

What is Dark Matter? and Classification of Effective operators

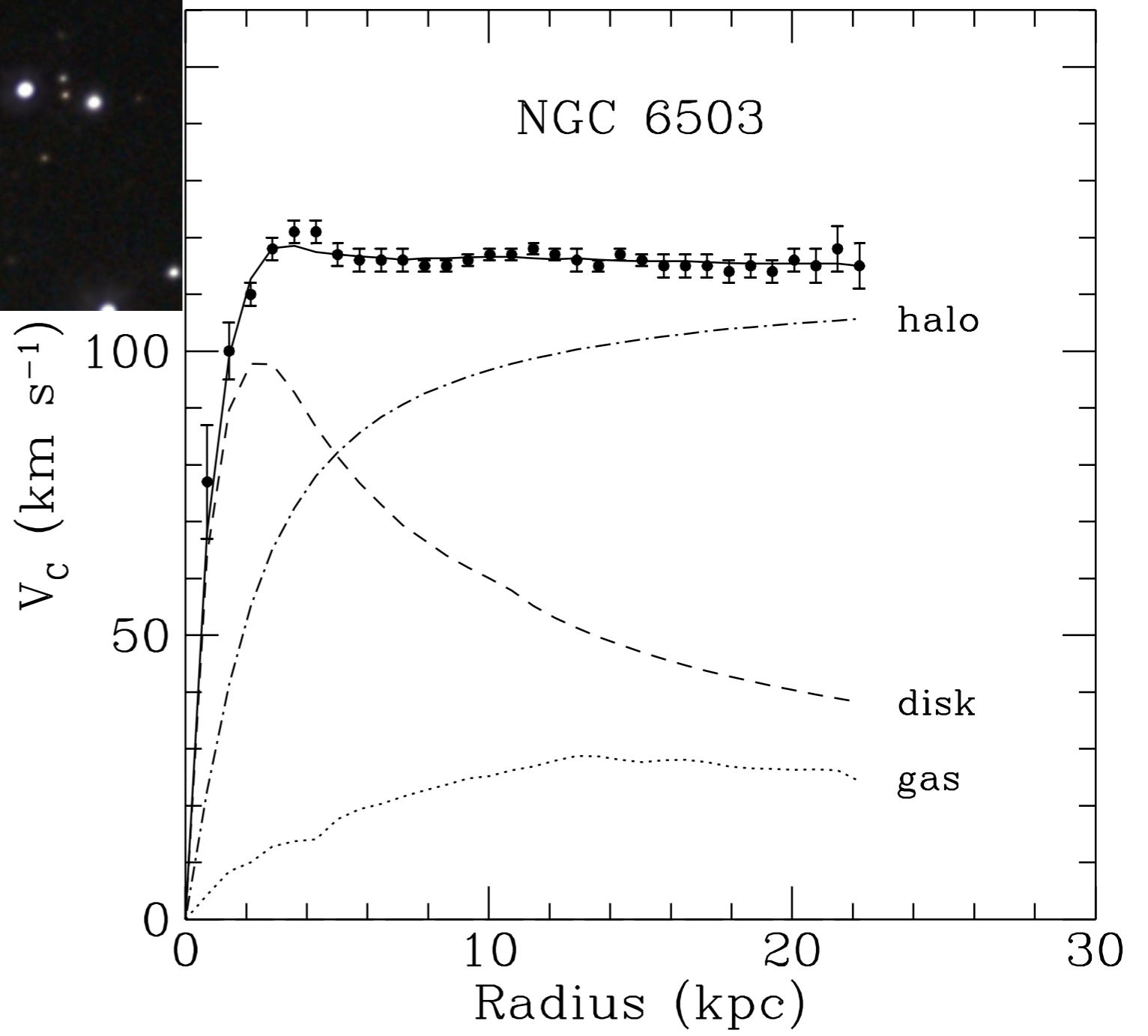
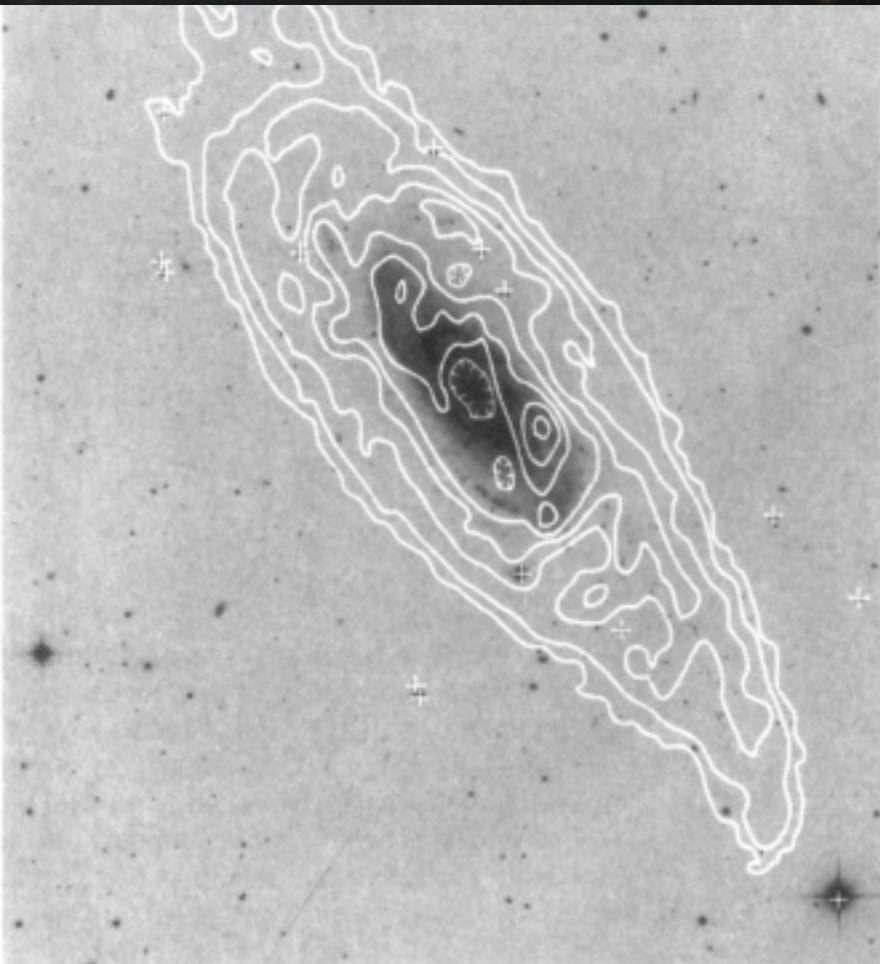
Hitoshi Murayama (Berkeley, Kavli IPMU, DESY)
Arnold Sommerfeld Seminar, July 25, 2019



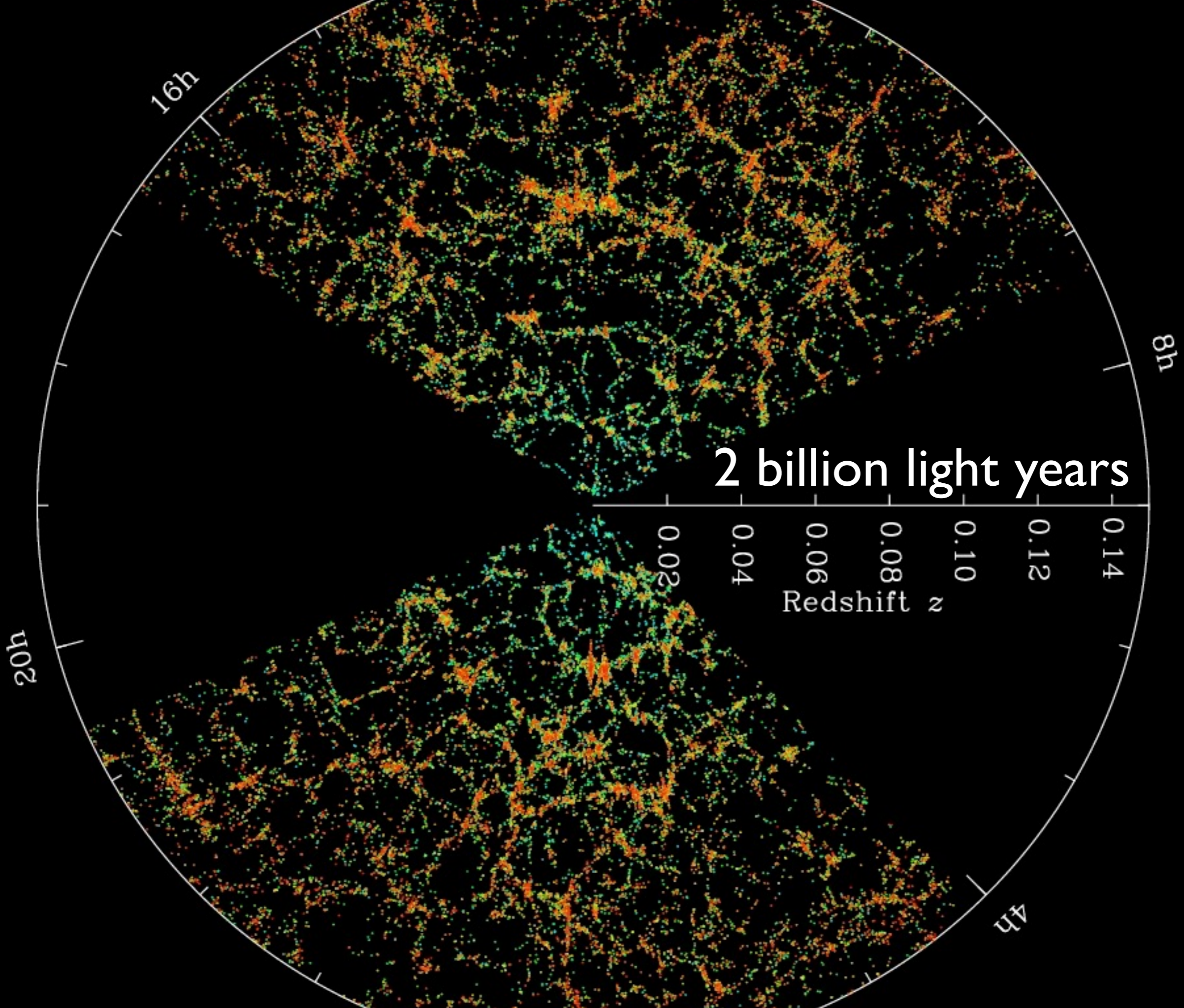
What is Dark Matter?

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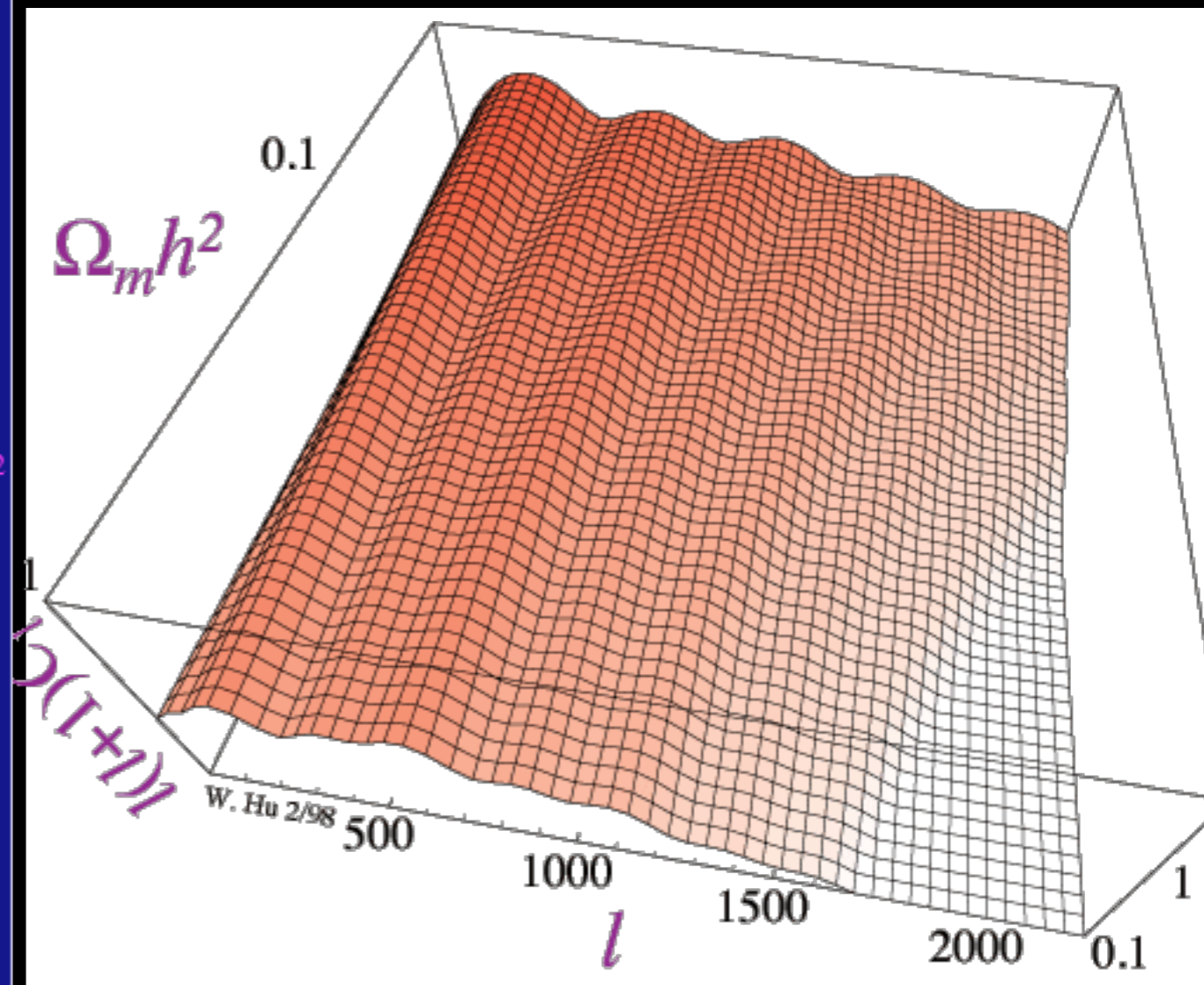
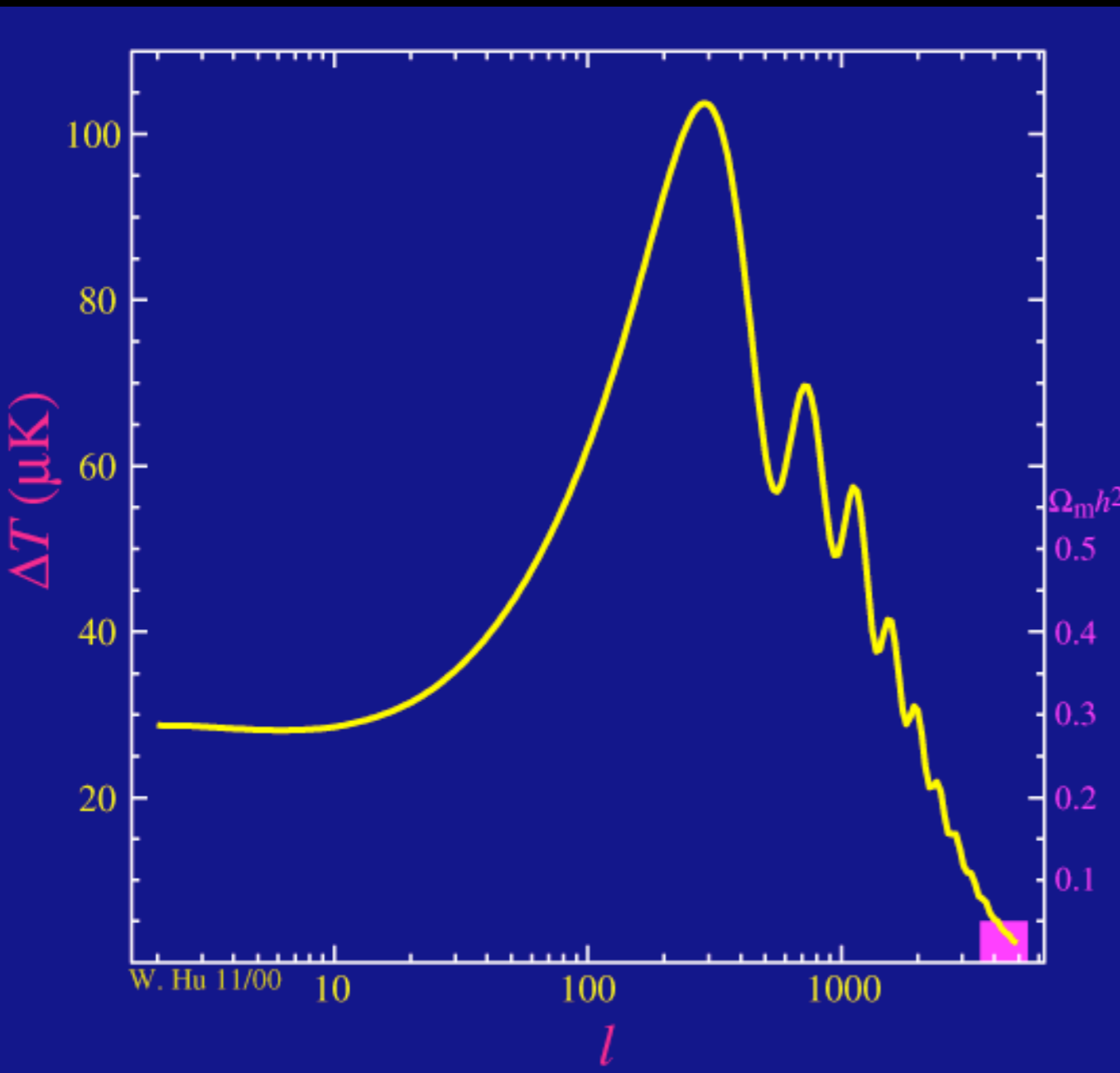




