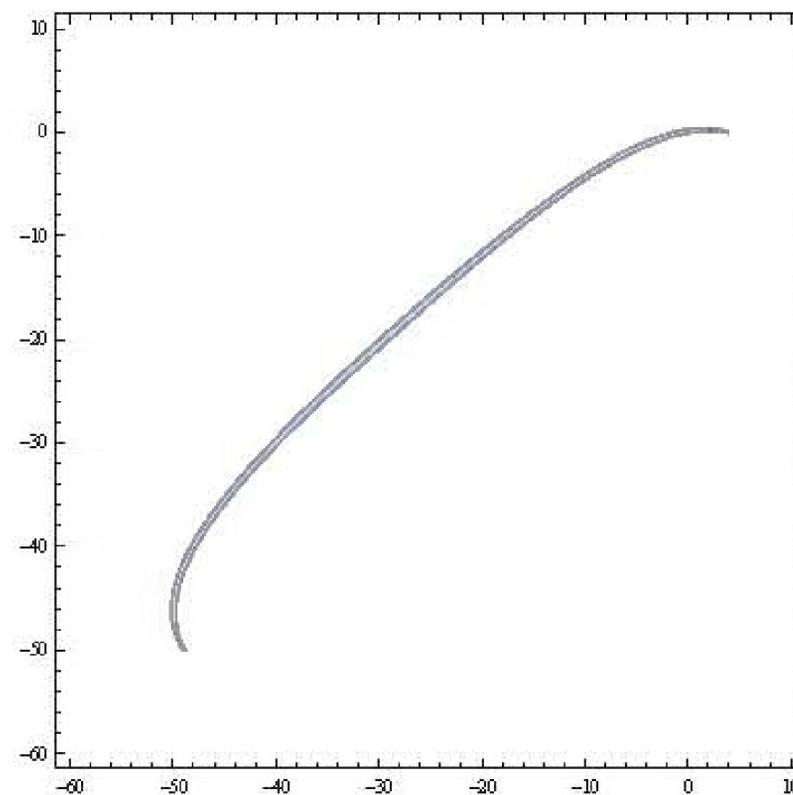
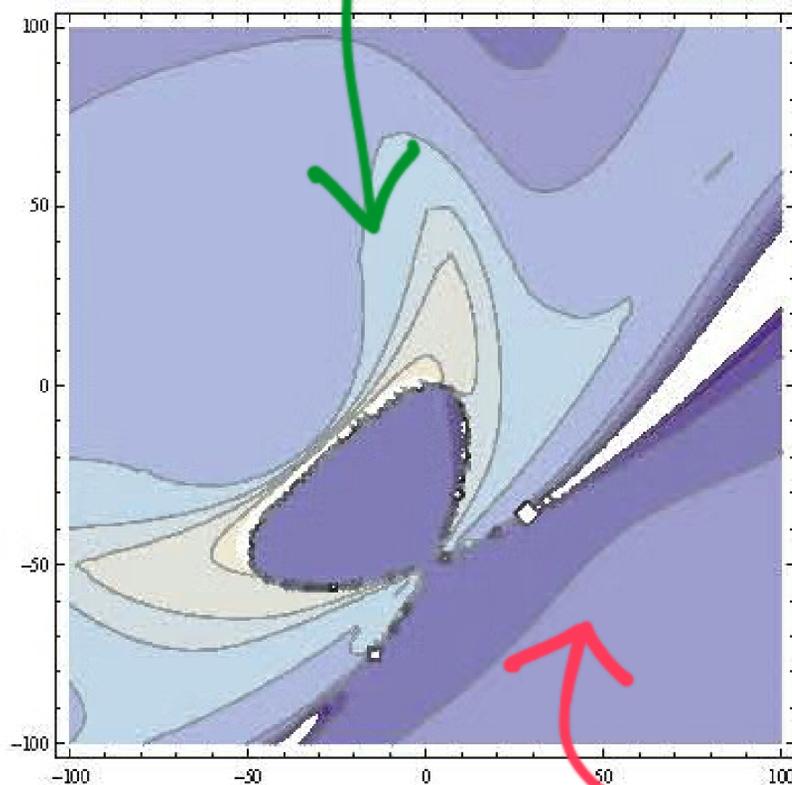
