

# Maximally Supersymmetric Solutions in Supergravity

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# Introduction

- ▶ Supersymmetric solutions in supergravity are well studied, many are classified.
- ▶ They are believed to uplift to solutions of the full string theory and are related to string solitons, i.e. non-perturbative string theory.
- ▶ Moreover anti-de Sitter solutions feature prominently in AdS/CFT correspondence.
- ▶ Here: **Maximally supersymmetric solutions.**
  1. Classification of all background space-times for all gauged and deformed supergravity theories in  $D \geq 3$  space-time dimensions.
  2. AdS solutions and their moduli spaces.

# Outline / Results

A) Set the stage to discuss solutions in a generic framework.

B) Solutions *without* fluxes:

- ▶ Only two possible cases:

$Mink_D$  and  $AdS_D$   
(only gauged / deformed SUGRAs) 

C) Solutions *with* (non-trivial) fluxes:

- ▶ Exist only for a small class of theories.
  - ▶ Solutions coincide with those of the corresponding ungauged theories.
  - ▶ For these theories all solutions are known and classified.
- Exhaustive list of solutions.

D)  $AdS_D$  solutions

- ▶ Characterize gaugings.
- ▶ Moduli spaces for theories with coset scalar field space.

# Maximally supersymmetric backgrounds

- ▶ Classical solutions / backgrounds for which

$$\langle \delta_\epsilon B \rangle = \langle \delta_\epsilon F \rangle = 0.$$

 SUSY variations of  
bosonic / fermionic fields

- ▶ Here: bosonic solutions, i.e.  $\langle F \rangle = 0$

$$\Rightarrow \langle \delta_\epsilon B \rangle \sim \langle F \rangle = 0.$$

$\Rightarrow$  only remaining condition:  $\langle \delta_\epsilon F \rangle = 0$

- ▶ indep. supersymmetry paramters  $\epsilon \leftrightarrow$  preserved supercharges

# Supersymmetry variations

- ▶ gravitini:

$$\delta\psi_\mu^i = D_\mu\epsilon^i + (\mathcal{F}_{0\mu})_j^i \epsilon^j + A_{0j}^i \gamma_\mu \epsilon^j$$

fluxes from the gravity multiplet      gaugings / deformations

covariant derivative:  $D_\mu\epsilon^i = \nabla_\mu\epsilon^i + (\mathcal{Q}_\mu)_j^i \epsilon^j$

- ▶ spin-1/2 fermions from the gravity multiplet (“dilatin”):

$$\delta\chi^a = (\mathcal{F}_1)_i^a \epsilon^i + A_{1i}^a \epsilon^i$$

- ▶ spin-1/2 fermions from other multiplets (“gaugini”, “hyperini”, ...):

$$\delta\lambda^s = (\mathcal{F}_2)_i^s \epsilon^i + A_{2i}^s \epsilon^i$$

fluxes from other multiplets

# Solutions without fluxes (1)

- ▶ Easiest case: all fluxes vanish:

$$\mathcal{F}_{0\mu} = \mathcal{F}_1 = \mathcal{F}_2 = 0$$

- ▶ Supersymmetric solutions of gauged supergravity without background fluxes need to satisfy the Killing spinor equation:

$$\delta\psi_{\mu}^i = \nabla_{\mu}\epsilon^i + A_{0j}^i\gamma_{\mu}\epsilon^j = 0$$

- ▶ Integrability condition:

$$\left( \frac{1}{4}R_{\mu\nu}{}^{\alpha\beta}\delta_k^i + 2A_{0j}^iA_{0k}^j\delta_{\mu}^{\alpha}\delta_{\nu}^{\beta} \right) \gamma_{\alpha\beta}\epsilon^k = 0$$

- ⇒ Unbroken supersymmetry (without fluxes):  
**Mink<sub>D</sub>** and **AdS<sub>D</sub>** are the only possible solutions.