2 billion light years

Redshift z

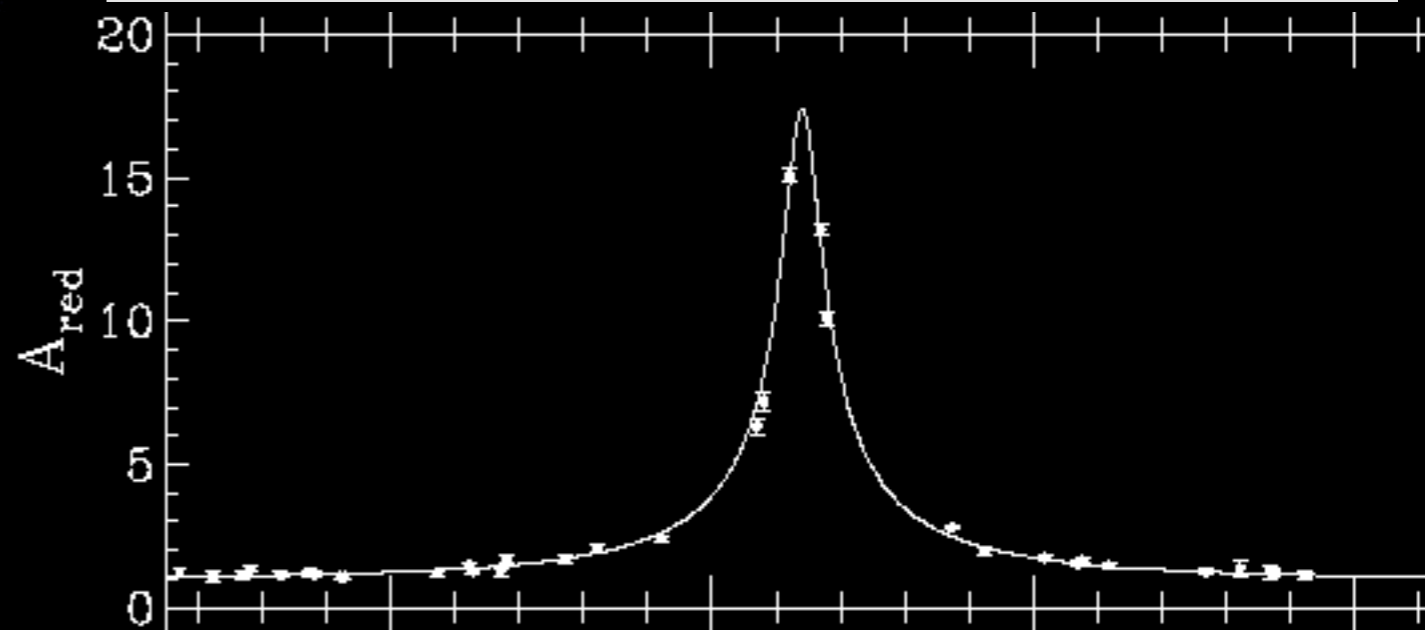
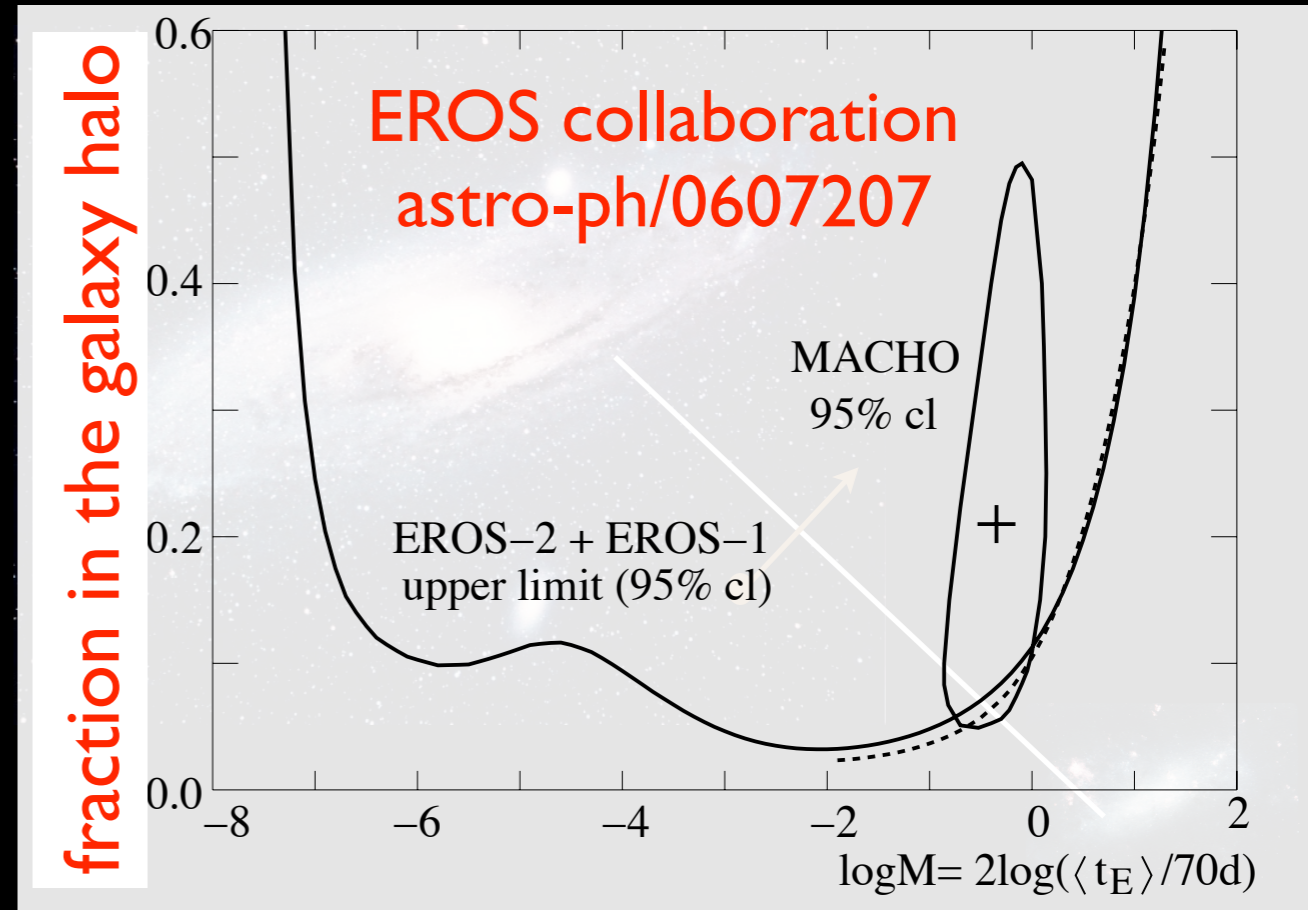
matter



Ω_m changes the overall heights of the peaks

Dim Stars? Black Holes?

Search for **MACHOs**
(Massive Compact Halo Objects)

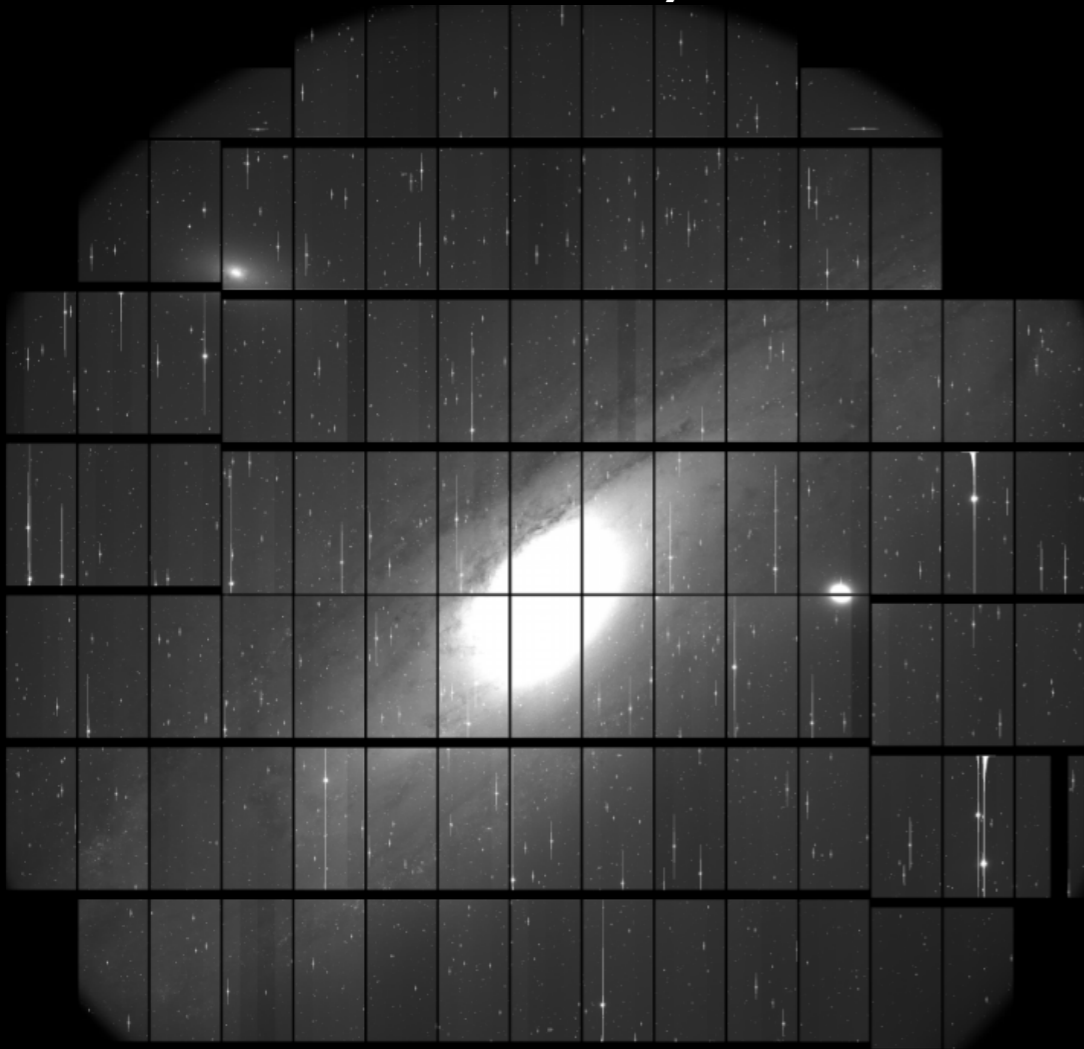


Not enough of them!

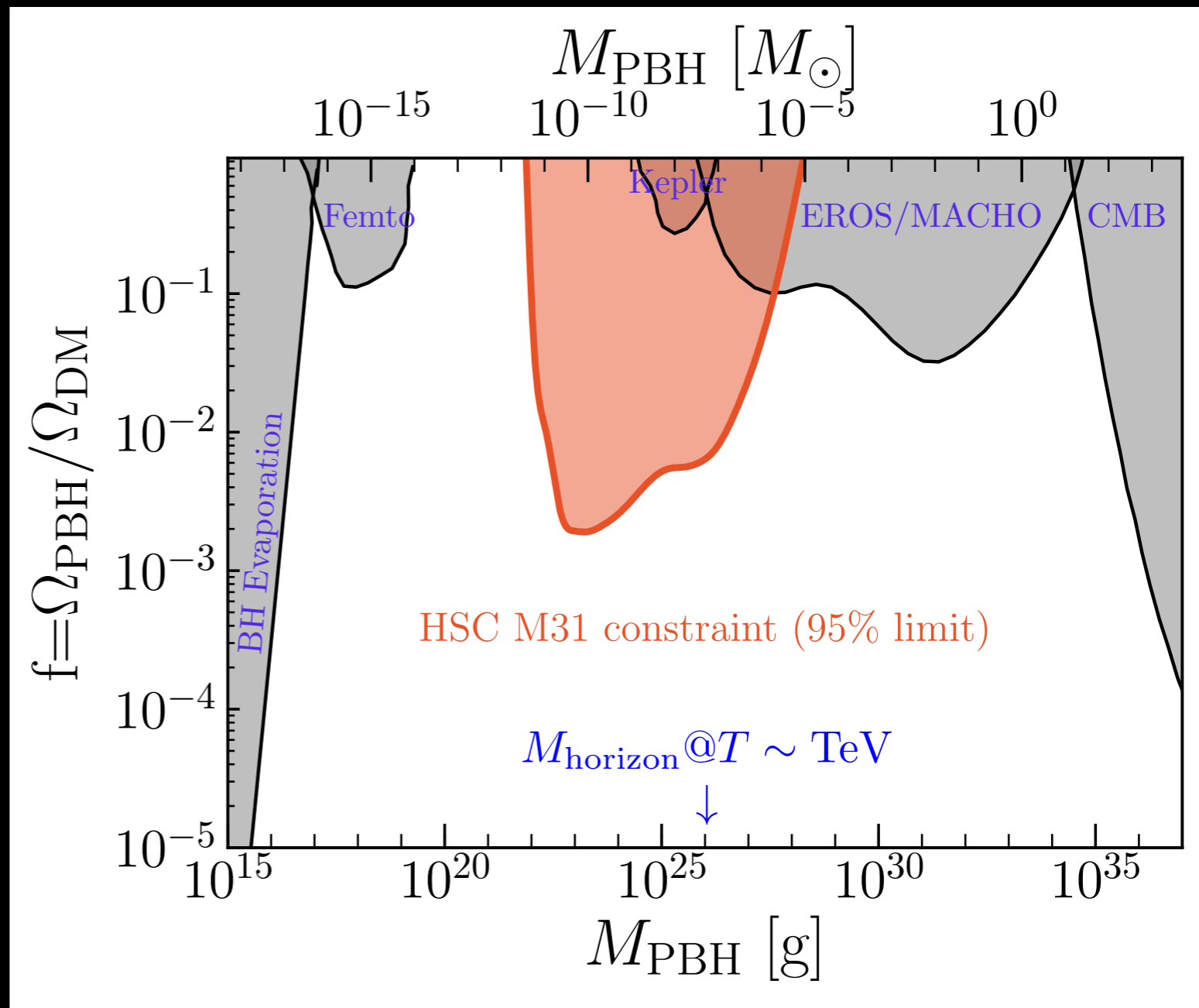
Best limit on Black Hole dark matter



observe Andromeda for one night
read out CCDs every 2 min



No detection \Rightarrow more stringent
upper bound, than 2yr Kepler data
(Griest et al.)





Mass Limits

“Uncertainty Principle”

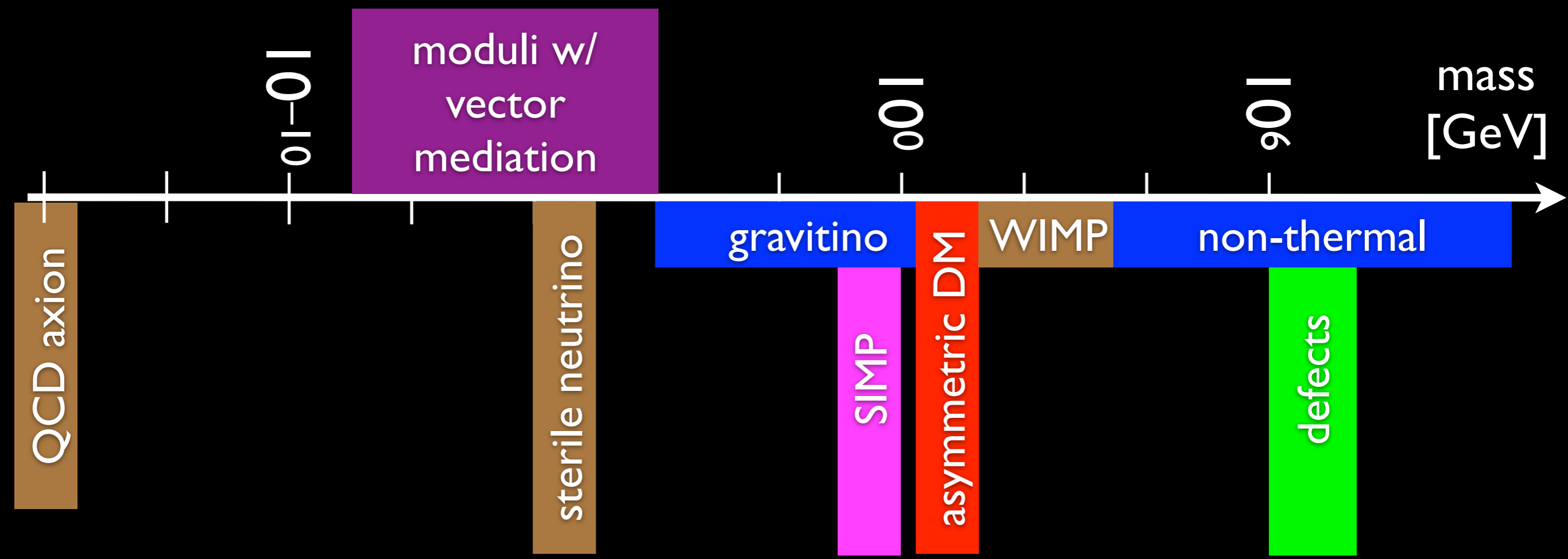
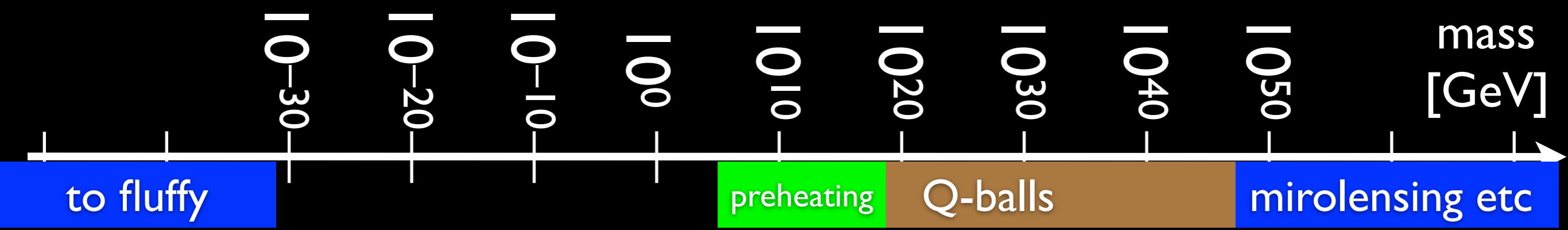
- Clumps to form structure

- imagine $V = G_N \frac{Mm}{r}$

- “Bohr radius”: $r_B = \frac{\hbar^2}{G_N M m^2}$

- too small $m \Rightarrow$ won't “fit” in a galaxy!

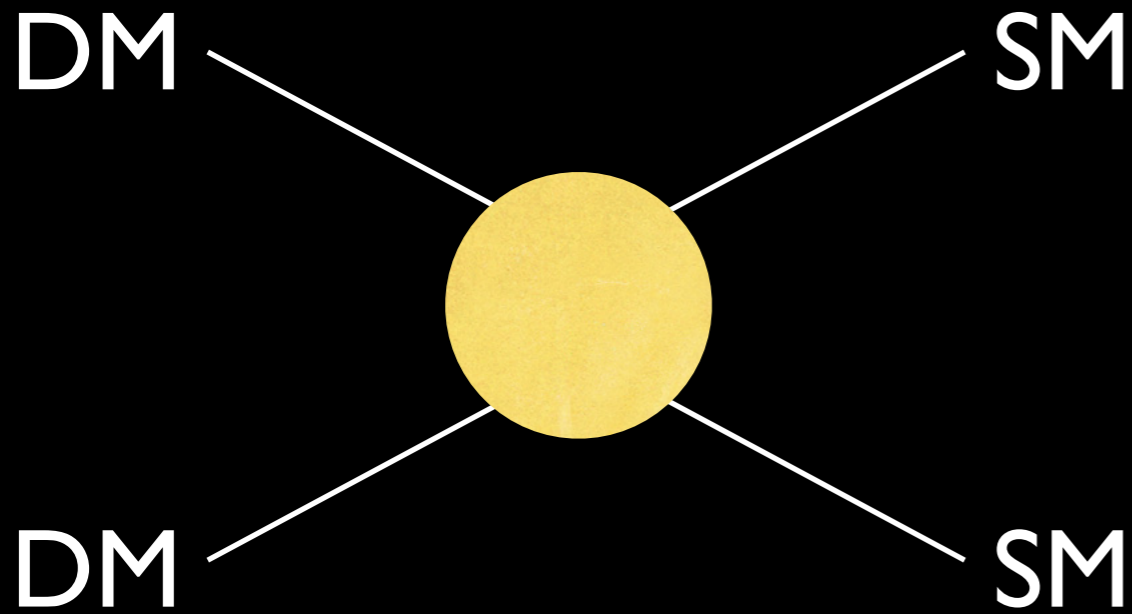
- $m > 10^{-22}$ eV “uncertainty principle” bound
(modified from Hu, Barkana, Gruzinov, astro-ph/0003365)



SIMP: dark hadrons
 $m \sim 0.3 \text{ GeV}$, $\sigma \sim 10^{-24} \text{ cm}^2$



WIMP Miracle



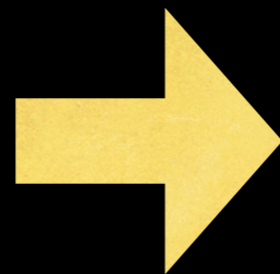
$$\langle \sigma_{2 \rightarrow 2\nu} \rangle \approx \frac{\alpha^2}{m^2}$$

$$\alpha \approx 10^{-2}$$

$$m \approx 300 \text{ GeV}$$

$$\frac{n_{\text{DM}}}{s} = 4.4 \times 10^{-10} \frac{\text{GeV}}{m_{\text{DM}}}$$