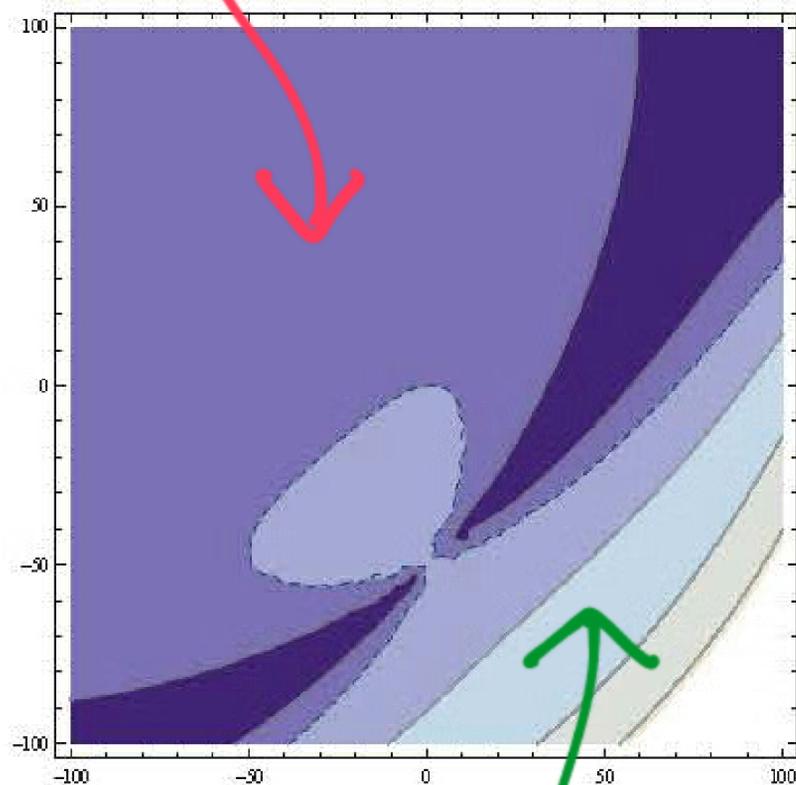
was studied with $M = 12$ and solutions found...



