

Moduli Stabilisation near the Conifold

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November, 2016

Workshop on Geometry and Physics, Ringberg

Moduli Stabilisation near the Conifold

String Moduli Stabilisation at the Conifold [arXiv:1605.06299](https://arxiv.org/abs/1605.06299)

Ralph Blumenhagen, DH, Florian Wolf

Motivation

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Large field inflation in string theory

Motivation

axionic shift symmetry of an inflaton protects from corrections

$$\mathcal{L}(\phi) = \frac{1}{2}(\partial\phi)^2 - \frac{1}{2}m^2\phi^2 + \sum_{i=1}^{\infty} c_i \phi^{2i} \Lambda^{4-2i}$$



Motivation

axionic shift symmetry of an inflaton protects from corrections

string theory has many axionic moduli

★ **Ramond-Ramond p-forms**

gauge symmetry  shift symmetry

★ **complex structure moduli**

at special points in moduli space

Motivation

axionic shift symmetry of an inflaton protects from corrections

string theory has many axionic moduli

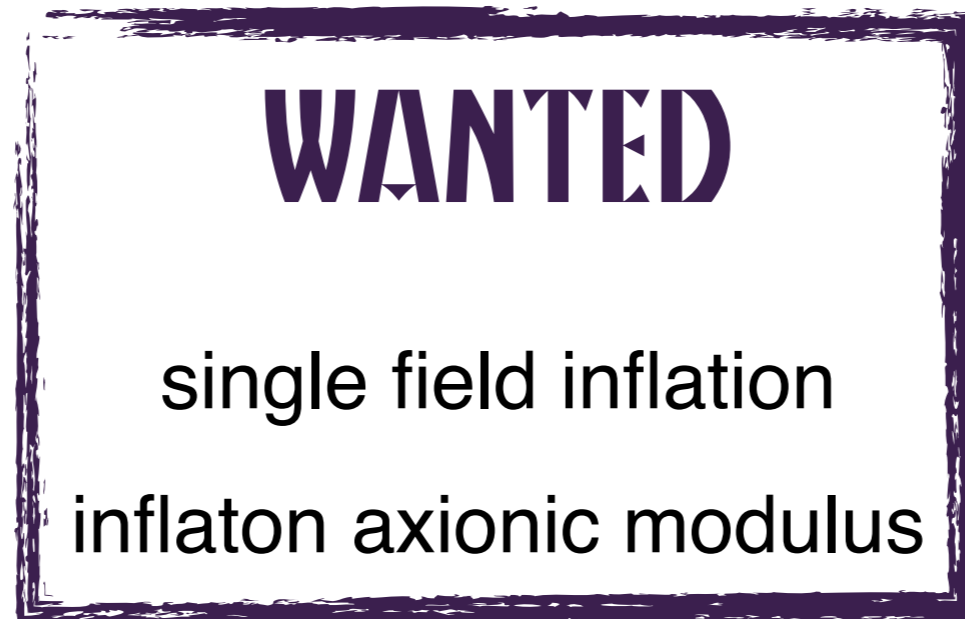
★ **axion monodromy inflation**

potential **polynomial**, generated by fluxes

★ **natural, aligned, .. inflation**

potential **periodic**, generated by e.g. instantons

Moduli stabilisation



moduli stabilisation leaves only one axion light

not easy at all

[Hebecker, Mangat, Rompineve, Witkowski]

[Blumenhagen, DH, Plauschinn]

[Escobar, Landete, Marchesano, Regalado], [Landete, Marchesano, Wieck]

[Palti]

...

How to stabilise moduli for aligned inflation?

Aligned inflation

[Kim, Nilles, Peloso]

- two axions with sub-Planckian axion decay constant
- two axionic linear combinations, one heavy, the other

aligned inflation potential

$$V(\zeta) = \Lambda^4 \left[1 - \cos \left(\frac{\zeta}{f_{eff}} \right) \right]$$

with super-Planckian effective

axion decay constant

So what leads to a *cos*-potential?