“weak” coupling
“weak” mass scale

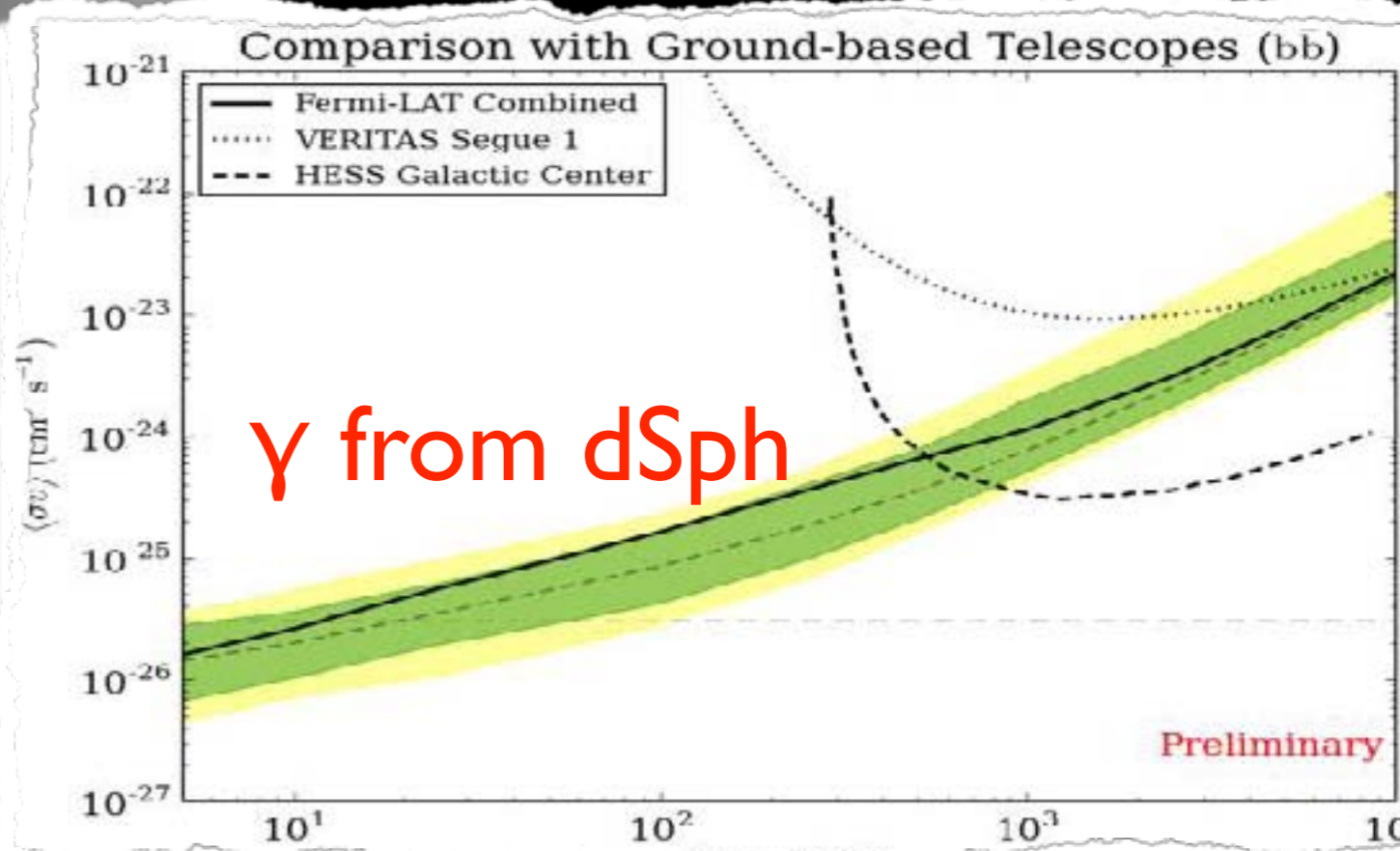
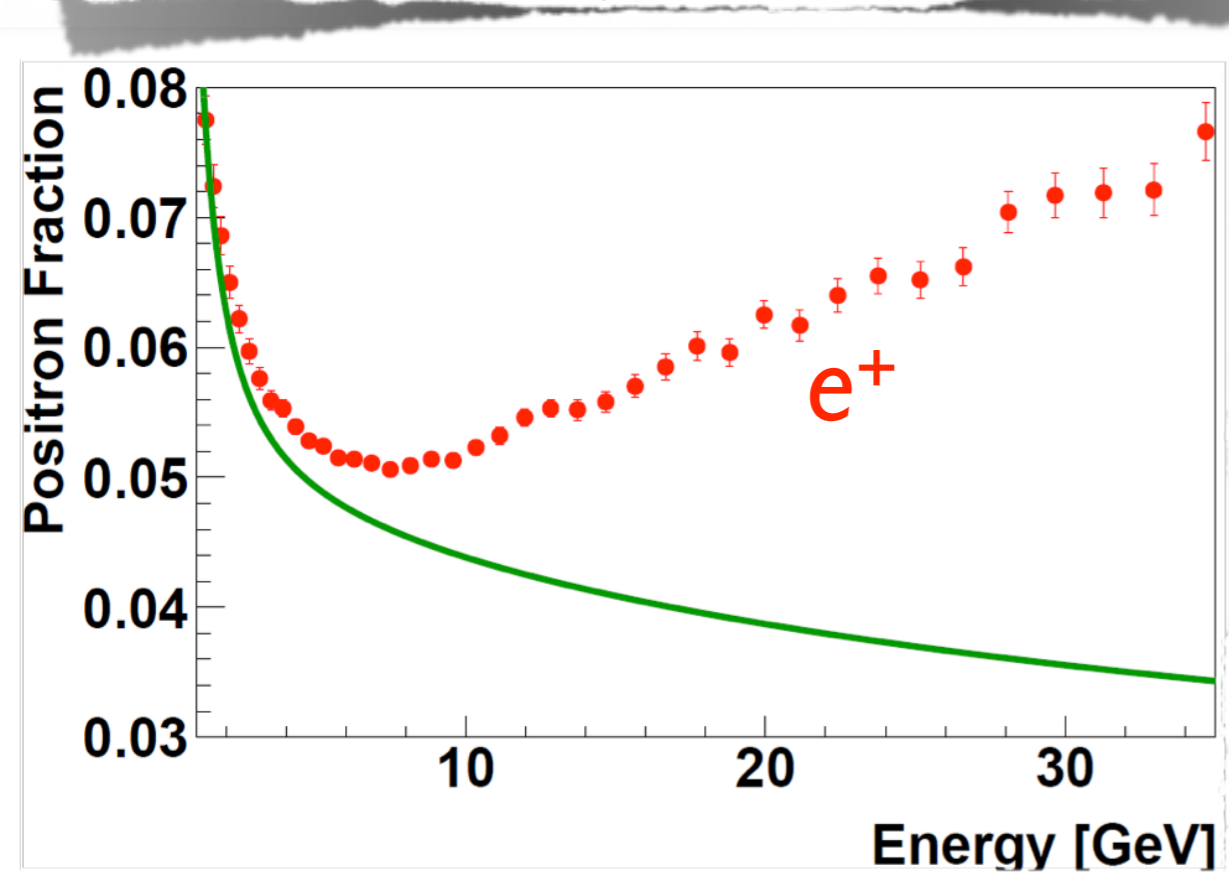
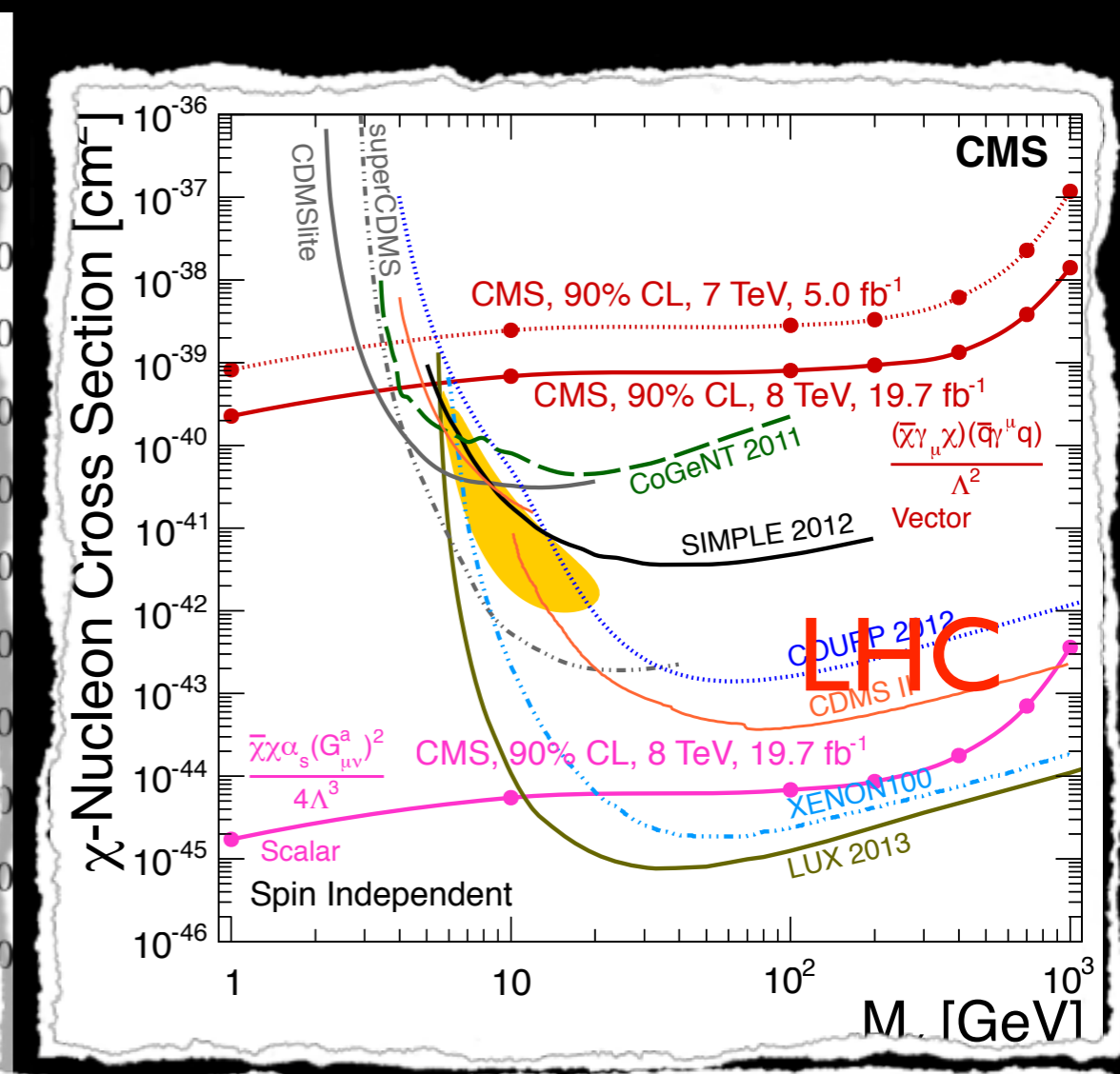
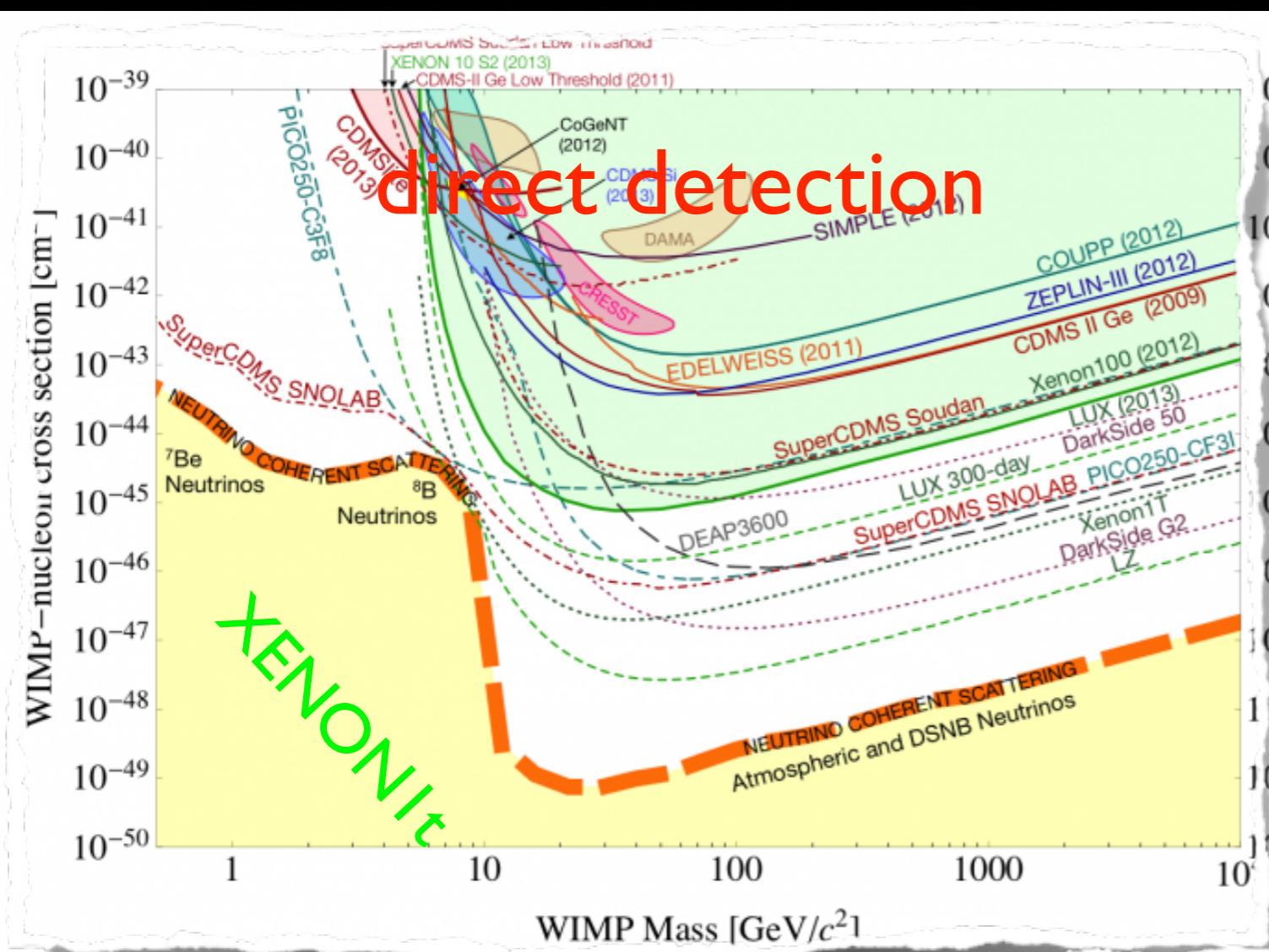


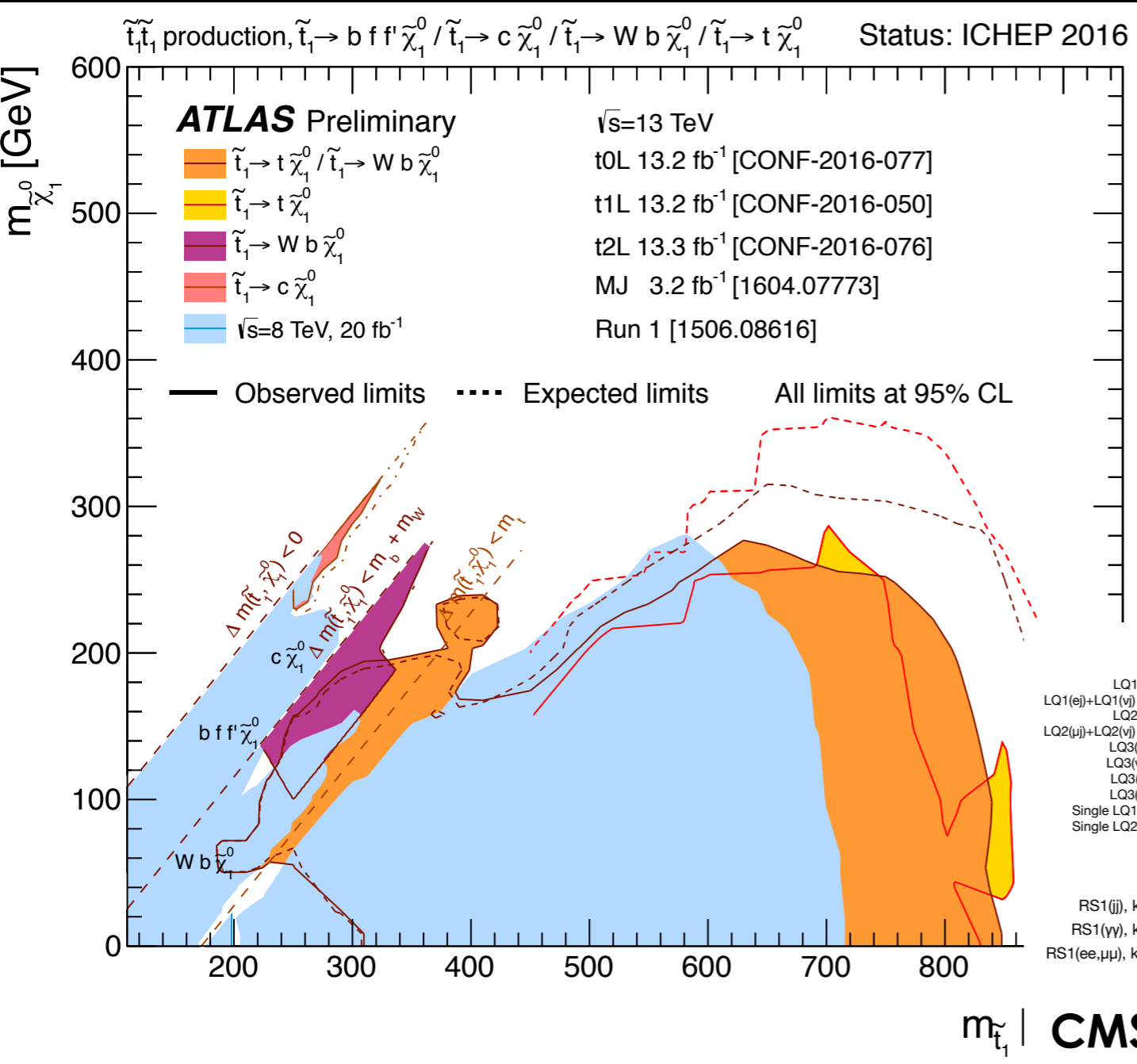
correct abundance

Miracle²

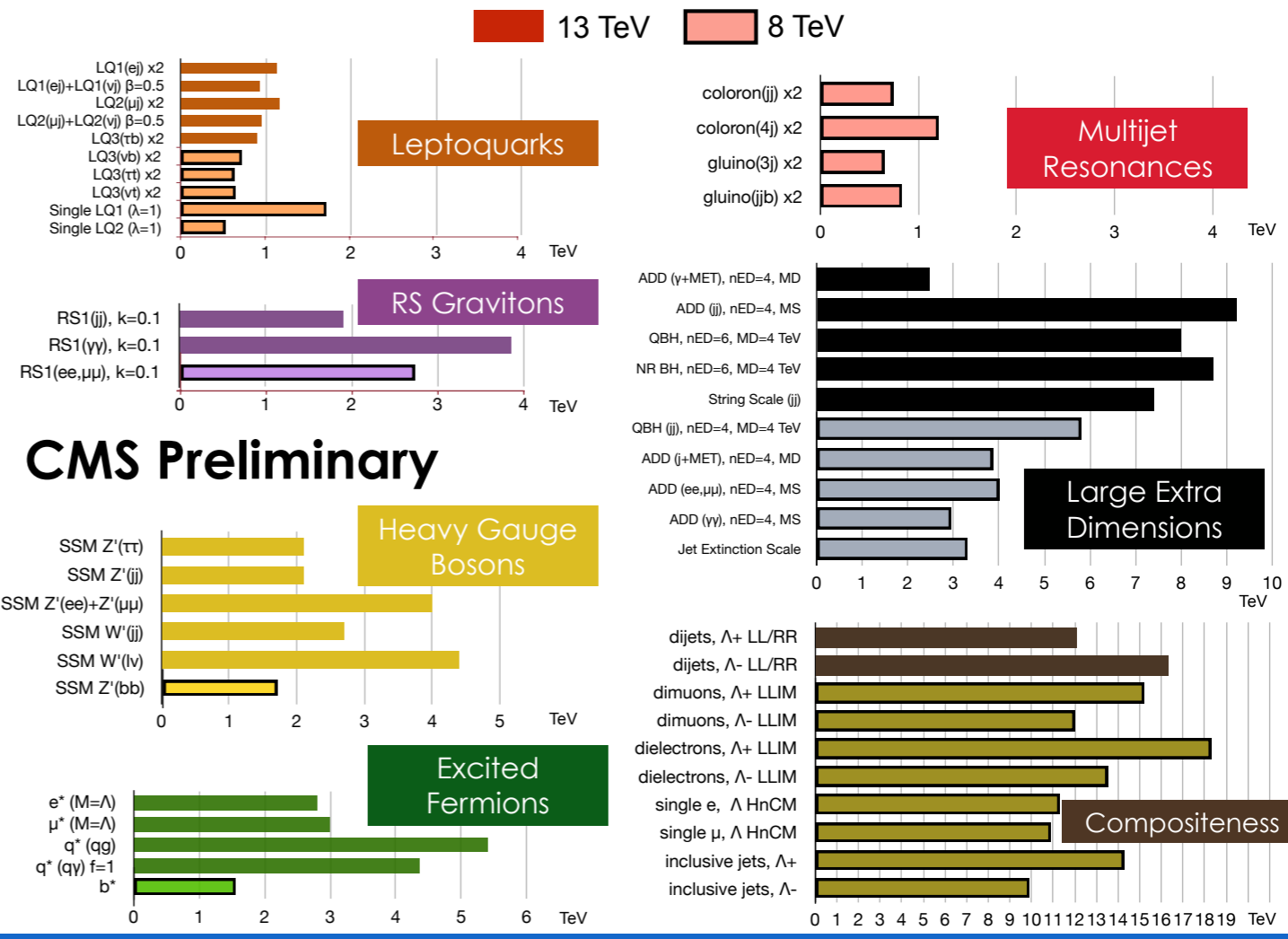
sociology

- We used to think
 - need to solve problems with the SM
 - hierarchy problem, strong CP, etc
 - it is great if a solution also gives dark matter candidate as an *option*
 - big ideas: supersymmetry, extra dim
 - probably because dark matter problem was not so established in 80's