6 fluxes
fixed for
solution...

AdS...

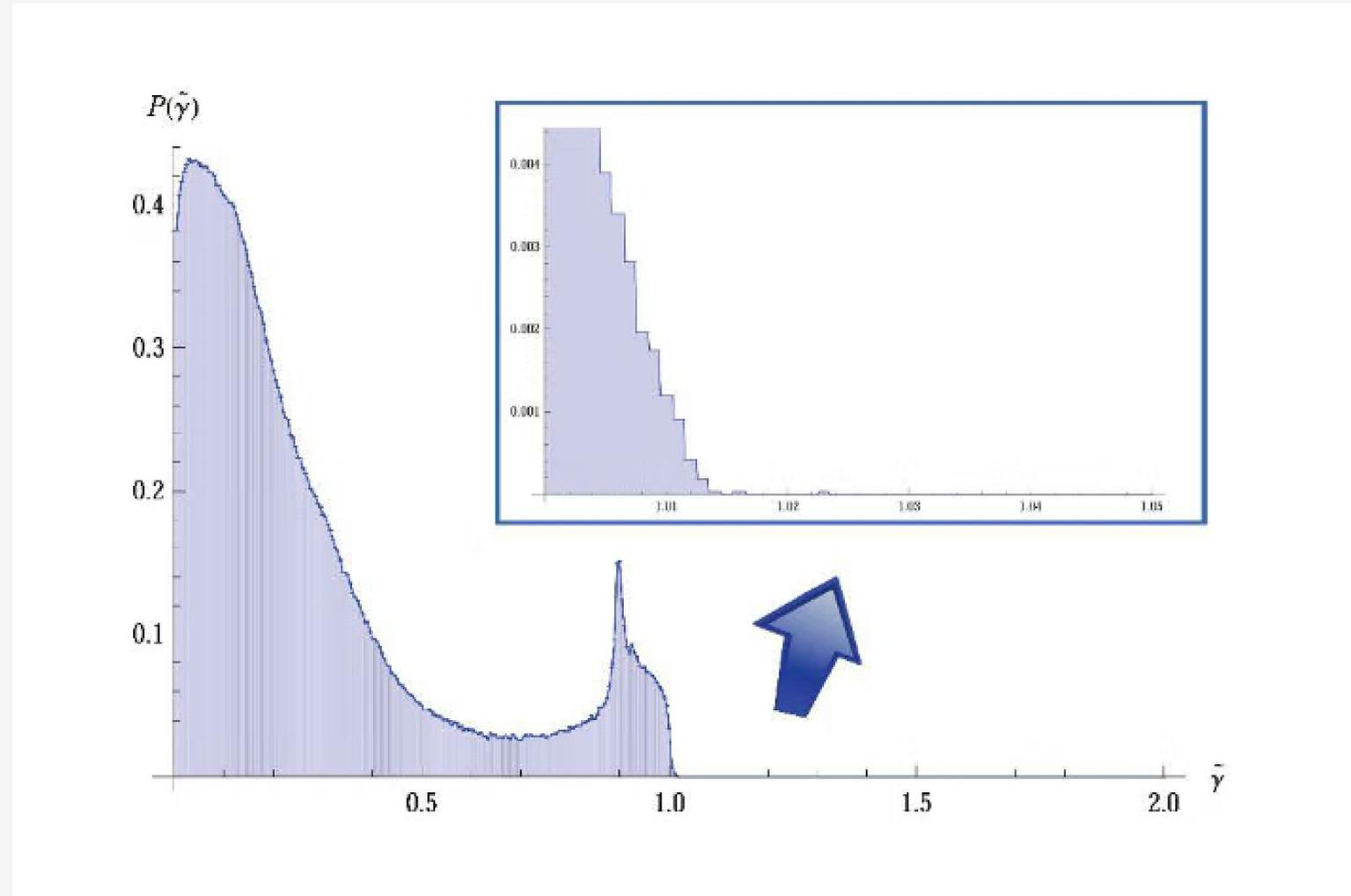
No tachyons!



dS!

Tachyons...

Good overlap!
(magnified)



... much fewer dS with large energy than expected from RMT ...

\Rightarrow conspiracy?

In 1301.7073 (J. Blabach, G. Dibitetto, U.D)

The full nonisotropic case was studied for

$$\text{II B} + \text{T, TU} = \text{II A} + \text{SU}^2, \text{SU}^3$$

$$N = 16 + 3 + 9 = 24 + 3 + 1 = 28 = 2 \times 14$$

... again solutions were found, Fully stable in

all 14 directions. Genetic algorithm used,

but see 1406.4866 (Kallosh, Linde, Verma, Wrase)

	Sol. 1		Sol. 2	
$V_0 \equiv V(\Phi_0)$	8.85×10^{-6}		1.58×10^{-5}	
$\tilde{\gamma} \equiv \frac{ DW ^2}{3 W ^2}$	1.00008		1.00012	
Normalised masses (m^2/V_0)	1.85148×10^6	1.78128×10^6	262778	25970
	1.31064×10^6	1.27212×10^6	160140	15360
	113890	94187.9	24219.8	11296.
	23907.3	11397.5	9273.53	7282.9
	5290.32	1478.37	4155.21	1745.3
	1353.35	607.799	1306.57	343.39
	17.9045	2.85612×10^{-3}	24.7202	11.450

The real values...