The conifold

conifold condition e.g. quintic $|Z| \sim u = 5(\psi - 1) = 0$

vicinity of the conifold $|Z| \ll 1$

[Giddings, Kachru, Polchinski]

backreaction of fluxes leads to a warped CY

$$ds^2 = e^{2A(y)} \eta_{\mu\nu} dx^\mu dx^\nu + e^{-2A(y)} g_{\mu\nu} dy^\mu dy^\nu$$

strongly warped regime not so well controllable

The vicinity of the conifold

★ negligible warping ..

low energy sugra

$$e^{-4A(y)} \sim 1 + \frac{1}{(\mathcal{V}|Z|^2)^{\frac{2}{3}}}$$

$$\mathcal{V}|Z|^2 \gg 1$$

three-cycle volume stays large $V(A) = \mathcal{V}^{\frac{1}{2}} \left| \int_A \Omega_3 \right| = (\mathcal{V}|Z|^2)^{\frac{1}{2}}$

we get this e.g. by requiring

$$M_{string} \gg M_Z$$

The vicinity of the conifold

★ exponential mass hierarchies

single field inflation

origin: log-term in the periods

$$X^1 \sim Z \log Z + \dots$$

example quintic:

consider axio-dilaton + c.s.

$$W = f X^1 + i h S F_1 - i h' S X^0$$

$$D_Z W = 0 \quad D_S W = 0$$

The vicinity of the conifold

★ exponential mass hierarchies

single field inflation

origin: log-term in the periods

$$X^1 \sim Z \log Z + \dots$$

example quintic:

$$Z \sim C e^{-\frac{2\pi h}{f} S}$$

$$M_Z \sim \frac{1}{\mathcal{V}|Z|}$$

$$M_{mod} \sim \frac{1}{\mathcal{V}}$$

What about the volume moduli?

The conic LARGE volume scenario

we only considered the GVW part

Reminder

$$\mathcal{V}|Z|^2 \gg 1$$

$$Z \sim e^{-sth} \quad \Rightarrow \quad \mathcal{V} \gg e^{2sth}$$

LARGE volume scenario natural candidate
to stabilise the volume in the controllable regime
near the conifold

The conic LARGE volume scenario

$$K = -2 \log \left(\tau_b^{\frac{3}{2}} - \tau_s^{\frac{3}{2}} + \frac{\zeta}{2} \text{Re}(S)^{\frac{3}{2}} \right)$$

assume superpotential

$$W = W_0 + A_s Z^N e^{-a_s T_s}$$

standard LVS with

$$A_s \rightarrow A_s Z^N$$

simple example quintic

$$Z \sim \exp \left(-\frac{2\pi h}{f} S \right)$$

$$\mathcal{V} |Z|^2 \sim \exp \left[\frac{a_s}{g_s} \left(\frac{h(N-2)}{f} + (4\zeta)^{\frac{2}{3}} \right) \right]$$

large if $N > 1$, else fluxes can fix it

Shift symmetries of complex structure moduli

Shift symmetry of complex structure moduli

[Garcia-Extebarria, Grimm, Valenzuela]

Kalb-Ramond field and RR fields: shift symmetry from **gauge** symmetry

complex structure moduli shift symmetric at special points

large complex structure

$$K_{c.s.} = -\log f(\Re U_i)$$

imaginary parts of c.s. moduli axionic

$$\Im U_i \rightarrow \Im U_i + a$$

Shift symmetry of complex structure moduli

[Garcia-Extebarria, Grimm, Valenzuela]