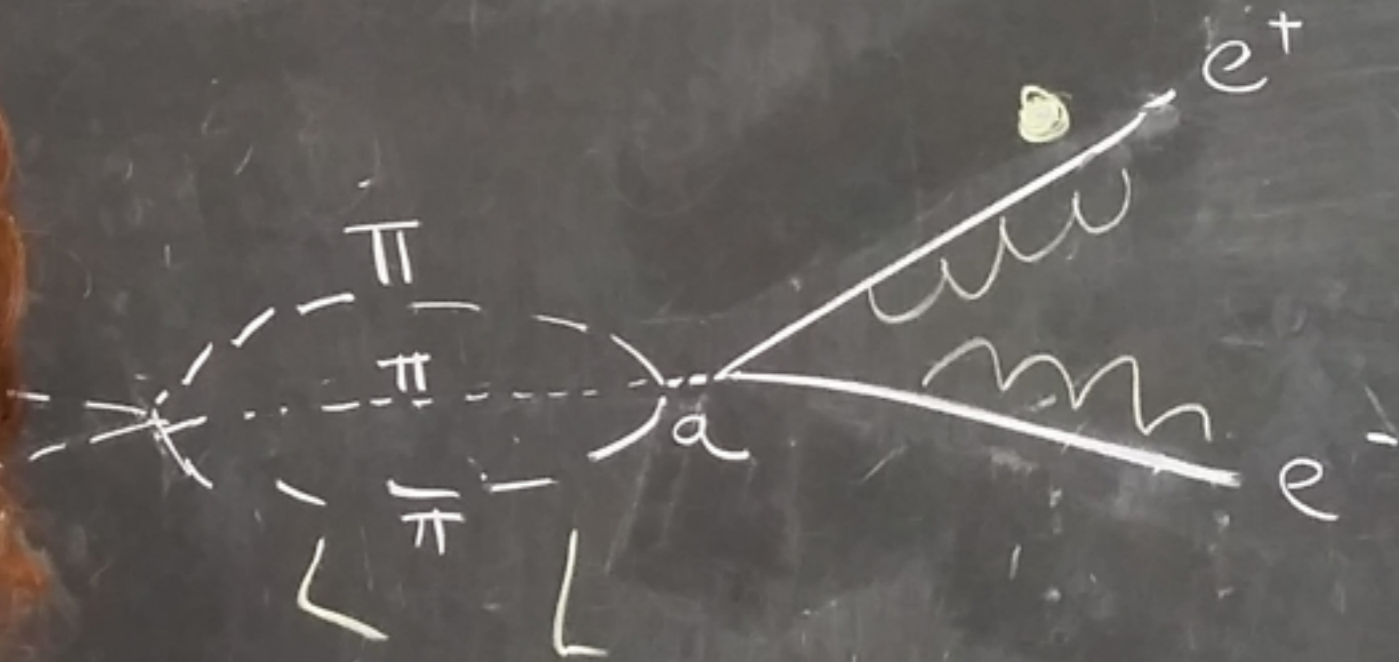
no sign of new physics that explains naturalness!



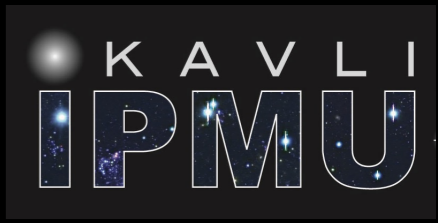
recent thinking

- dark matter definitely exists
 - naturalness problem may be optional?
- need to explain dark matter on its own
- perhaps we should decouple these two
- do we really need big ideas like SUSY?
- perhaps we can solve it with ideas more familiar to us?

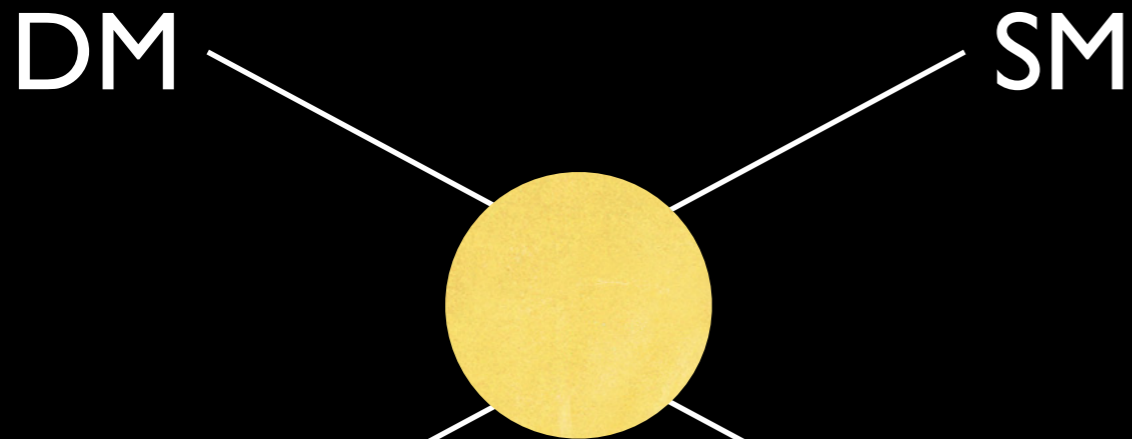
Seminar in Berkeley
Strongly Interacting Massive Particle
(SIMP)



Yonit Hochberg



Miracles



$$\langle \sigma_{2 \rightarrow 2\nu} \rangle \approx \frac{\alpha^2}{m^2}$$

$$\alpha \approx 10^{-2}$$

$$m \approx 300 \text{ GeV}$$

WIMP miracle!

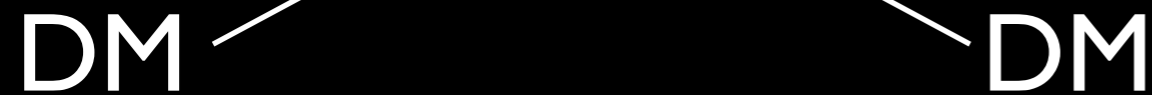
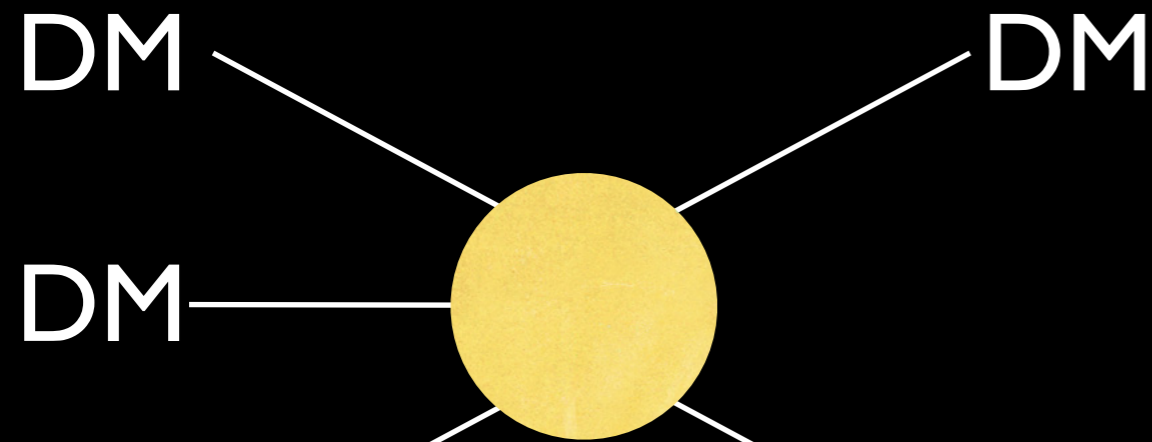
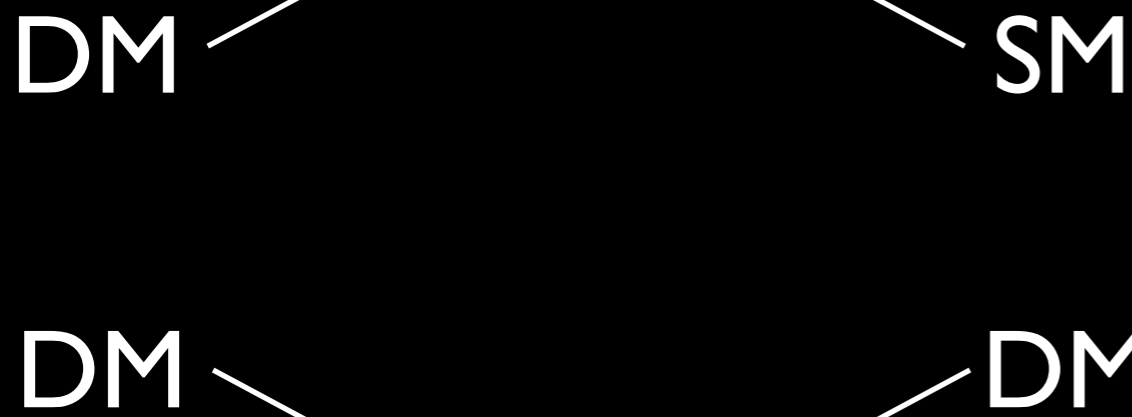
$$\frac{n_{\text{DM}}}{s} = 4.4 \times 10^{-10} \frac{\text{GeV}}{m_{\text{DM}}}$$

$$\langle \sigma_{3 \rightarrow 2\nu^2} \rangle \approx \frac{\alpha^3}{m^5}$$

$$\alpha \approx 4\pi$$

Hochberg, Kuflik,
Volansky, Wacker

$$m \approx 300 \text{ MeV} \text{ arXiv:1402.5143}$$



SIMP miracle!

~~LEE WEINBERG~~ FREEZE-OUT

3 → 2
Back of the envelope calculation

$$\Gamma_{\text{ann}} \simeq H|_{\text{freezeout}}$$

$$\Gamma_{\text{ann}} \simeq n_{\text{dm}}^2 \langle \sigma v^2 \rangle_{3 \rightarrow 2}$$

$$H \simeq \frac{T^2}{M_{\text{pl}}}$$

$$m_{\text{dm}} n_{\text{dm}} \sim m_p n_b$$

$$\langle \sigma v^2 \rangle_{3 \rightarrow 2} \simeq \frac{\alpha^3}{m_{\text{dm}}^2}$$

$$n_b \sim \eta_b s$$

Eric Kuflik

$$\eta_b \simeq T_{\text{eq}}/m_p$$

$$s \simeq T^3$$

$$\Gamma_{\text{ann}} \simeq \frac{T_{\text{eq}}^2 \alpha^3}{x_F^3 \times m_{\text{dm}}} =$$

$$H \simeq \frac{m_{\text{dm}}^2}{M_{\text{pl}} x_F^2}$$



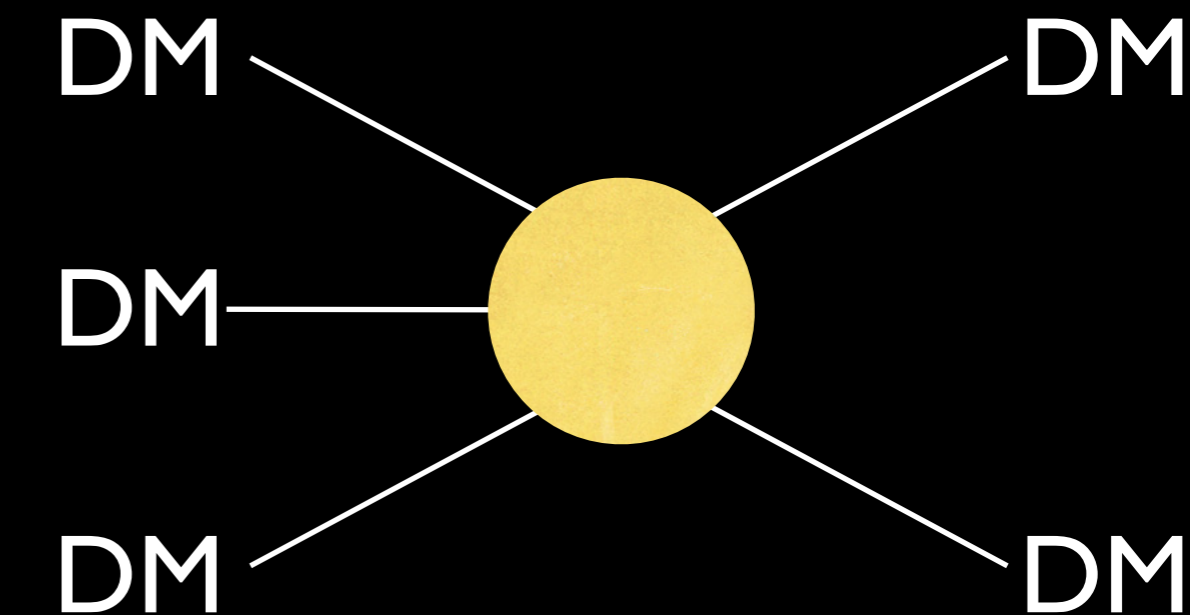
SIMPLest Miracle

- Not only the mass scale is similar to QCD
- dynamics itself can be QCD! Miracle³
- DM = pions
- e.g. $SU(4)/Sp(4) = S^5$

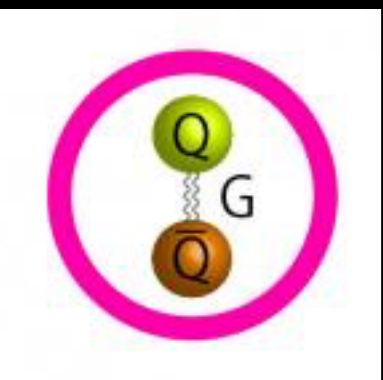
$$\mathcal{L}_{\text{chiral}} = \frac{1}{16f_\pi^2} \text{Tr} \partial^\mu U^\dagger \partial_\mu U$$

$$\mathcal{L}_{\text{WZW}} = \frac{8N_c}{15\pi^2 f_\pi^5} \epsilon_{abcde} \epsilon^{\mu\nu\rho\sigma} \pi^a \partial_\mu \pi^b \partial_\nu \pi^c \partial_\rho \pi^d \partial_\sigma \pi^e + O(\pi^7)$$

$$\pi_5(G/H) \neq 0$$



Phys.Rev.Lett. 115 (2015) 021301

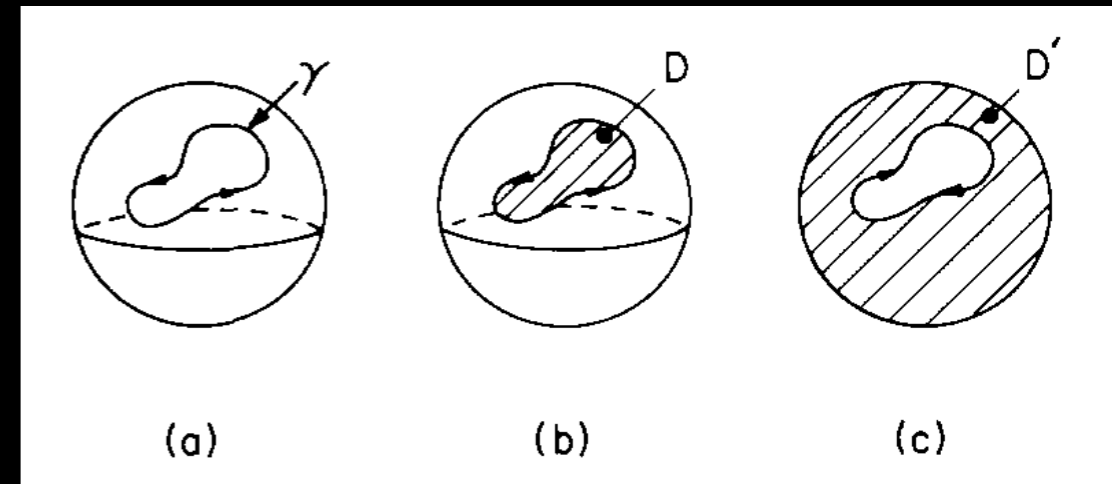


SIMPlEst Miracle

- $SU(2)$ gauge theory with four doublets
- $SU(4)=SO(6)$ flavor symmetry
- $\langle q^i q^j \rangle \neq 0$ breaks it to $Sp(2)=SO(5)$
- coset space $SO(6)/SO(5)=S^5$
- $\pi_5(S^5)=\mathbb{Z} \Rightarrow$ Wess-Zumino term
- $\mathcal{L}_{WZ} = \epsilon_{abcde} \epsilon^{\mu\nu\rho\sigma} \pi^a \partial_\mu \pi^b \partial_\nu \pi^c \partial_\rho \pi^d \partial_\sigma \pi^e$

Wess-Zumino term

- $SU(N_c)$ gauge theory
 - $\pi_5(SU(N_f)) = \mathbb{Z}$ ($N_f \geq 3$)
- $Sp(N_c)$ gauge theory
 - $\pi_5(SU(2N_f)/Sp(N_f)) = \mathbb{Z}$ ($N_f \geq 2$)
- $SO(N_c)$ gauge theory
 - $\pi_5(SU(N_f)/SO(N_f)) = \mathbb{Z}$ ($N_f \geq 3$)



Witten

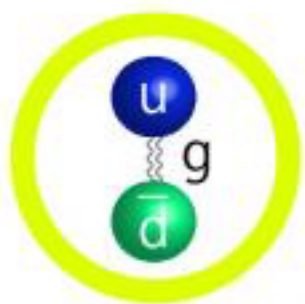
LAGRANGIANS

Quark theory

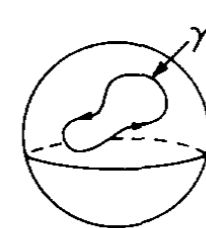
$$\mathcal{L}_{\text{quark}} = -\frac{1}{4} F_{\mu\nu}^a F^{\mu\nu a} + \bar{q}_i i \not{D} q_i - \frac{1}{2} m_Q J^{ij} q_i q_j + h.c.$$

Sigma theory

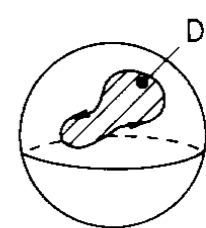
$$\mathcal{L}_{\text{Sigma}} = \frac{f_\pi^2}{16} \text{Tr} \partial_\mu \Sigma \partial^\mu \Sigma^\dagger - \frac{1}{2} m_Q \mu^3 \text{Tr} J \Sigma + h.c. - \frac{i N_c}{240 \pi^2} \int \text{Tr} (\Sigma^\dagger d\Sigma)^5$$