A_{U_1} B_{U_1}	0.437796	-0.223916	-0.41977		
A_{U_2} B_{U_2}	-0.10521	-0.716205	-0.074538		
A_{U_3} B_{U_3}	-0.110376	-0.952221	-0.597945		
a_0 b_0	0.249727	1.26278	-0.0809612		
$a_1^{(i)}$ $b_1^{(i)}$	-0.385558	-1.00688	0.763615		
	3.24742	3.23091	-0.129007		
	-3.42277	-1.65416	-0.258342		
$a_2^{(i)}$ $b_2^{(i)}$	-0.678952	0.269654	-1.69969		
	4.77196	-2.82032	0.141055		
	-4.74382	3.85504	2.04863		
a_3 b_3	-0.626634	0.244097	-0.0281293		
$c_0^{(i)}$	0.210031		0.22		
	0.680146		-0.		
	-0.154349		0.28		
$c_1^{(ij)}$	0.934579	0.932678	0.134729	1.00141	-1.
	1.3931	0.978188	0.351523	-0.335711	-0.
	-1.12313	-0.707561	-0.313208	-0.628606	0.

Is it OK to have nongeo fluxes?

At least some can be understood in M_{111}

To appear (U.D., J.P. Derendinger, G. Dibitetto, A. Guarino)

M_{111} and work in progress,

Uplifting and a generic problem

So far all branes and orientifolds are smeared...
OK if SUSY... No forces, does not matter how
you put things on top of each other. But when
non-SUSY?

No force!

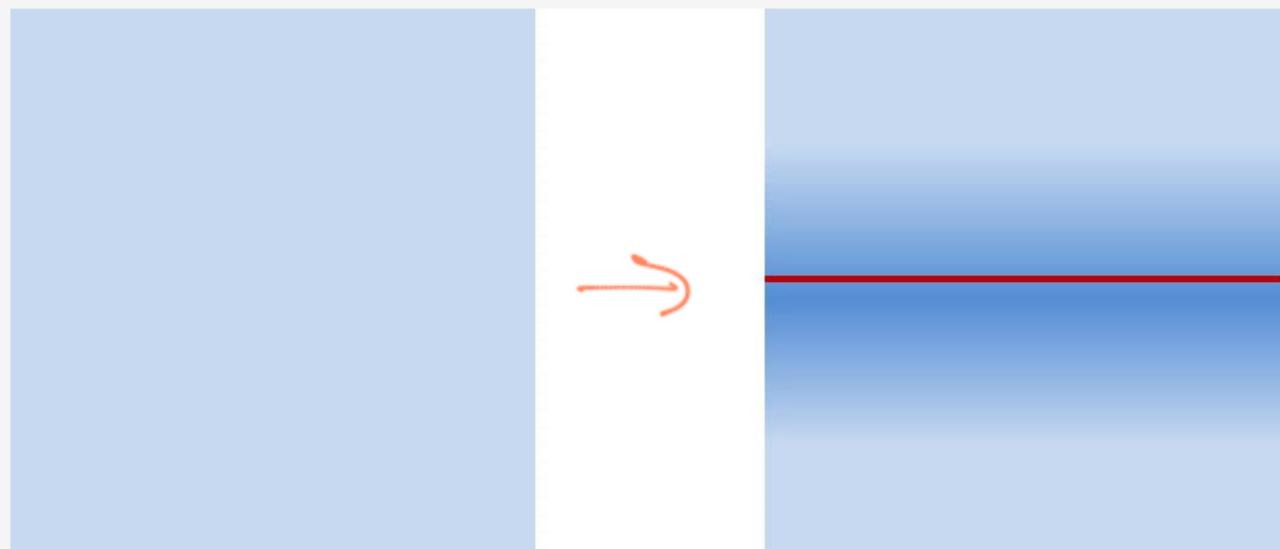


Force...



Lift to dS by adding antibrane to
SUSY-flux configuration...

Typically what happens



... flux attracted to antibrane and piles up!

Selection of papers...