Kalb-Ramond field and RR fields: shift symmetry from **gauge** symmetry

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conifold

example \mathbf{IP}_{11226} **[12]**

$$Z = |Z|e^{2\pi i\theta}$$

$$K_{c.s.} = -\log\left(\frac{1}{2\pi}|Z|^2 \log|Z|^2 + C|Z|^2 + A + \Re Y + B \Re Y^2\right)$$

$$\theta \rightarrow \theta + \text{const}$$

$$\Im Y \rightarrow \Im Y + \text{const}$$

in general
higher orders
break

$$\theta \rightarrow \theta + n$$

Realising inflation

Procedure for inflation

example $W_{eff} = f_1 + f_2 U + f_3 S + \Delta W_{ax}$

1. stabilise non-inflatonic moduli

configurations with minimum number of fluxes (simple)

leave inflatonic direction flat

2. turn on a small inflaton-dependent term

e.g. polynomial for axion monodromy inflation

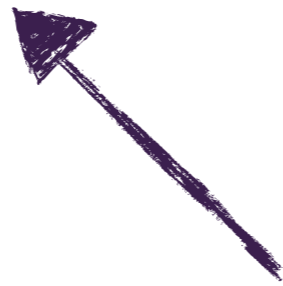
here: **exponential** term from integrating out conic modulus

Aligned inflation example

$$W = f_1 + f_2 U + f_3 S + f_1 (\alpha Z + \beta Z \log Z + \gamma + \dots)$$

Aligned inflation example

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one axionic direction not stabilised

Aligned inflation example

$$W = f_1 + f_2 U + f_3 S + f_1 (\alpha Z + \beta Z \log Z + \gamma + \dots)$$

$$D_S W = 0 \quad D_U W = 0 \quad \text{schematically.. some subtleties}$$

$$\text{integrate out } Z \quad M_Z \sim \frac{1}{\mathcal{V}|Z|}$$

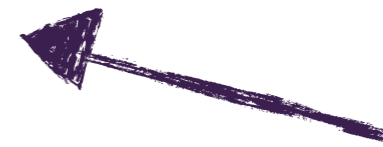
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$$W = f_1 + f_2 U + f_3 S + f_1 (\alpha e^{-(f_4 S + f_5 U)} + \dots)$$



ΔW_{ax}

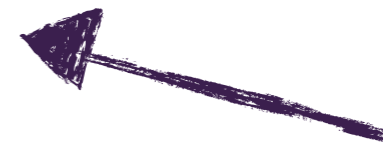
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$$M_{mod} \sim \frac{1}{\mathcal{V}}$$

$$M_{ax} \sim \frac{|Z|}{\mathcal{V}}$$

ΔW_{ax}

exponentially light axionic inflaton

Periods.. example $IP_{11226}[12]$

remember $K_{c.s.} = -\log\left(\frac{1}{2\pi}|Z|^2 \log|Z|^2 + C|Z|^2 + A + \Re Y + B \Re Y^2\right)$

Periods

$$F_0 = 1, \quad F_1 = Z$$

$$F_2 = c_1 + c_2 Y + c_3 Z + c_4 Y^2 + c_5 Y Z + c_6 Z^2 + \dots$$

$$X^0 = \tilde{c}_1 + \tilde{c}_2 Y + \tilde{c}_3 Z + \tilde{c}_4 Y^2 + \tilde{c}_5 Y Z + \tilde{c}_6 Z^2 + \dots$$

$$X^2 = r_1 + r_2 Y + r_3 Y^2 + i_1 Y Z + r_4 Z^2 + \dots$$

$$X^1 = -\frac{1}{2\pi i} Z \log Z + \tilde{r}_1 + \tilde{r}_2 Y + \tilde{i}_1 Z + \dots$$

we want: terms e.g. linear in Z & Y and a Log term

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some non-trivial cancellations can lead to simpler potentials

Periods.. example $IP_{11226}[12]$

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even if we could control a part, we cannot control ...

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Summary

- ★ **region near the conifold**

exponential mass hierarchies with negligible warping

- ★ **potential for aligned inflation**

generated purely by the form of the periods

- ★ **exact form of the periods crucial**

Aligned inflation

more details

- potential
- mass hierarchies
- weak gravity conjecture
- challenges

Florian Wolf's talk

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THANK YOU!