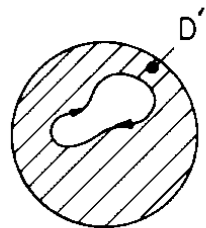
Pion theory



(a)



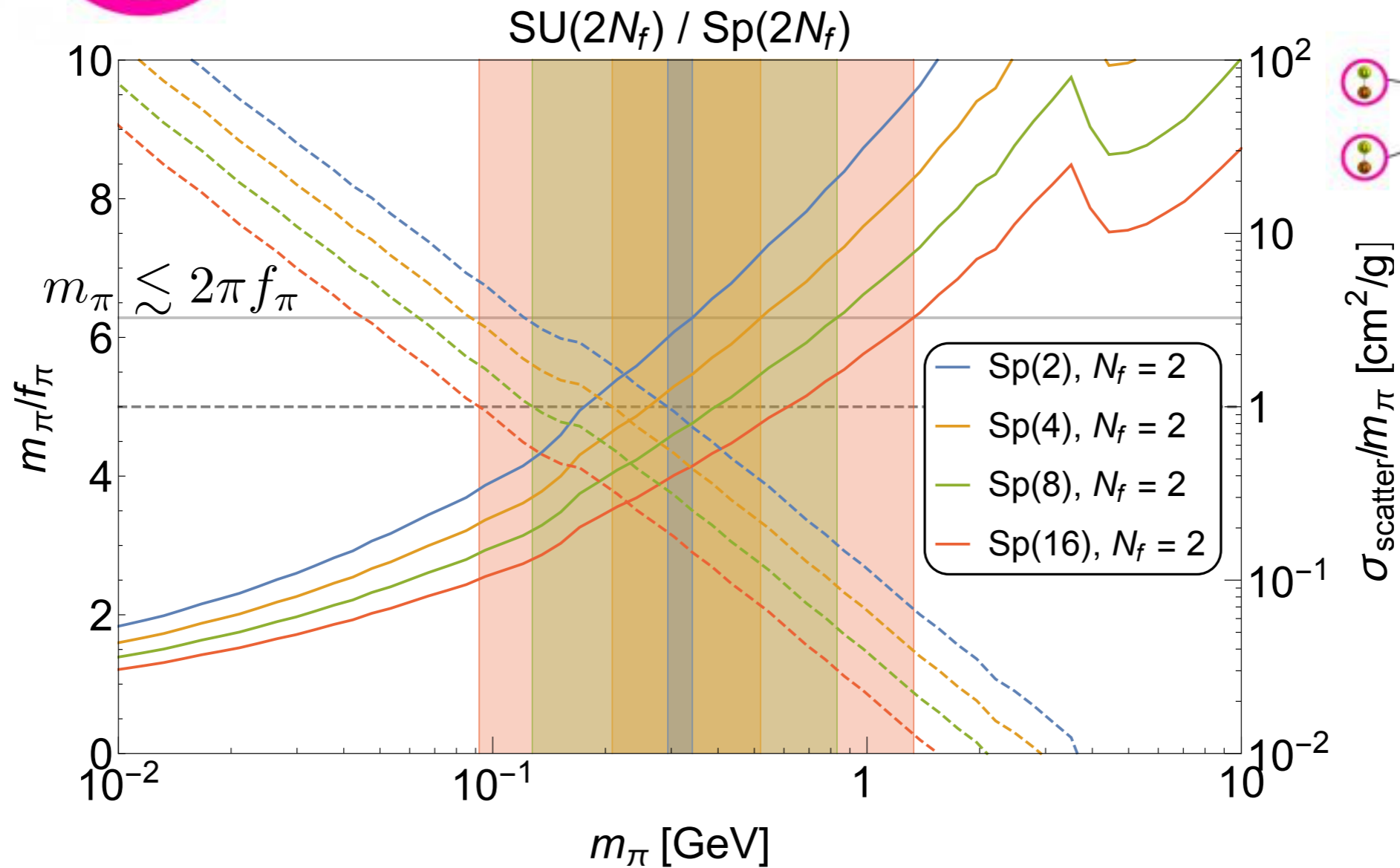
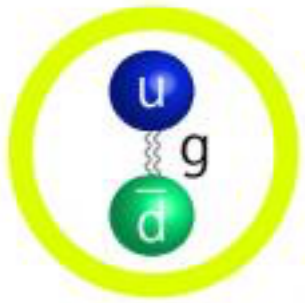
(b)



(c)

$$\mathcal{L}_{\text{pion}} = \frac{1}{4} \text{Tr} \partial_\mu \pi \partial^\mu \pi - \frac{m_\pi^2}{4} \text{Tr} \pi^2 + \frac{m_\pi^2}{12 f_\pi^2} \text{Tr} \pi^4 - \frac{1}{6 f_\pi^2} \text{Tr} (\pi^2 \partial^\mu \pi \partial_\mu \pi - \pi \partial^\mu \pi \pi \partial_\mu \pi) + \frac{2 N_c}{15 \pi^2 f_\pi^5} \epsilon^{\mu\nu\rho\sigma} \text{Tr} [\pi \partial_\mu \pi \partial_\nu \pi \partial_\rho \pi \partial_\sigma \pi] + \mathcal{O}(\pi^6)$$

The Results



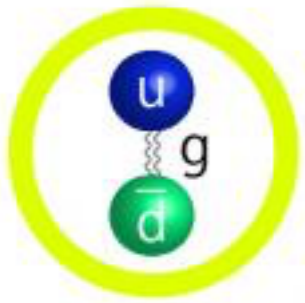
Solid curves: solution to Boltzmann eq.

Dashed curves: along that solution

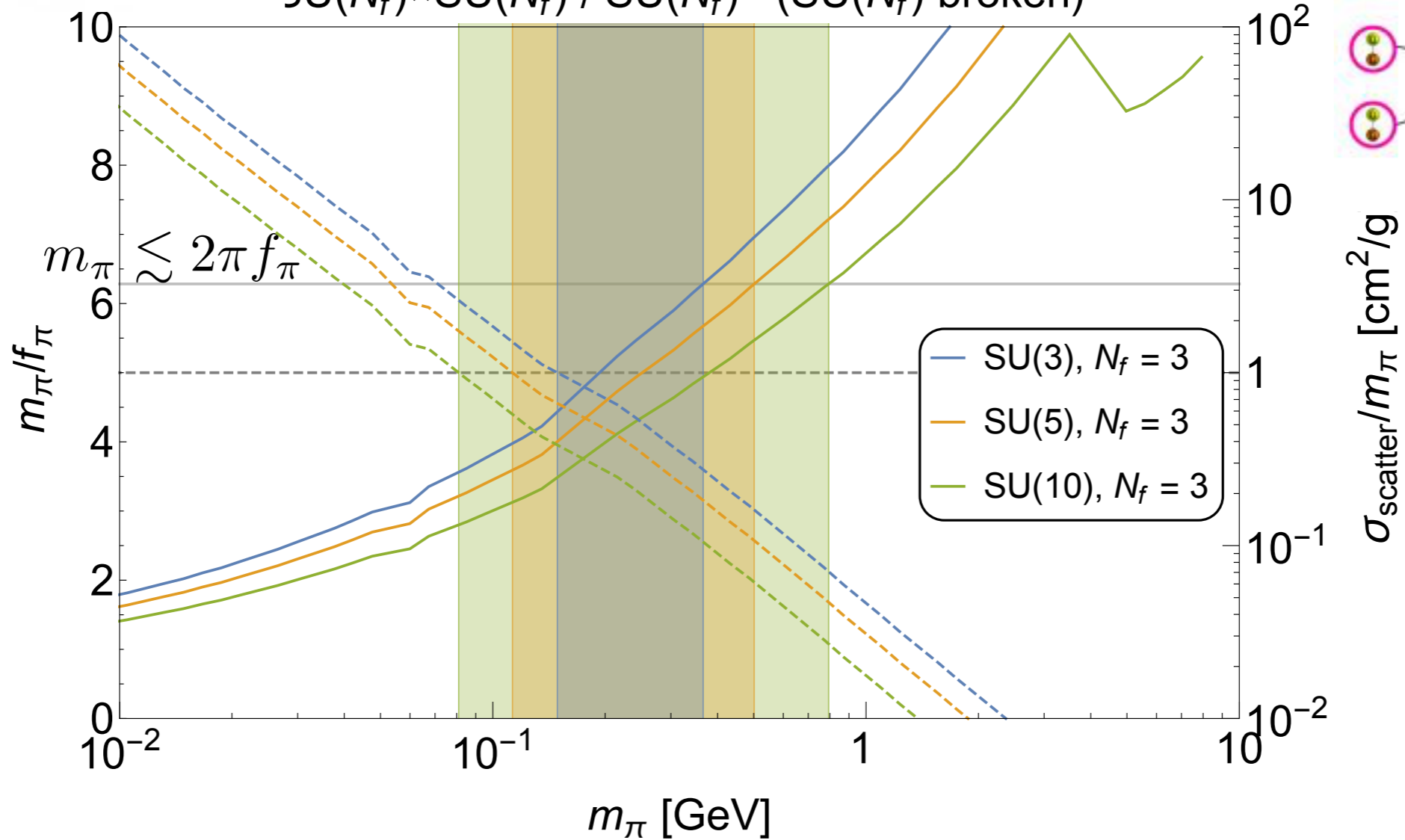
$$\frac{m_\pi}{f_\pi} \propto m_\pi^{3/10}$$

$$\frac{\sigma_{\text{scatter}}}{m_\pi} \propto m_\pi^{-9/5}$$

The Results



$3U(N_f) \times SU(N_f) / SU(N_f)$ ($SU(N_f)$ broken)



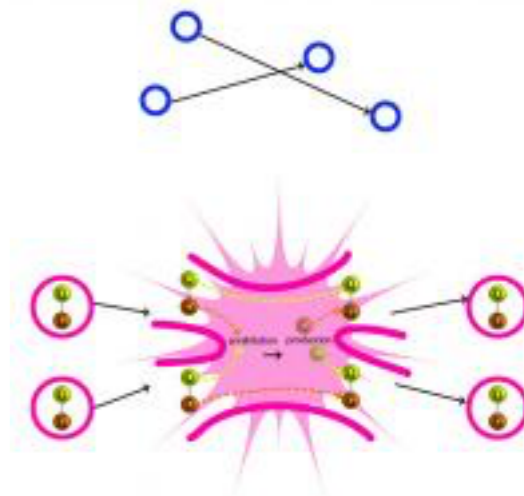
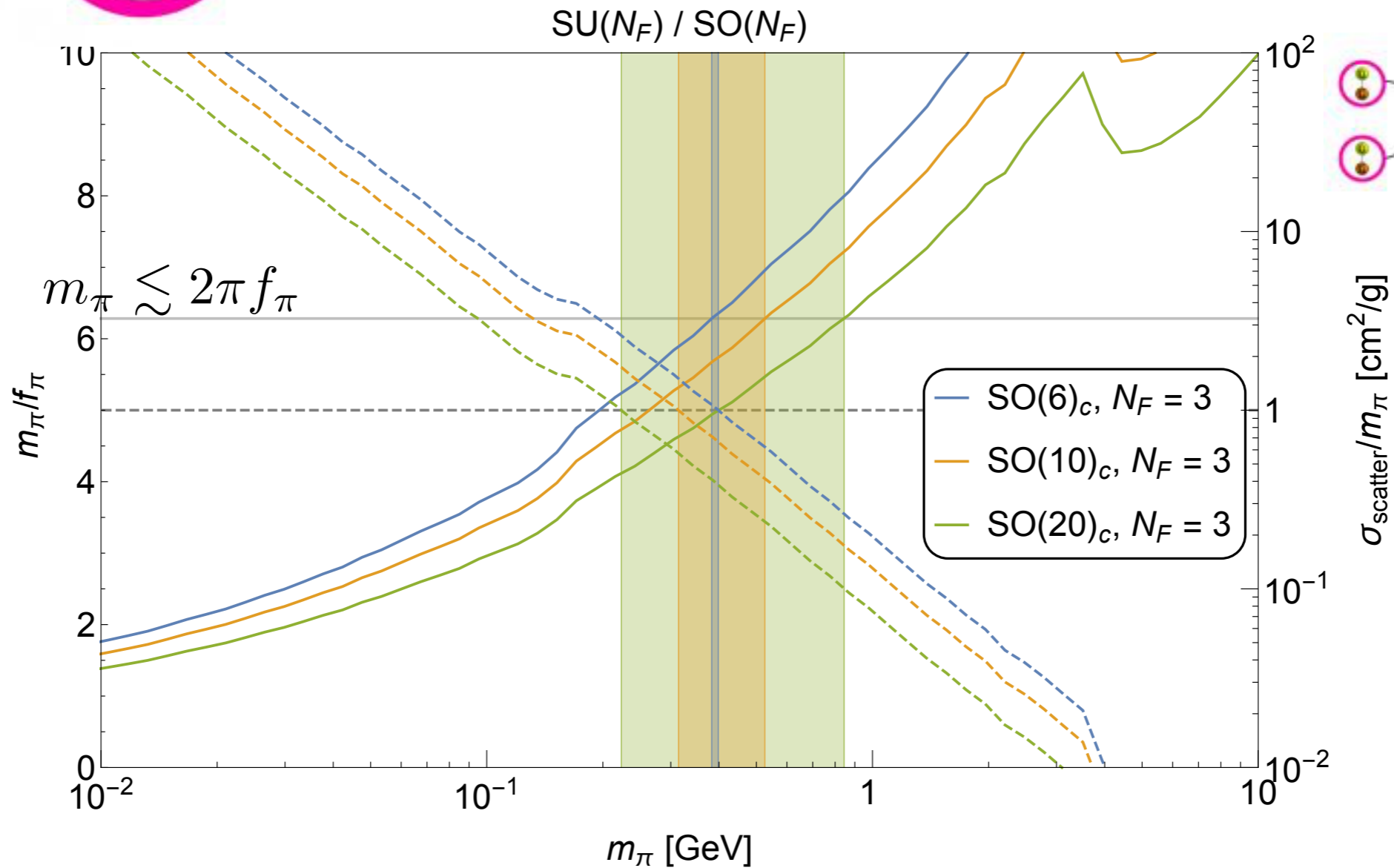
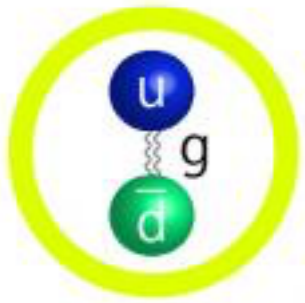
Solid curves: solution to Boltzmann eq.

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$$\frac{m_\pi}{f_\pi} \propto m_\pi^{3/10}$$

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The Results



Solid curves: solution to Boltzmann eq.

Dashed curves: along that solution

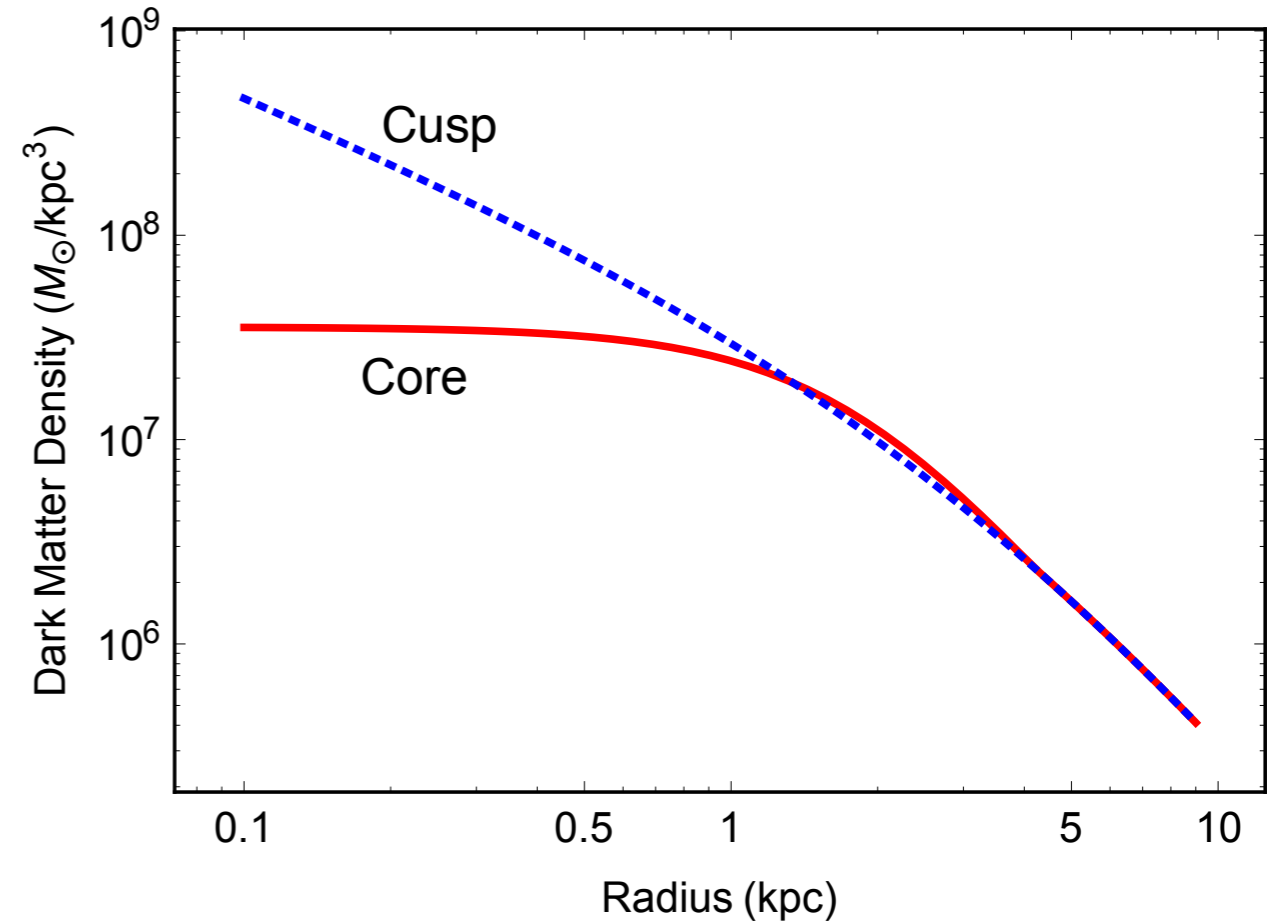
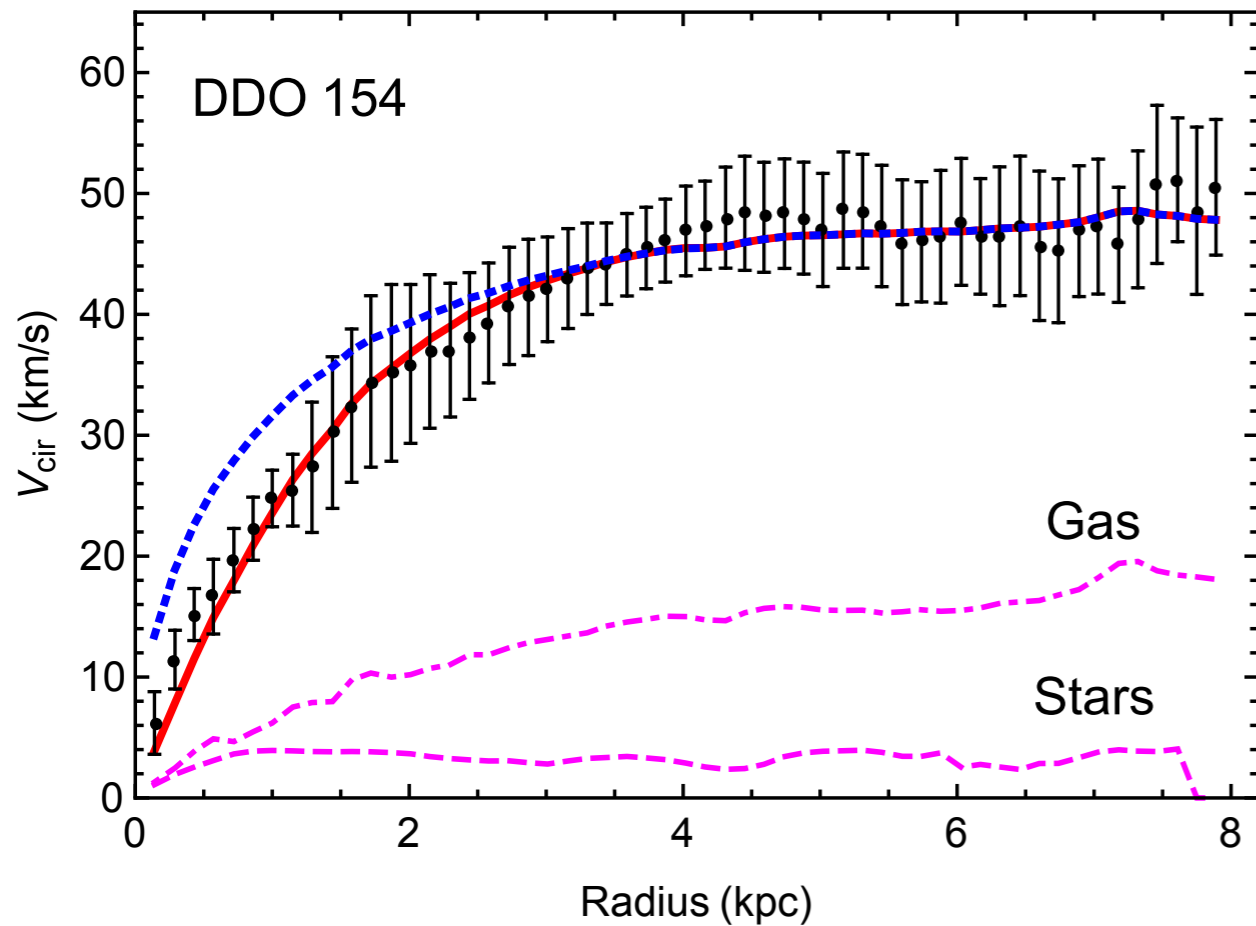
$$\frac{m_\pi}{f_\pi} \propto m_\pi^{3/10}$$