McGuirk, Shiu, Sumitomo 0910,4581

Bena, Grana, Halmagyi 0912,3519

Dymarshy 1102,1734

Bena, Biebold, Grana, Halmagyi, Massai 1106.6165

Bena, Junghans, Kuperstein, van Riet, Wrase 1205,1798

Bena, Grana, Kuperstein, Massai 1206.6369

Bena, Grana, Kuperstein, Massai 1212,4828

Dymarshy, Massai 1310,0015

...

Get singularity in H^2 ...

For $\overline{D6}$ in IIA this can be proved using ...

$$dF_2 = F_0 H_3 + Q_8$$

... sources F_2 , which tells H_3 how to move ...

$$d * H_3 = - * F_2 F_0$$

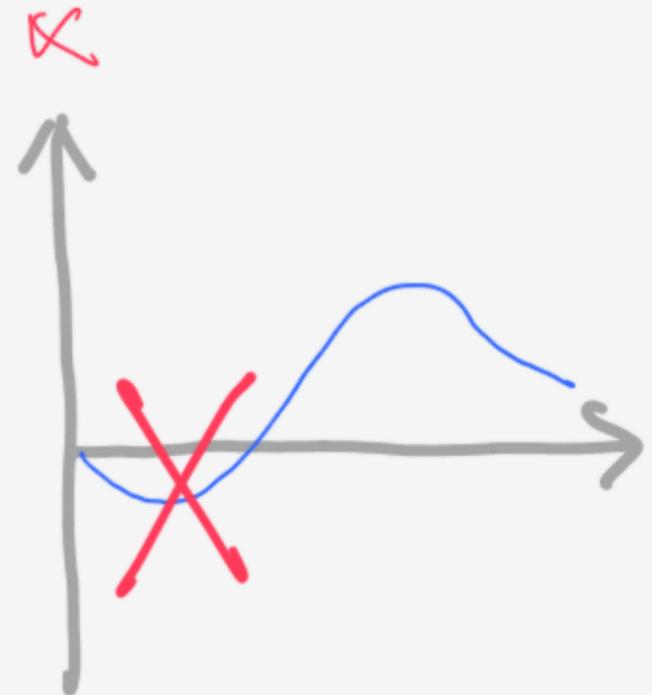
Using $H_3 = \int F_0 *_{3} \mathbb{1}$ and $F_2 = *_{3} dx$
 one gets at $x' = 0$!

Blabach, U.D.,
 Jonghans, van Riet,
 Wrase, Zagerman
 (105,4879, 111,2605)

$$\text{sign } x'' = \text{sign } x$$



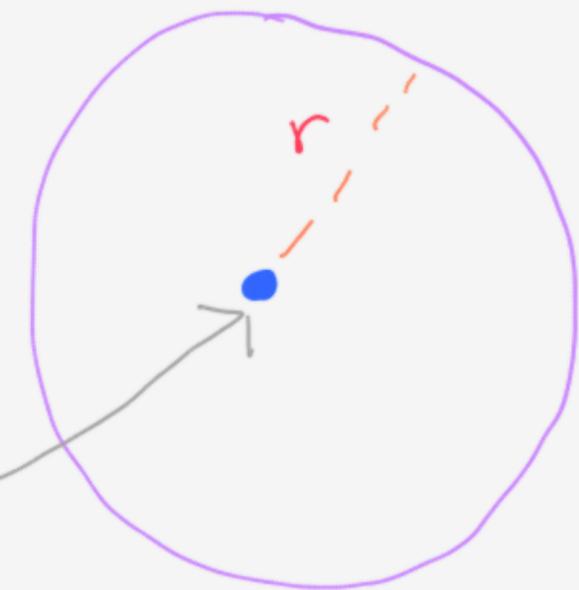
$$x(0) \neq 0 \Rightarrow \text{singularity!}$$



... in case of $\overline{03}$ in KS you can use the
 results of 1301.5647 (F. Gautason, D. Jungkars, M. Zagermann)
 further developed in to appear (J. Blöbäck, U. D., D. Jungkars,
 T. van Riet, S. C. Vargas) to the form!

$$0 \cong S_{DBI}^0 + S_{WZ} + \oint_r \mathcal{B}$$

lives on source



independent
of r

$$M_{ADM} \neq 0 \Rightarrow \oint_r \vec{B} \neq 0 \Rightarrow S_{WZ} \neq 0 \Rightarrow \kappa(0) \neq 0$$

\Rightarrow singularity!