## Solutions without fluxes (2)

- ▶ Spin-1/2 variations:

$$\delta\chi^a = A_{1i}^a \epsilon^i = 0, \quad \delta\lambda^s = A_{2i}^s \epsilon^i = 0$$

### Algebraic conditions

$$A_0^2 = -\frac{\Lambda}{2(D-1)(D-2)} \mathbb{1}, \quad A_1 = A_2 = 0$$

- ▶ Compare with the potential:

$$V = -c_0 \text{tr}(A_0^\dagger A_0) + c_1 \text{tr}(A_1^\dagger A_1) + c_2 \text{tr}(A_2^\dagger A_2),$$

- ▶ **Mink<sub>D</sub>** ( $\Lambda = 0$ ) solutions exist for all theories.
- ▶ For  $D \leq 7$  most gauged theories admit **AdS<sub>D</sub>** ( $\Lambda < 0$ ) solutions.  
(See second half of this talk.)

[de Alwis, Louis, McAllister, Triendl, Westphal][Louis, Triendl][Louis, SL][Louis, Triendl, Zagermann][Louis, Muranaka]

# Solutions with non-trivial flux (1)

- ▶ For more “interesting” solutions: Allow for non-vanishing flux.
- ▶ Firstly: spin-1/2 variations:

$$\delta\chi^a = (\mathcal{F}_1)_i^a \epsilon^i + A_{1i}^a \epsilon^i = 0, \quad \delta\lambda^s = (\mathcal{F}_2)_i^s \epsilon^i + A_{2i}^s \epsilon^i = 0$$

- ▶ Unbroken supersymmetry:

$$\mathcal{F}_1 = \mathcal{F}_2 = A_1 = A_2 = 0 \quad \Rightarrow \quad \mathcal{F}_{0\mu} = 0$$

- ▶ Generically no non-trivial fluxes possible!

## Two exceptions

1. theories without  $\chi$ 's in the gravitational multiplet.
2. chiral theories with selfdual fluxes.

## Solutions with non-trivial flux (2)

- ▶ Secondly: gravitino variations:

$$\delta\psi_{\mu}^i = \nabla_{\mu}\epsilon^i + (\mathcal{Q}_{\mu})_j^i + (\mathcal{F}_{0\mu})_j^i \epsilon^j + A_{0j}^i \gamma_{\mu}\epsilon^j = 0$$

- ▶ Integrability condition:

$$\left( \frac{1}{4} R_{\mu\nu\rho\sigma} \gamma^{\rho\sigma} \delta_j^i - (\mathcal{H}_{\mu\nu})_j^i + \dots \right) \epsilon^j = 0,$$

where  $\mathcal{H}_{\mu\nu}$  is the field strength corresponding to  $\mathcal{Q}_{\mu}$ .

- ▶ Unbroken supersymmetry:

$$\mathcal{H}_{\mu\nu} = 0 \quad \Rightarrow \quad \mathcal{Q}_{\mu} = A_0 = 0$$

(See also [Hristov,Looyestijn,Vandoren] [Gauntlett,Gutowski] [Akyol,Papadopoulos])

- ▶ The supersymmetry variations take the same form as in the ungauged / undeformed case.

## Solutions with non-trivial flux (3)

SUGRAs with non supersymmetry breaking flux:

dimension	supersymmetry	q	possible flux	classification
$D = 11$	$N = 1$	32	$F^{(4)}$	[Figueroa-O'Farrill,Papadopoulos]
$D = 10$	IIB	32	$F_+^{(5)}$	[Figueroa-O'Farrill,Papadopoulos]
$D = 6$	$N = (2, 0)$	16	$5 \times F_+^{(3)}$	[Chamseddine,Figueroa-O'Farrill,Sabra]
$D = 6$	$N = (1, 0)$	8	$F_+^{(3)}$	[Gutowski,Martelli,Reall]
$D = 5$	$N = 2$	8	$F^{(2)}$	[Gauntlett,Gutowski,Hull,Pakis,Reall]
$D = 4$	$N = 2$	8	$F^{(2)}$	[Tod]

The maximally supersymmetric solutions are classified:

- ▶  $AdS_p \times S^{(D-p)}$  and  $AdS_{(D-p)} \times S^p$ , for  $p$ -form flux.
- ▶ Hpp-wave as Penrose-limit of  $AdS \times S$  solutions.