$$\frac{\sigma_{\text{scatter}}}{m_\pi} \propto m_\pi^{-9/5}$$

DDO 154 dwarf galaxy



DDO 154 dwarf galaxy



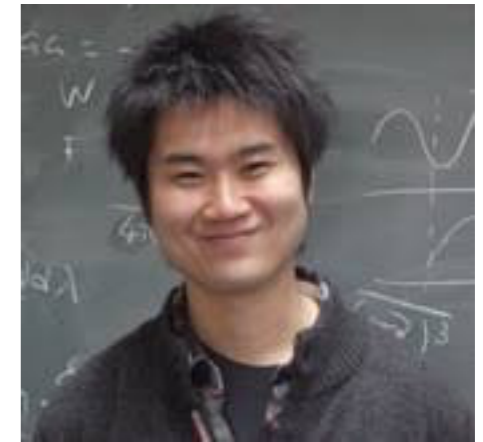
can be explained if dark matter scatters against itself
Need $\sigma/m \sim 1 \text{ b} / \text{GeV}$

only astrophysical information beyond gravity

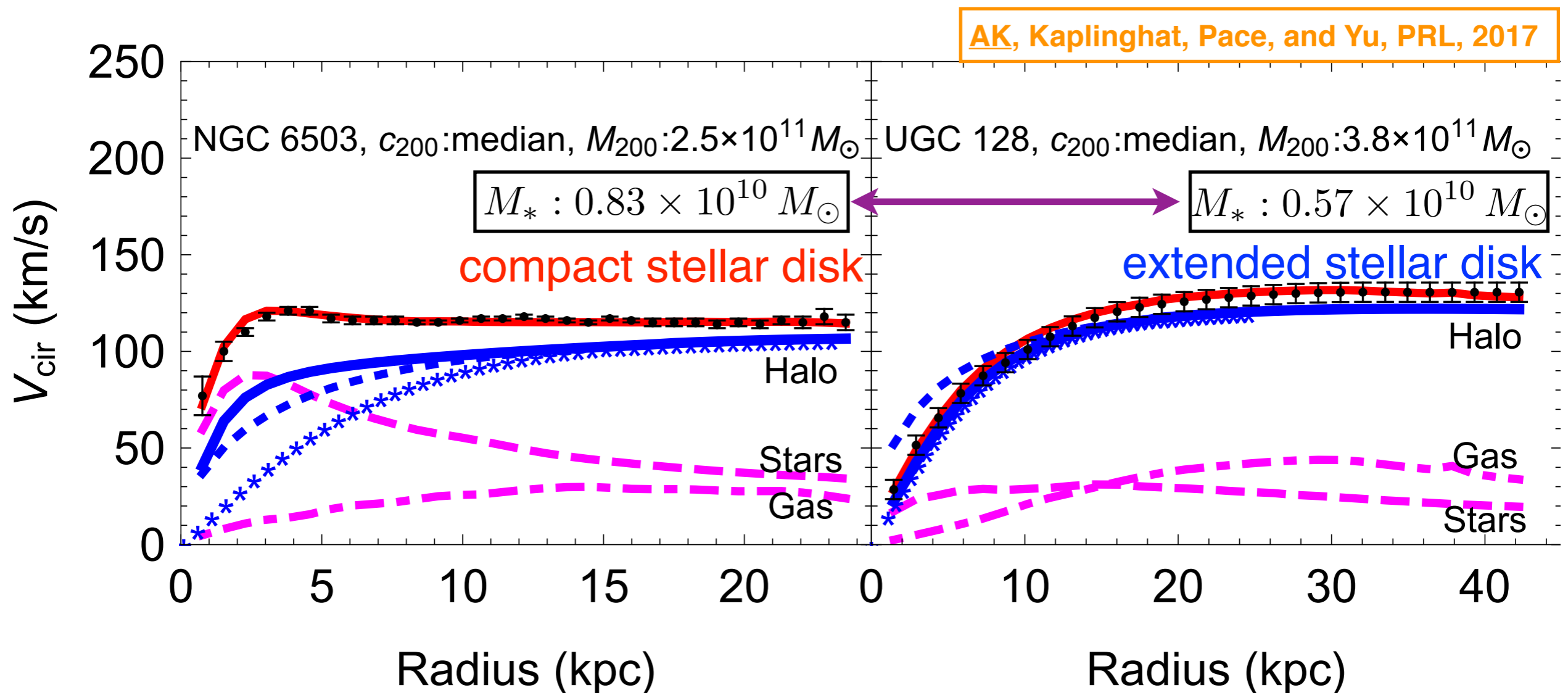
Diversity in stellar distribution

Similar outer circular velocity and stellar mass, but different stellar distribution

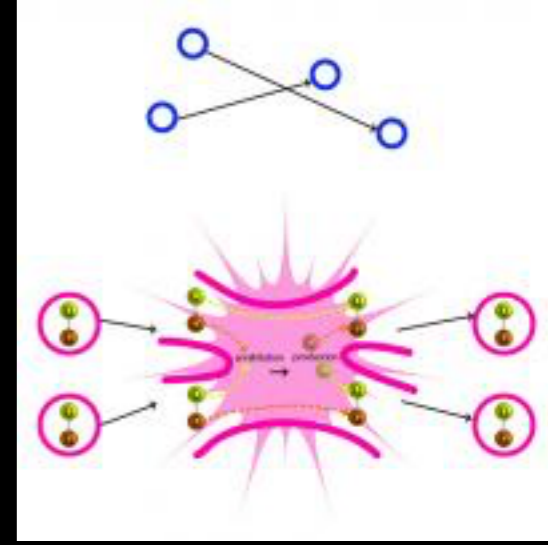
- compact \rightarrow redistribute SIDM significantly
- extended \rightarrow unchange SIDM distribution



Ayuki Kamada

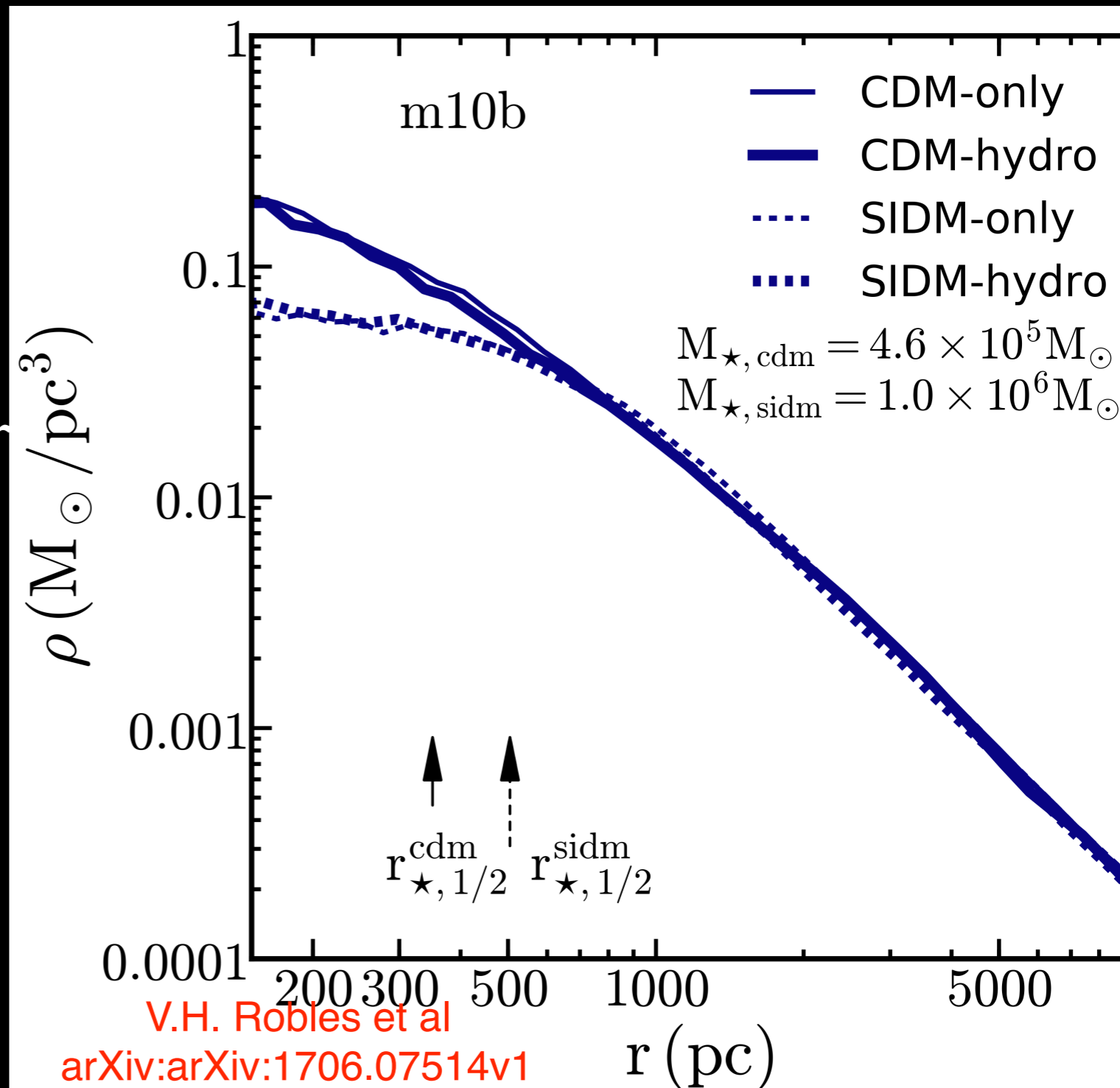


self interaction

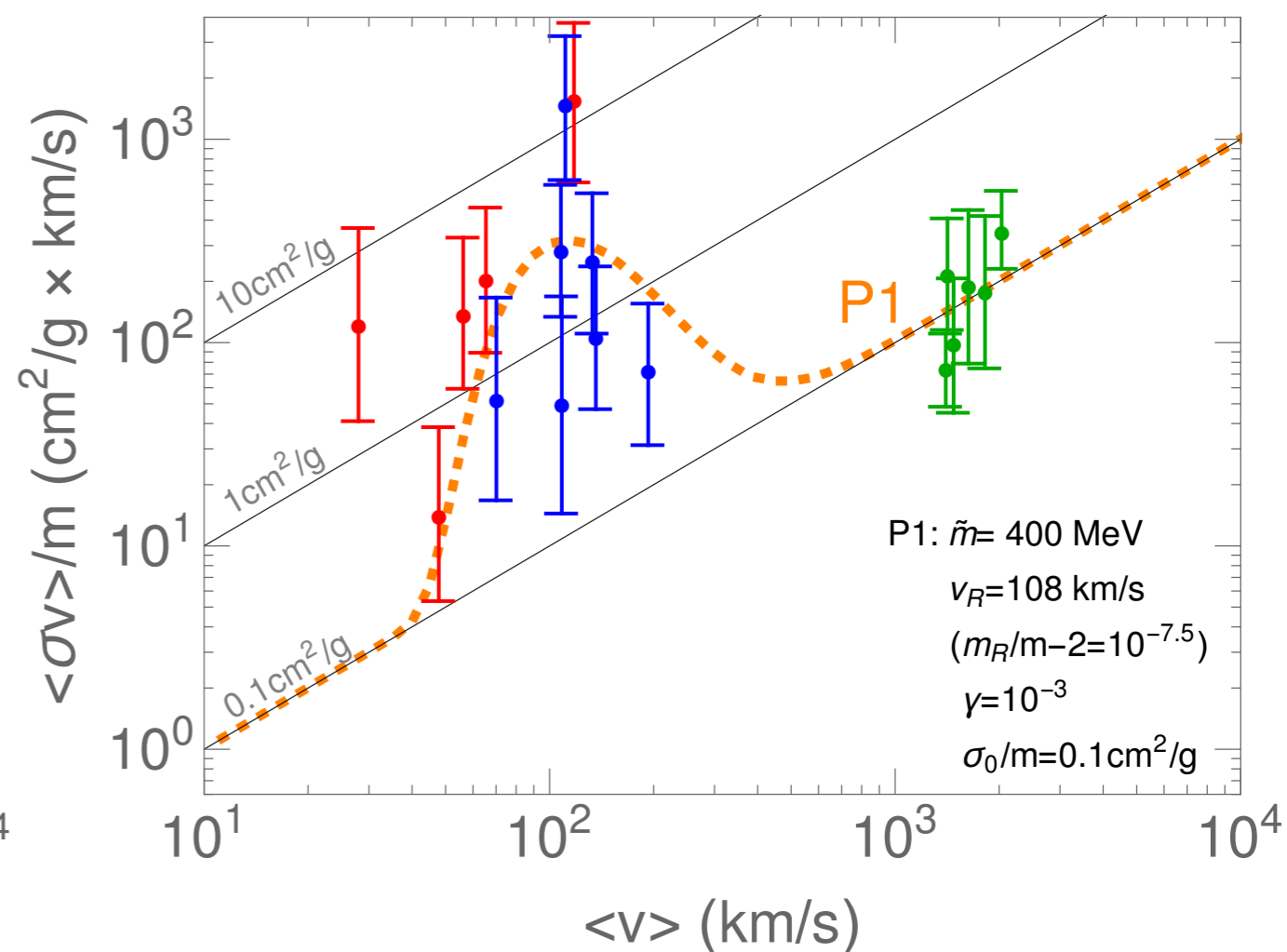
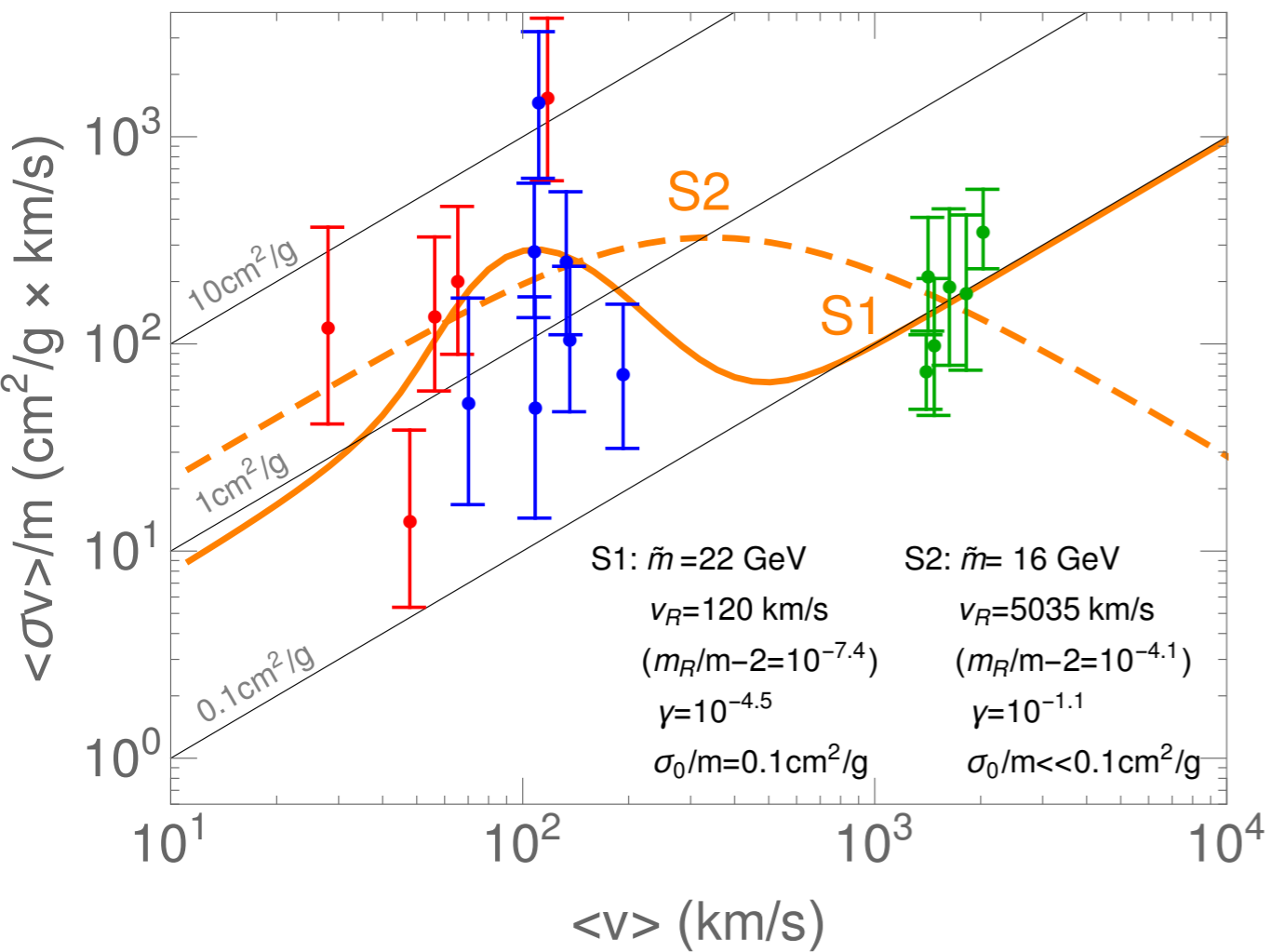


- $\sigma/m \sim \text{cm}^2/\text{g}$
 $\sim 10^{-24} \text{cm}^2 / 300 \text{MeV}$
- flattens the cusps in NFW profile
- suppresses substructure
- actually desirable for dwarf galaxies?