Note!

Not singularity in field at its source.

Rather, singularity in fluid attracted by source!

The dS-vacuum is always unstable;

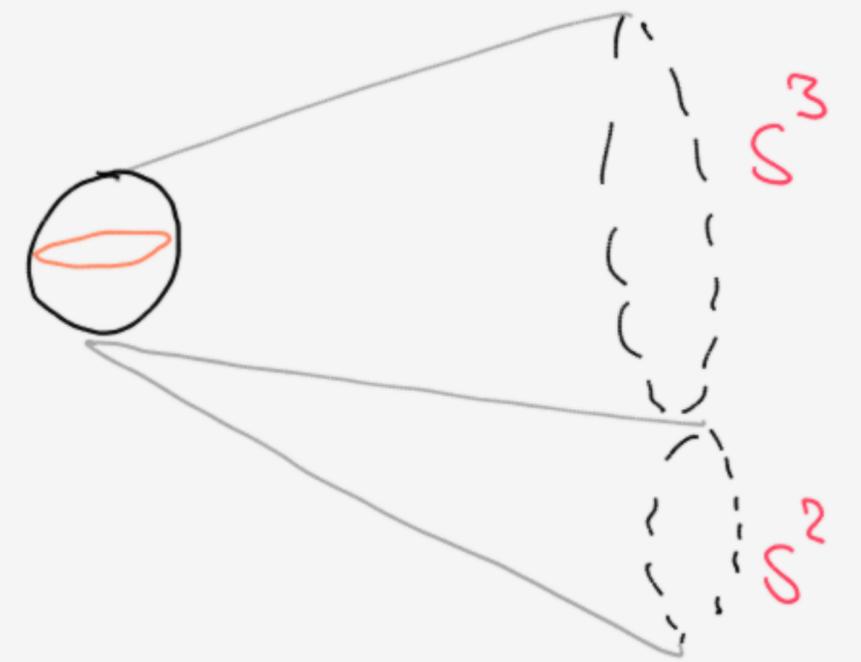
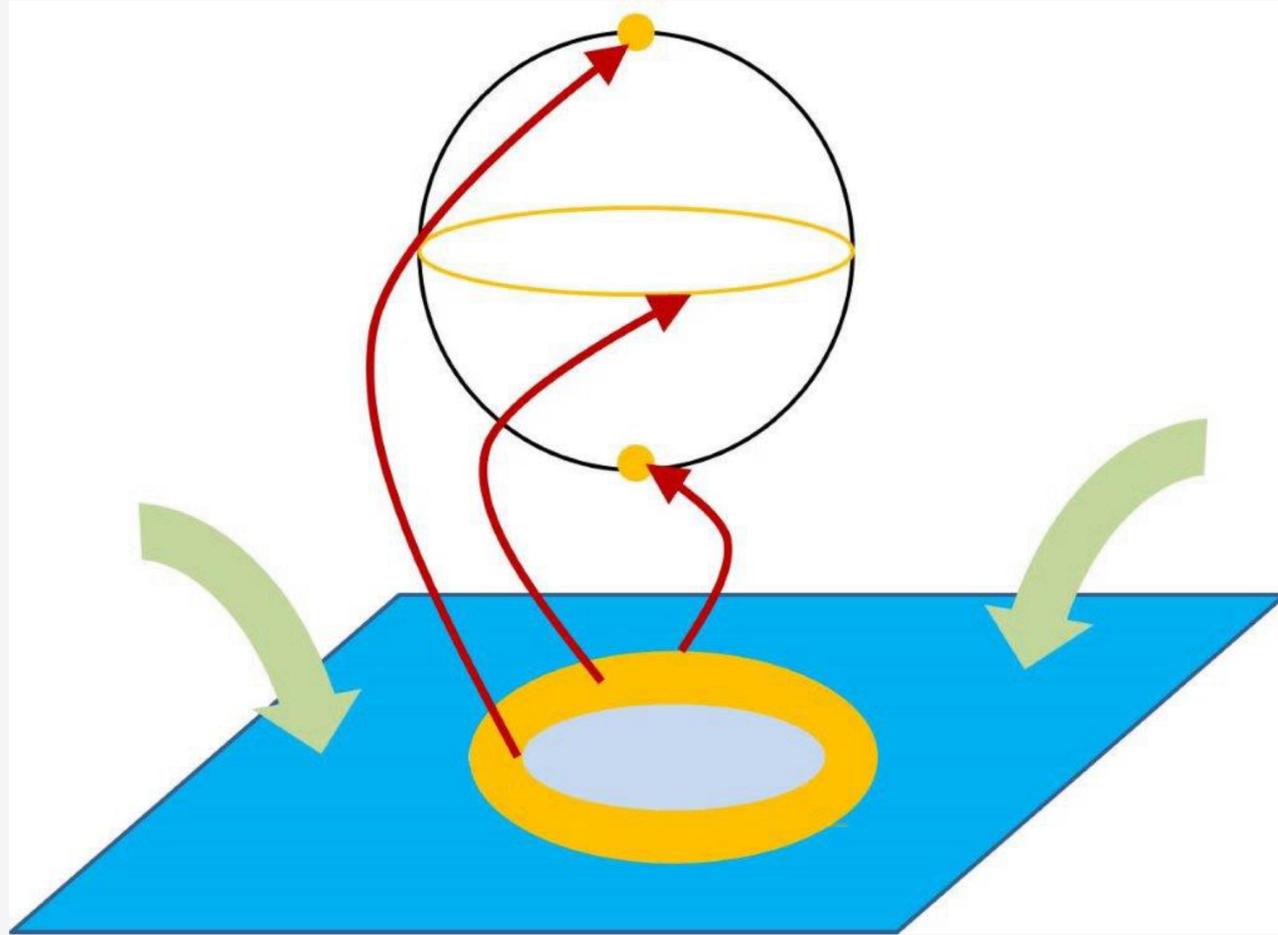