[Penrose;Gueven;Blau,Figueroa-O'Farrill,Hull,Papadopoulos]

- ▶ Exceptional solutions in  $D = 5$ . [Gauntlett,Gutowski,Hull,Pakis,Reall]

# All maximally supersymmetric solutions

(with non-trivial flux)

dim.	SUSY	q	$AdS \times S$	Hpp-wave	others
$D = 11$	$\mathcal{N} = 1$	32	$AdS_4 \times S^7$ $AdS_7 \times S^4$	KG <sub>11</sub>	-
$D = 10$	IIB	32	$AdS_5 \times S^5$	KG <sub>10</sub>	-
$D = 6$	$\mathcal{N} = (2, 0)$ $\mathcal{N} = (1, 0)$	16 8	$AdS_3 \times S^3$	KG <sub>6</sub>	-
$D = 5$	$\mathcal{N} = 2$	8	$AdS_2 \times S^3$ $AdS_3 \times S^2$	KG <sub>5</sub>	Gödel-like, NH-BMPV*
$D = 4$	$\mathcal{N} = 2$	8	$AdS_2 \times S^2$	KG <sub>4</sub>	-

\* = near-horizon limit of the BMPV [Breckenridge, Myers, Peet, Vafa] black hole

# AdS solutions

- ▶ So far: classification of background *space-time geometries*.
- ▶ Now: focus on *algebraic conditions* and *target space geometry*.
  
- ▶ Most promising: **AdS<sub>D</sub>** (here:  $D \geq 4$ ):
  - ▶ Conditions  $A_0^2 \sim \mathbb{1}$  and  $A_1 = A_2 = 0$  require gauging.
  - ▶ May have non-trivial moduli spaces.

## Motivation: AdS/CFT correspondence

AdS <sub>D</sub> solution with $q$ supercharges	$\leftrightarrow$	SCFT in $(D - 1)$ dim. with $q/2$ supercharges
gauge symmetry	$\leftrightarrow$	global symmetry
moduli space	$\leftrightarrow$	conformal manifold

# The gauged R-symmetry

- ▶ The conditions  $A_0^2 \sim \mathbb{1}$  and  $A_1 = A_2 = 0$  require  $Q_\mu \neq 0$ .
- ⇒ The R-symmetry group  $H_R$  must be gauged by

$$H_R^g \subseteq H_R,$$

where  $H_R^g$  needs to be generated by the vector fields from the gravity multiplet, i.e. the “*graviphotons*”.

- ▶  $H_R^g$  is uniquely determined to be the maximal subgroup of  $H_R$ , s.t.
  - a)  $A_0$  is  $H_R^g$  invariant, i.e.

$$[\mathfrak{h}_R^g, A_0] = 0$$

- b) The decomposition of the  $H_R$ -representation  $r_{GP}$  of the graviphotons w.r.t.  $H_R^g$  contains the adjoint representation of  $H_R^g$ , i.e.

$$r_{GP} \rightarrow \text{ad}_{H_R^g} \oplus \dots$$

# Coset spaces

- ▶ From now on: Focus on theories where the scalar field space is a coset:

$$\mathcal{M}_{scal} = \frac{G}{H}, \quad H = H_R \times H_{mat}$$

- ▶ Gauge group:  $G^g \subseteq G$ , where

$$G^g = G_R^g \times G_{mat}^g,$$

s.t.  $H_R^g$  is the maximal compact subgroup of  $G_R^g$ .

- ▶ Lie algebra of  $G$ :

$$\mathfrak{g} = \mathfrak{h} \oplus \mathfrak{k}$$

Notice:  $[\mathfrak{h}, \mathfrak{k}] \subseteq \mathfrak{k}$ , i.e.  $\mathfrak{k}$  transforms in some  $\mathfrak{h}$  representation.