SIDM
Spergel & Steinhardt
(2000)
now complete theory



Resonant scattering



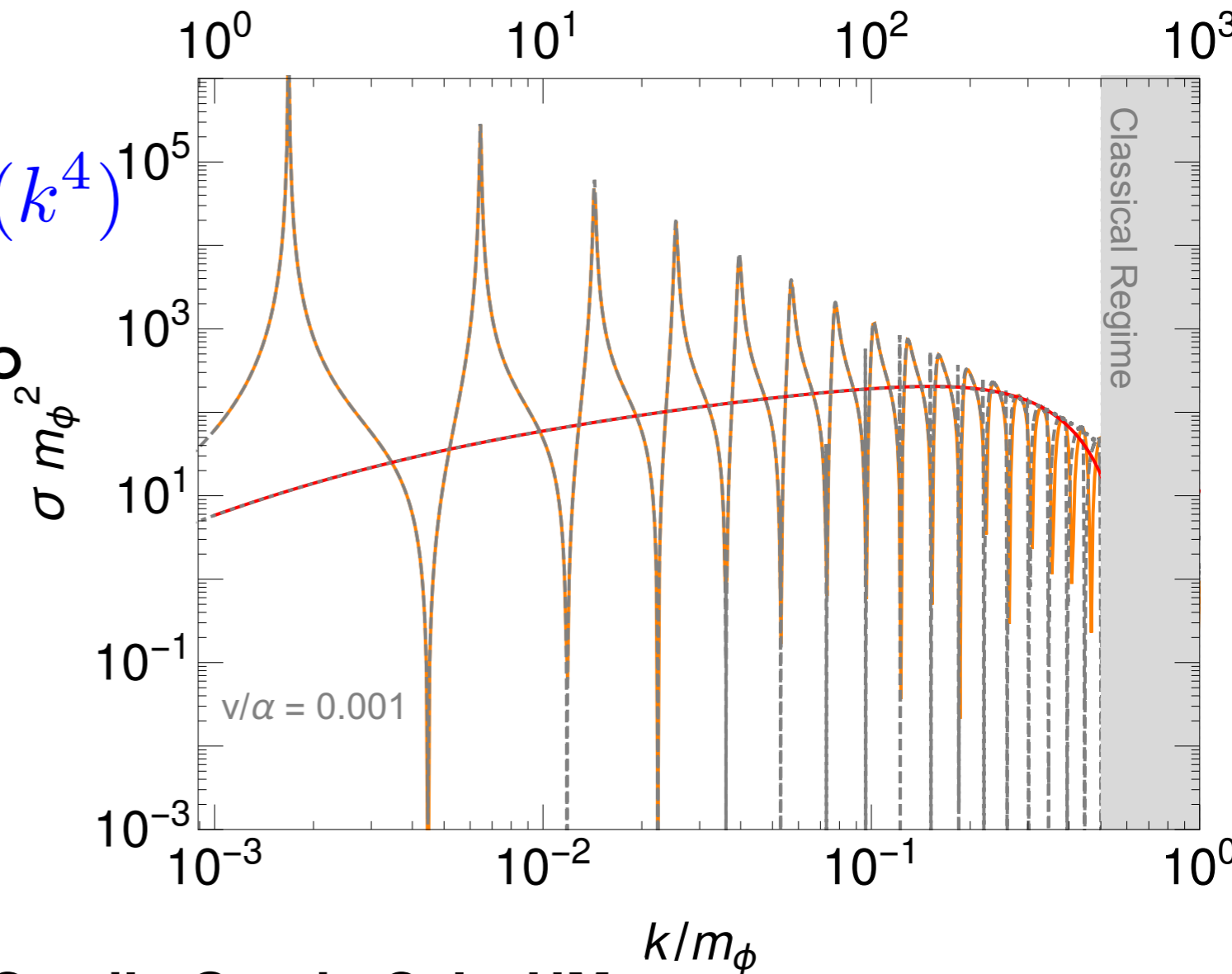
Unified description of SIDM

- Hans Bethe: effective range theory

$$k \cot \delta = -\frac{1}{a} + \frac{1}{2} r_e k^2 + O(k^4)$$

- only two parameters to describe scattering at low velocities
- fully unitary and non-perturbative
- ideal for simulations!

$$V = -\alpha \frac{e^{-m_\phi r}}{r}$$

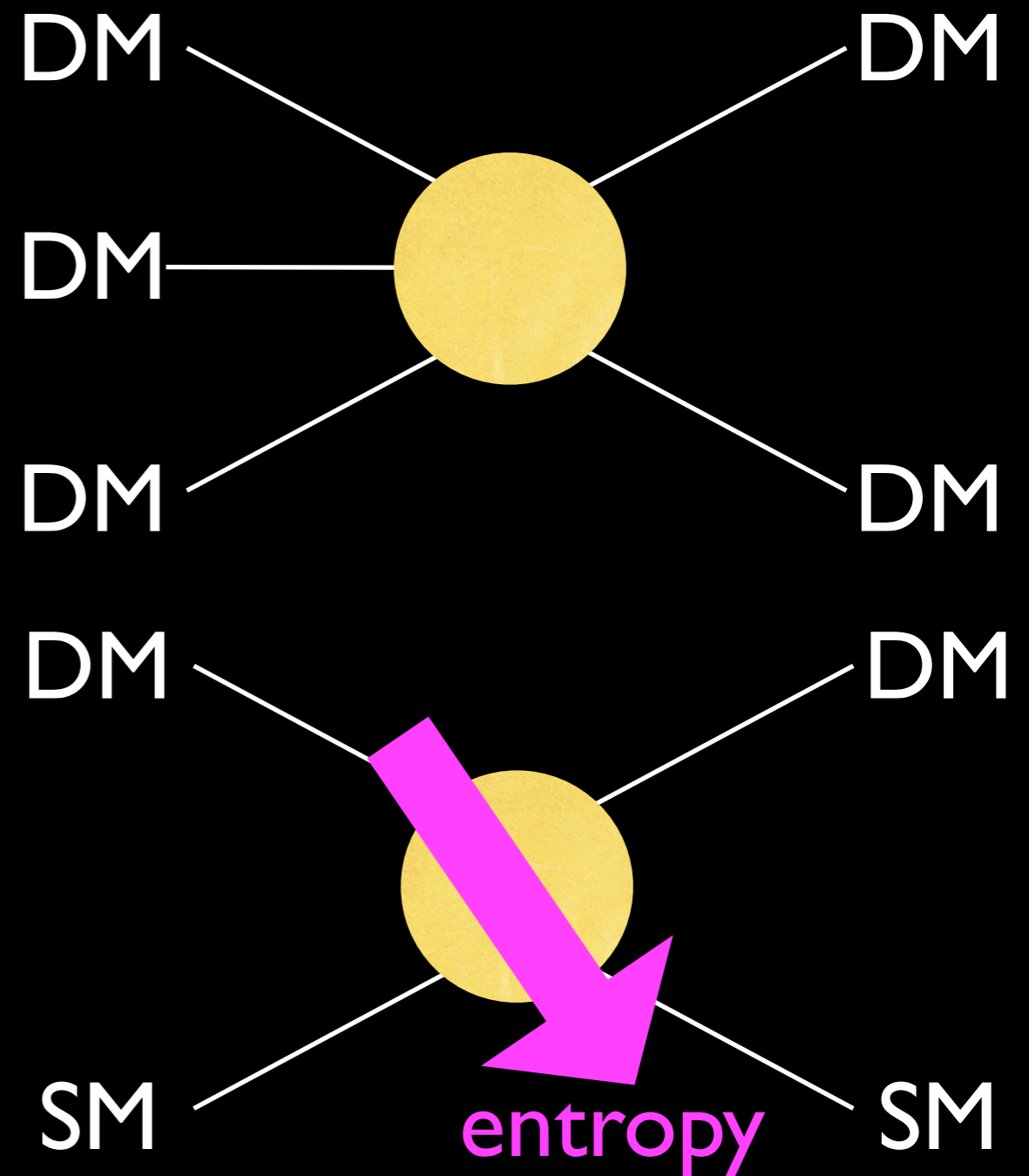


Xiaoyong Chu, Camilo Garcia-Cely, HM,
in preparation

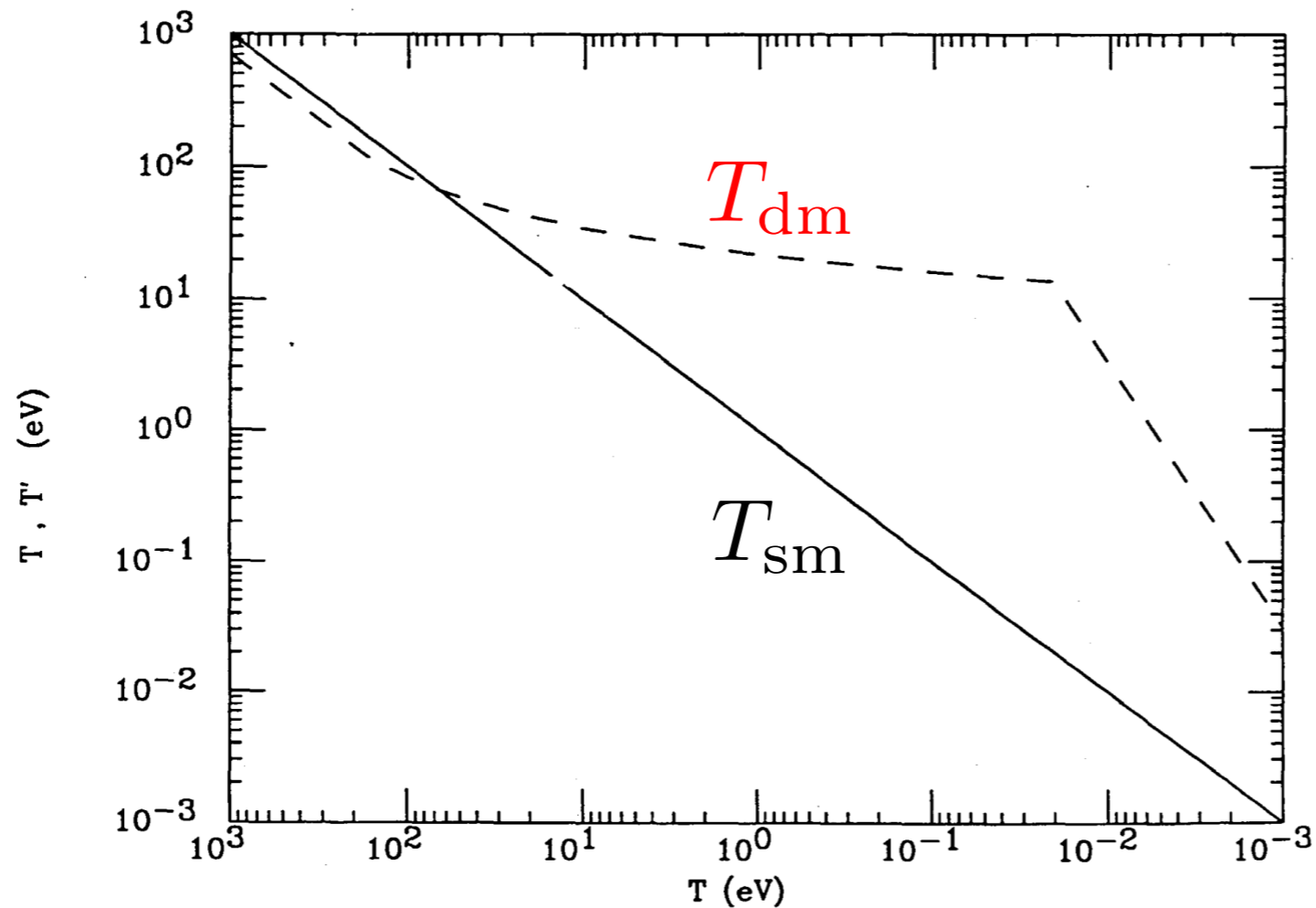


communication

- 3 to 2 annihilation
- excess entropy *must* be transferred to e^\pm, γ
- need communication at some level
- leads to experimental signal



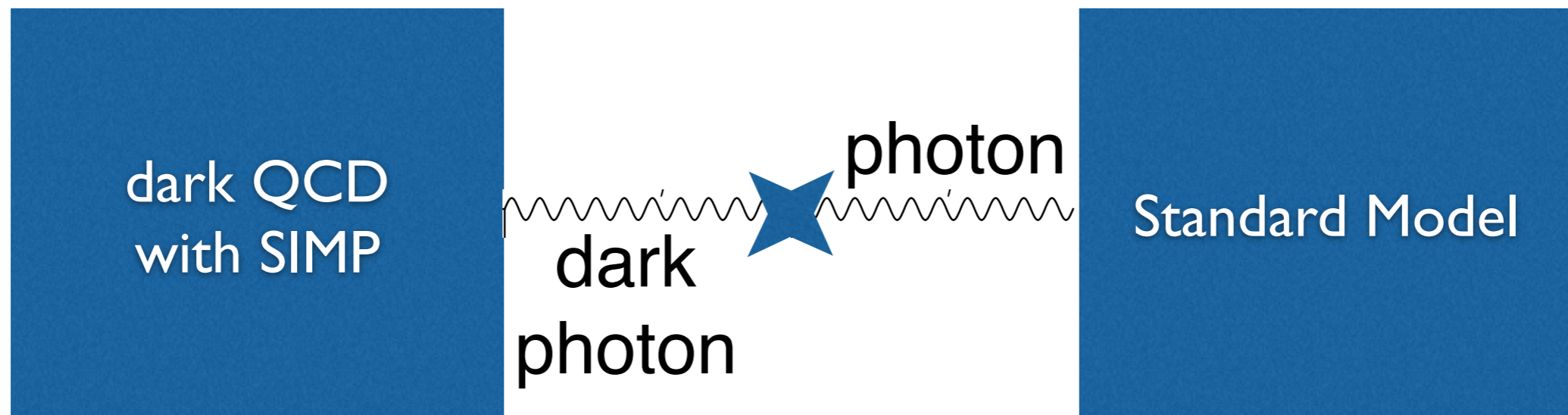
if totally decoupled



Carlson, Hall and Machacek,
Astrophys. J. 398, 43 (1992)

- $3 \rightarrow 2$ annihilations without heat exchange is excluded by structure formation, [de Laix, Scherrer and Schaefer, *Astrophys. J.* 452, 495 (1995)]

vector portal



$$\frac{\epsilon_\gamma}{2c_W} B_{\mu\nu} F_D^{\mu\nu}$$

Kinetically mixed U(1)

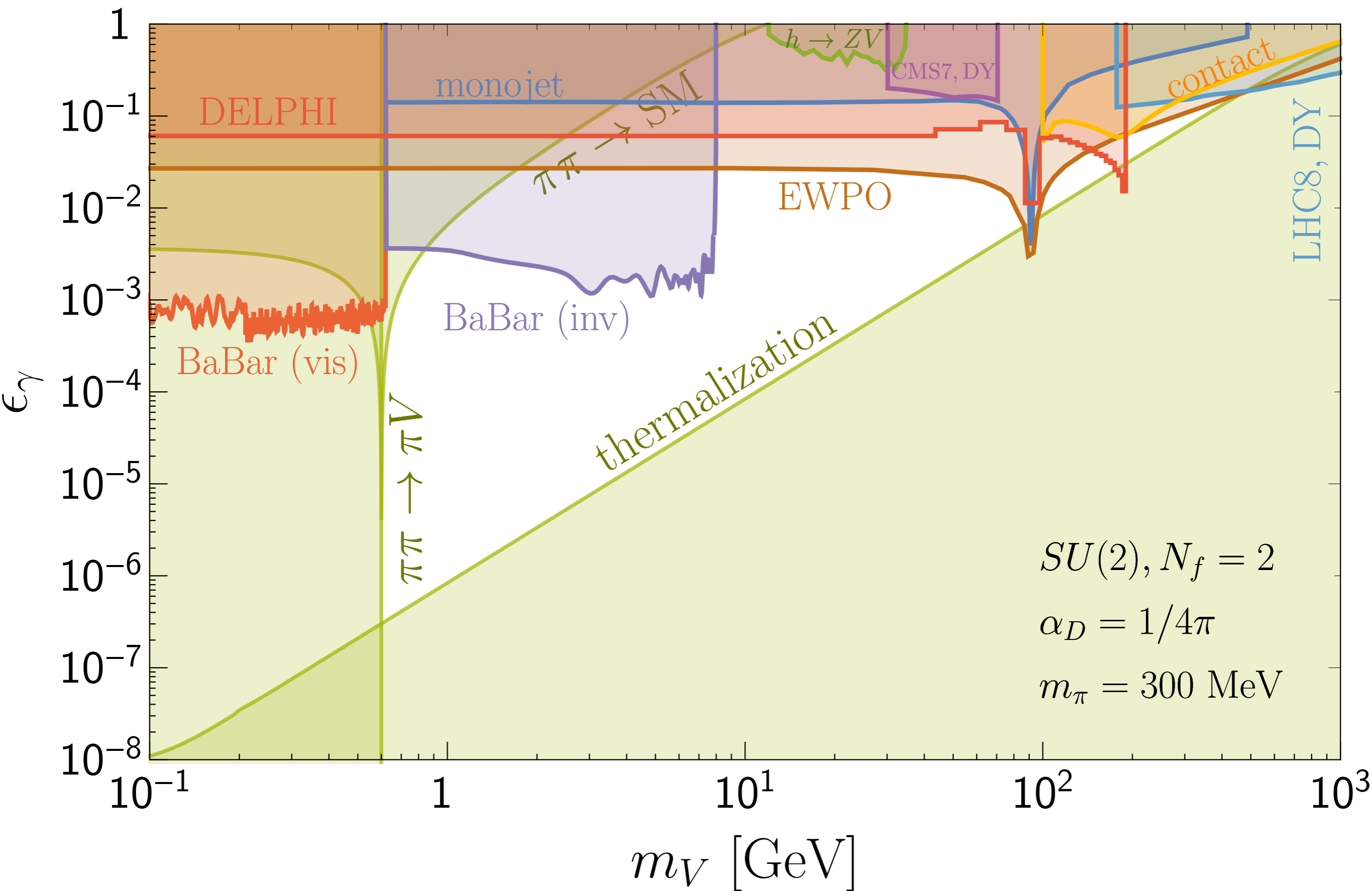
- e.g., the SIMPlest model
SU(2) gauge group with
 $N_f=2$ (4 doublets)
- gauge U(1)=SO(2)
 \subset SO(2) \times SO(3)
 \subset SO(5)=Sp(4)
- maintains degeneracy of quarks
- near degeneracy of pions for co-annihilation