KPV 0112197: barrier against $F_3 \wedge H_3$ annihilation,
with antibrane,

Claim! barrier lowered when flux
piled up on top of antibrane.

1202.1132 (J. Blabach, G.D., T. van Riet)

$\overline{D3}$ can end on NS5, Formation of NS5-bubbles
annihilate the $\overline{D3}$...



Klebanov-Strassler

Just like Schwinger effect

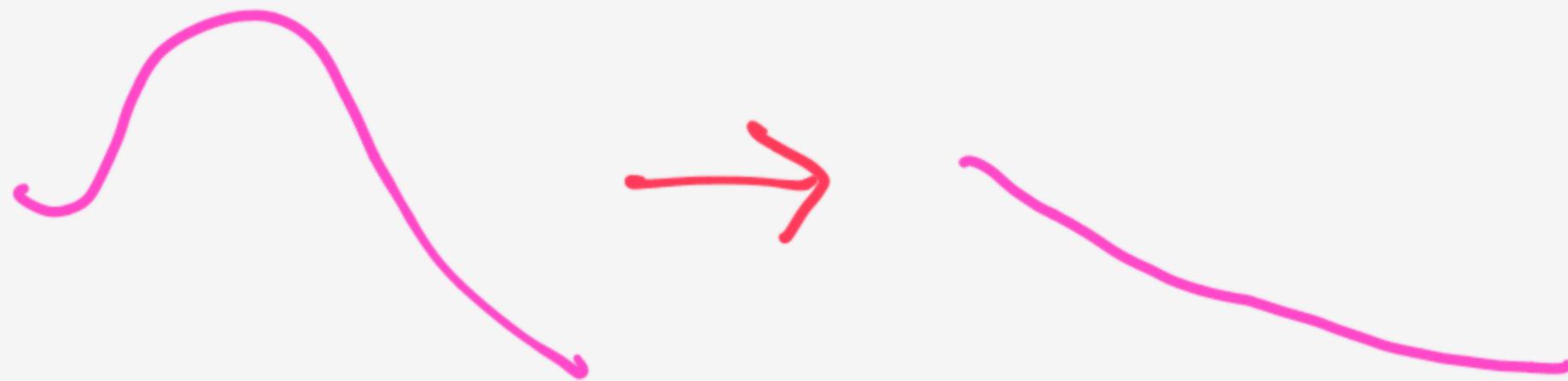


Gain energy through tunneling. Barrier lowered by higher \bar{E} .

Action for NS5:

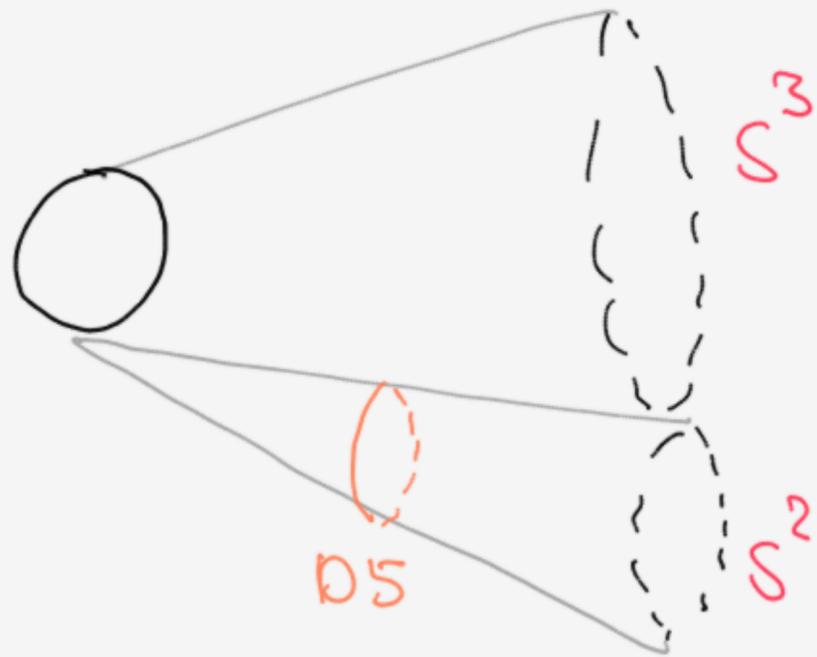
$$S = \frac{\mu_5}{g_s^2} \int d^6\xi \sqrt{-G_{11}(G_I + F)} - \mu_5 \int B_6$$

Barrier lowered, goes away, when $H \sim \ast d B_6$ grows...



Coupled polarization of $\overline{D3} \rightarrow D5$ prevent
build up?

Calculations suggest this does not happen...



Singularity resolved by time dependence?

Flux annihilated as it is piled up...

Time scale?

Conclusion: It remains a challenge to prove existence of longlived dS with anti-brane uplift.

Conclusions

Spontaneous ~~SUSY~~ (smeared)

... difficult without non-geo...

Explicit ~~SUSY~~

... instabilities against annihilation

A smaller landscape ...?

... string theory is more predictive!

A different landscape ...?

... do not focus on (meta)-stable dS, but

consider instead

Quintessence

Proof of principle ...

In 1310.8300 (J. Blåbäck, U.D., and G. DiBiello)

The presence of accelerated (unstable) solutions

in geometric IIA is investigated ...

~ 2-3 efoldings

Can one do better?