- ▶ Decompose  $\mathfrak{k}$  into  $\mathfrak{h}_R^g$  irreps:

$$\mathfrak{k} = \bigoplus_i \mathfrak{k}_i, \quad [\mathfrak{h}_R^g, \mathfrak{k}_i] \subseteq \mathfrak{k}_i$$

# The moduli space

- ▶ Flat directions of the potential  $V(\phi)$ :

$$\begin{aligned} \mathfrak{f} &= \{ \delta\phi \in \mathfrak{k} : V(\langle\phi\rangle + \delta\phi) = V(\langle\phi\rangle) \} \\ &= \{ \delta\phi \in \mathfrak{k} : [\delta\phi, \mathfrak{g}_R^g] \subseteq \mathfrak{g}_R^g \} \end{aligned}$$

( $\mathfrak{f}$  is the non-compact part of the normalizer  $N_{\mathfrak{g}}(\mathfrak{g}_R^g)$ )

- ▶ Observation:

$$\mathfrak{f} = \mathfrak{k}_R^g \oplus \mathfrak{k}_0$$

non-compact part of  $\mathfrak{g}_R^g$   
 $\rightarrow$  Goldstone bosons

$\mathfrak{h}_R^g$  singlets in  $\mathfrak{k}$ , i.e.  
 $[\mathfrak{h}_R^g, \mathfrak{k}_0] = 0$   
 $\rightarrow$  span the moduli space

- ▶ The moduli  $\mathfrak{k}_0$  are the non-compact part of  $\mathfrak{g}_0 = \mathfrak{h}_0 \oplus \mathfrak{k}_0$  (which is the centralizer  $C_{\mathfrak{g}}(\mathfrak{g}_R^g)$ ).

$$\Rightarrow \mathcal{M}_{AdS} = \frac{G_0}{H_0}$$

# All $\text{AdS}_D$ solutions for $D \geq 4$

dim.	SUSY	$H_R^g$	$\mathcal{M}_{\text{AdS}}$	
$D = 7$	$\mathcal{N} = 4$	$\text{USp}(\mathcal{N})$	•	[Pernici, Pilch, van Nieuwenhuizen]
	$\mathcal{N} = 2$		•	[Louis, SL, Rüter (to app.)]
$D = 6$	$\mathcal{N} = (1, 1)$	$\text{SU}(2) \otimes \text{SU}(2)$	•	[Romans]
$D = 5$	$\mathcal{N} = 8$	$\text{SU}(4)$	$\frac{\text{SU}(1, 1)}{\text{U}(1)}$	[Günaydin, Romans, Warner] [Louis, SL, Rüter]
	$\mathcal{N} = 6$	$\text{U}(\mathcal{N}/2)$	•	[Ferrara, Porrati, Zaffaroni]
	$\mathcal{N} = 4$		$\frac{\text{SU}(1, p)}{\text{U}(1) \times \text{SU}(p)}$	[Corrado, Günaydin, Warner, Zagermann] [Louis, Triendl, Zagermann]
	$\mathcal{N} = 2$		Kähler	[Louis, Muranaka]
$D = 4$	$\mathcal{N} = 8$	$\text{SO}(\mathcal{N})$	•	[de Wit, Nicolai][Louis, SL, Rüter]
	$\mathcal{N} = 5, 6$		•	[de Wit, Nicolai]
	$\mathcal{N} = 4$		•	[Louis, Triendl]
	$\mathcal{N} = 3$		•	
	$\mathcal{N} = 2$		Kähler	[de Alwis, et al.]
	$\mathcal{N} = 1$		real	[de Alwis, et al.]

→ Perfect agreement with marginal deformations of SCFTs [Cordova, Dumitrescu, Intriligator] 16 / 17

# Summary

- ▶ Systematic classification of all maximally supersymmetric supergravity backgrounds:
  - ▶ no fluxes: only flat space-time and anti-de Sitter.
  - ▶ fluxes: - generically not possible
    - in the gauged case: same solutions as for ungauged theories
    - all solutions are known and classified
- ▶ General recipe for the computation of  $\text{AdS}_D$  moduli spaces for theories with coset scalar manifold.

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Thank You!