$$SU(4)/Sp(4) = S^5$$

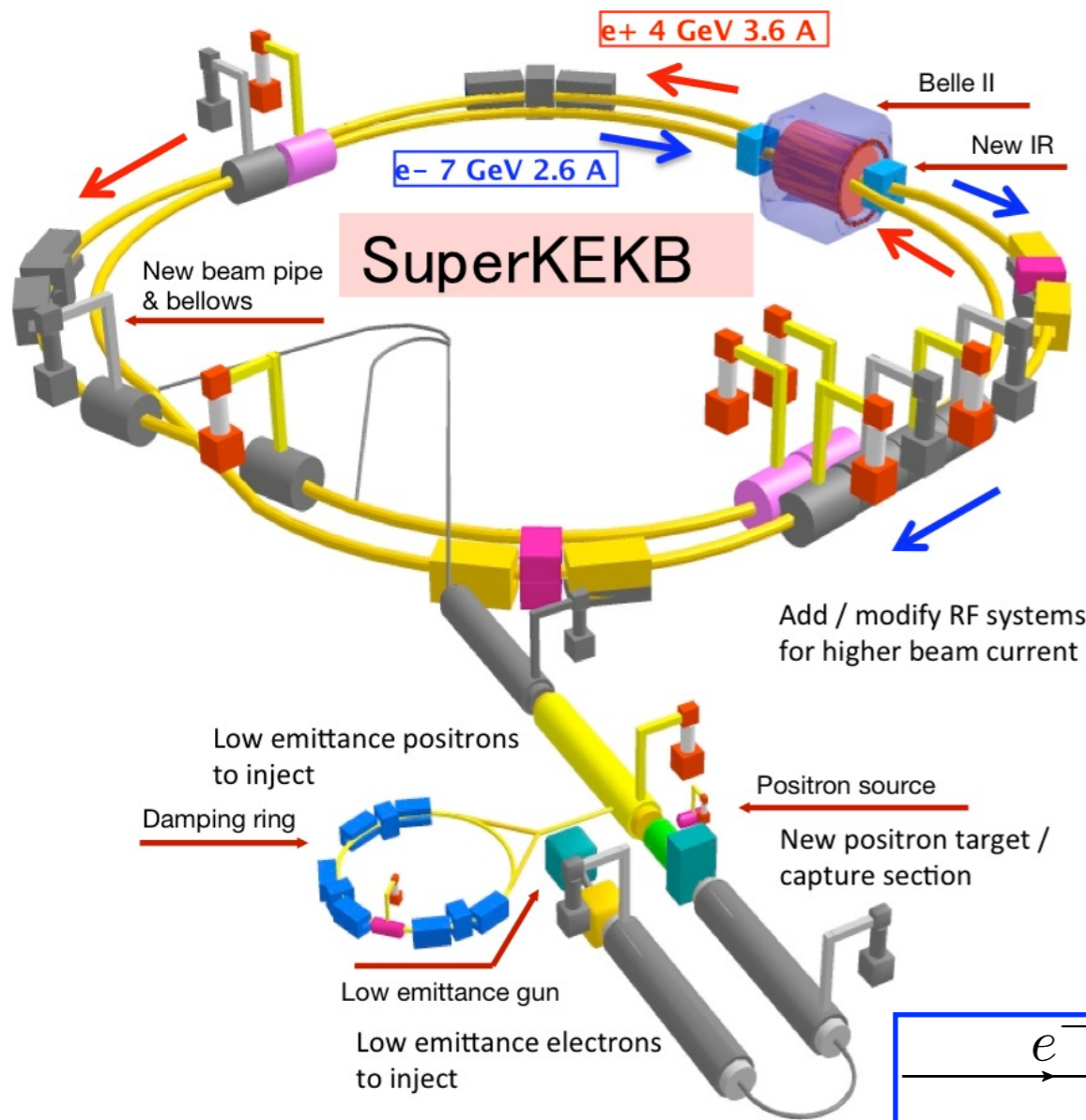
$$(q^+, q^+, q^-, q^-)$$

$$(\pi^{++}, \pi^{--}, \pi_x^0, \pi_y^0, \pi_z^0)$$

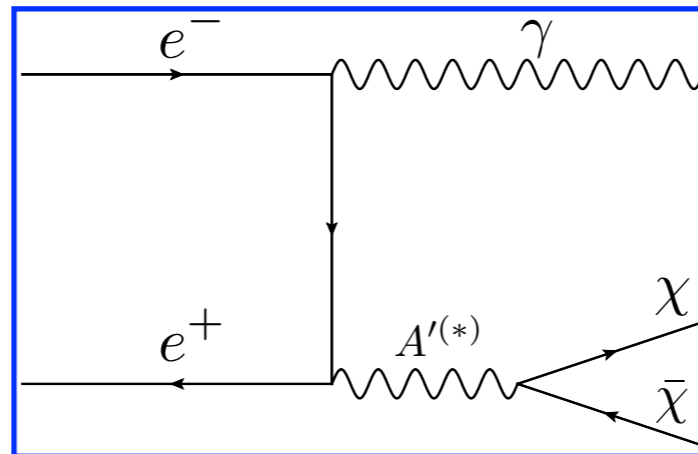
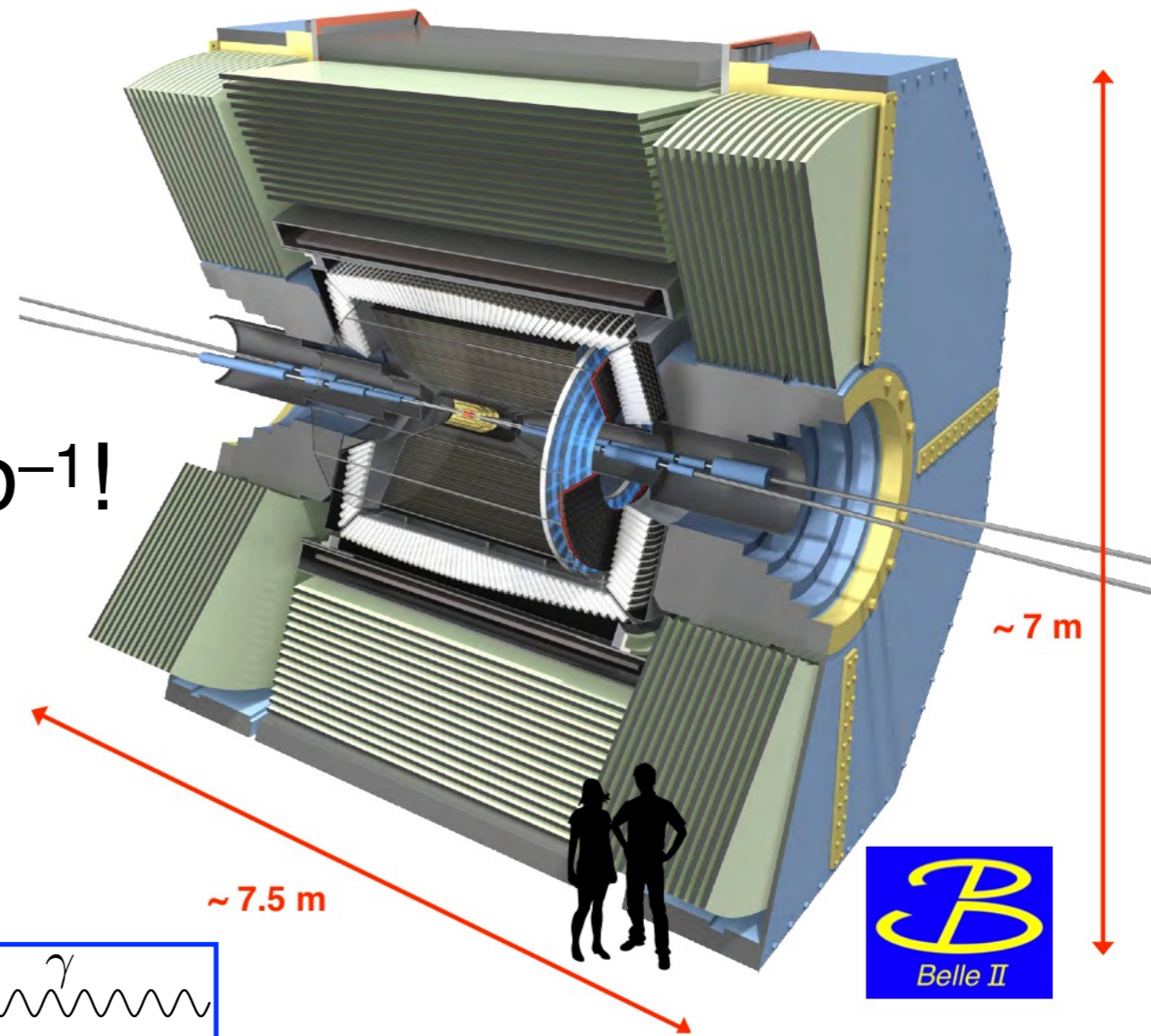
$$\frac{\epsilon_\gamma}{2c_W} B_{\mu\nu} F_D^{\mu\nu}$$



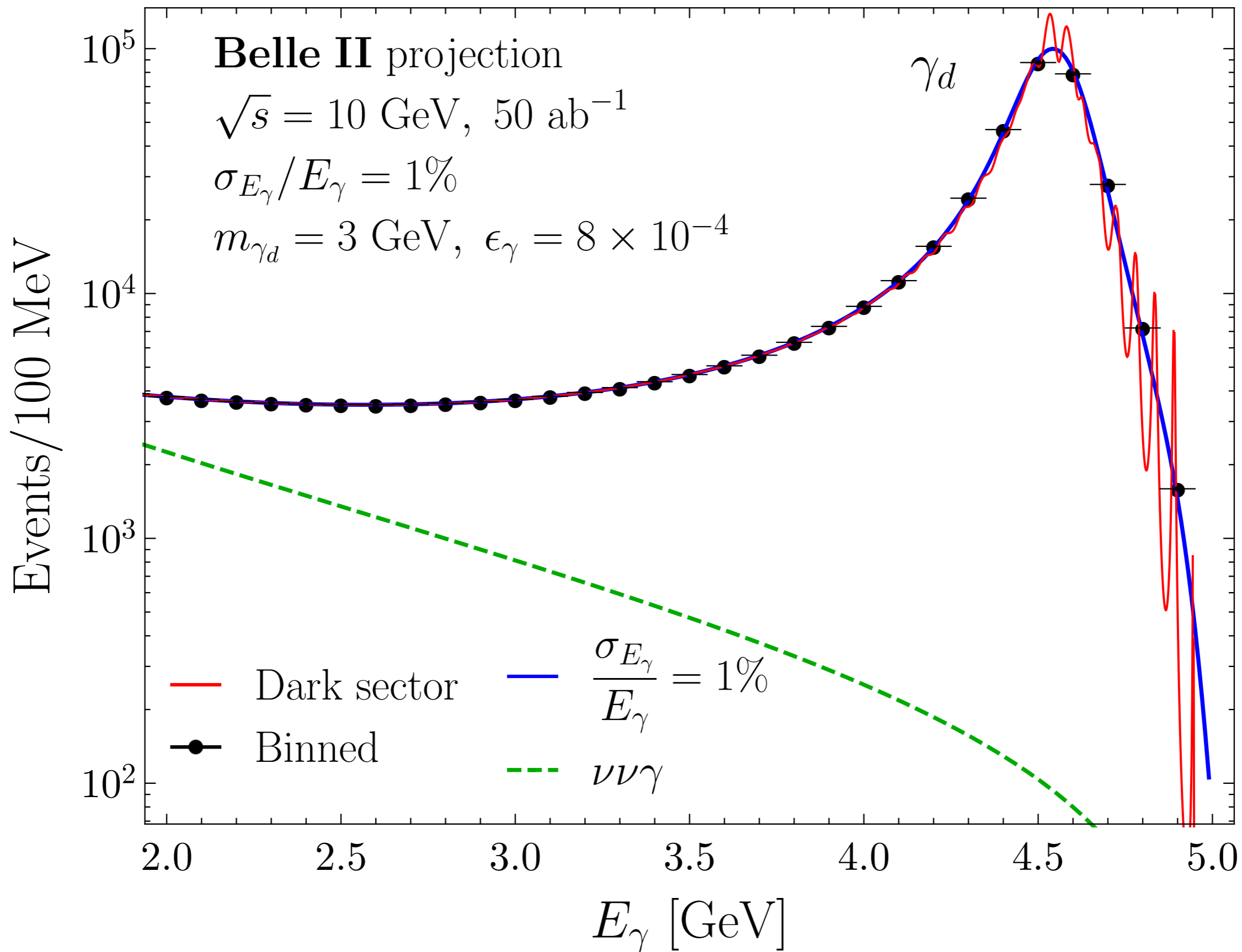
Super KEK B & Belle II

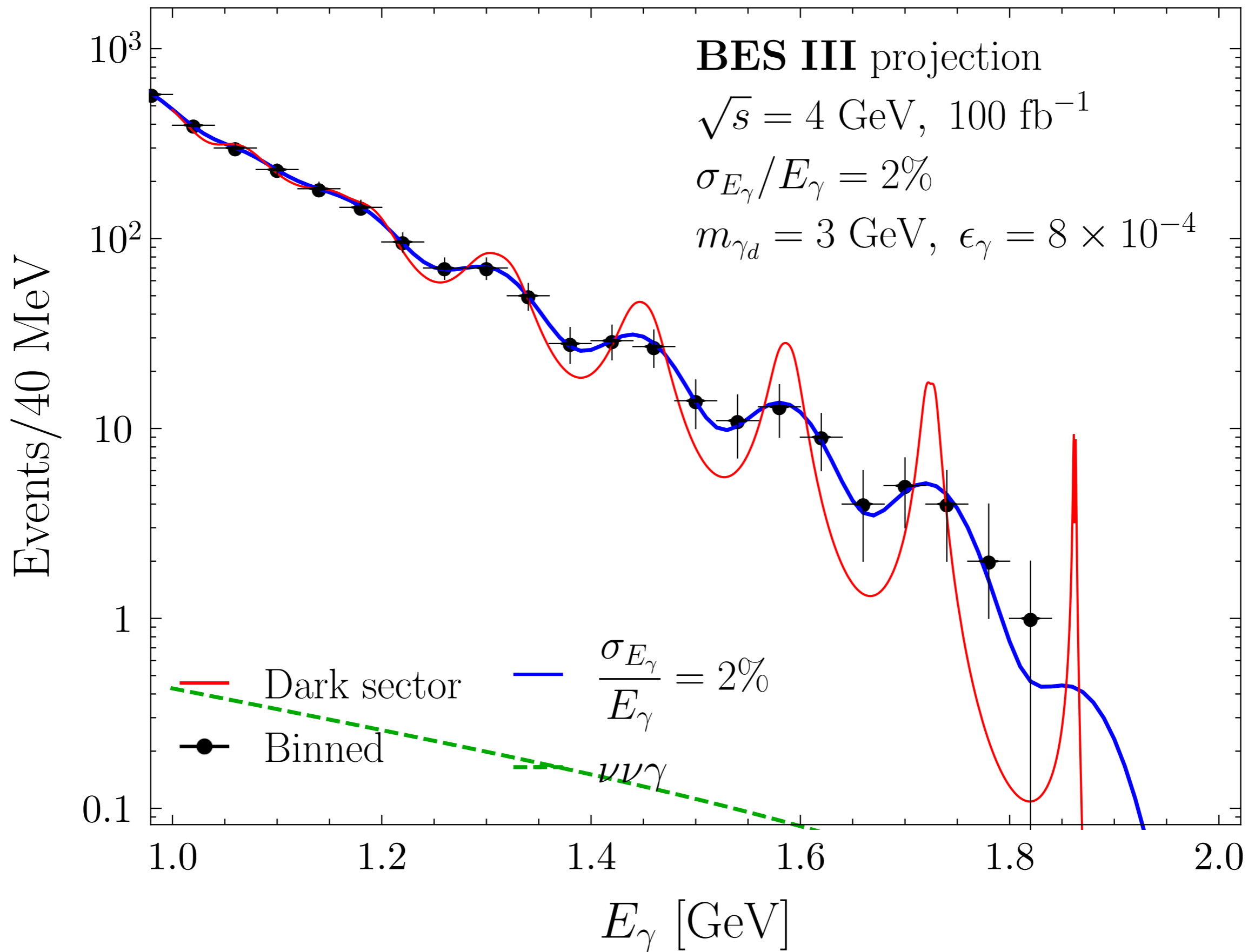


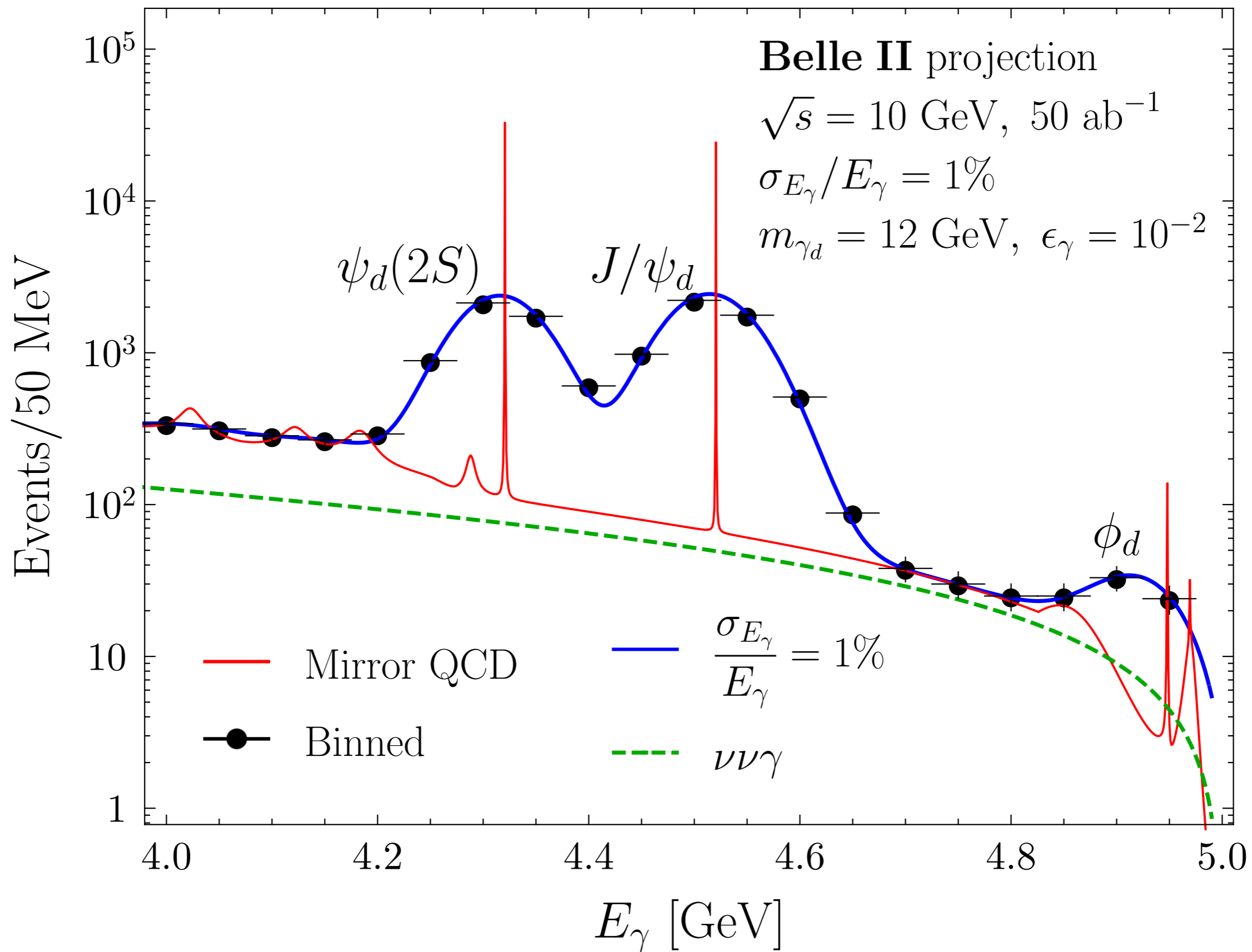
50 ab^{-1} !



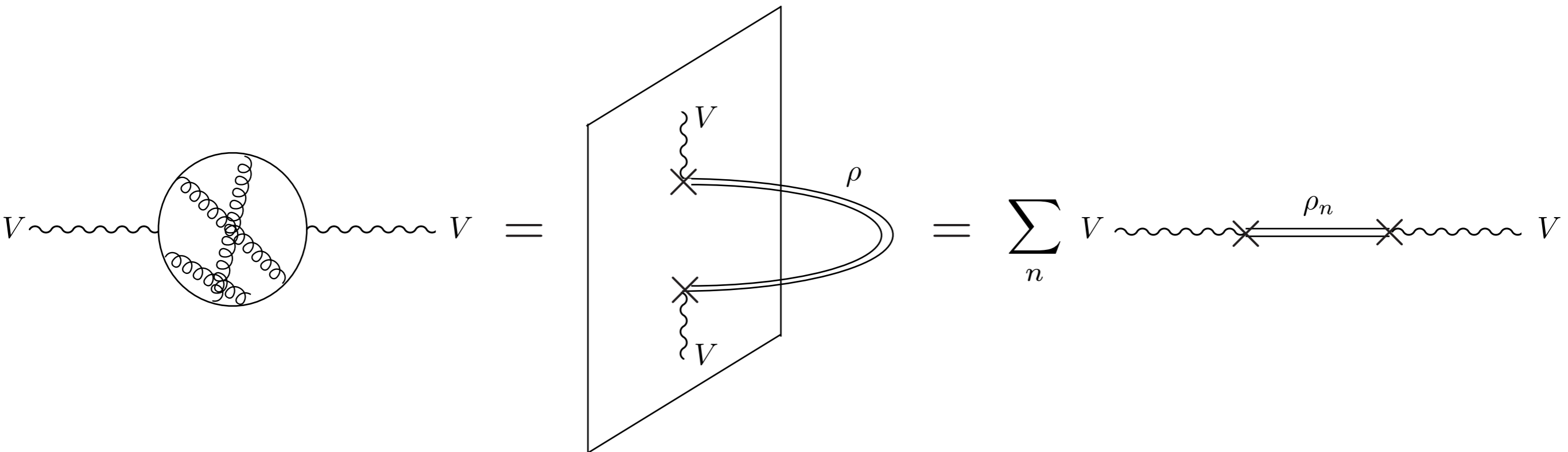
$$E_\gamma = \frac{\sqrt{s}}{2} \left(1 - \frac{M_{\text{inv}}^2}{s} \right)$$



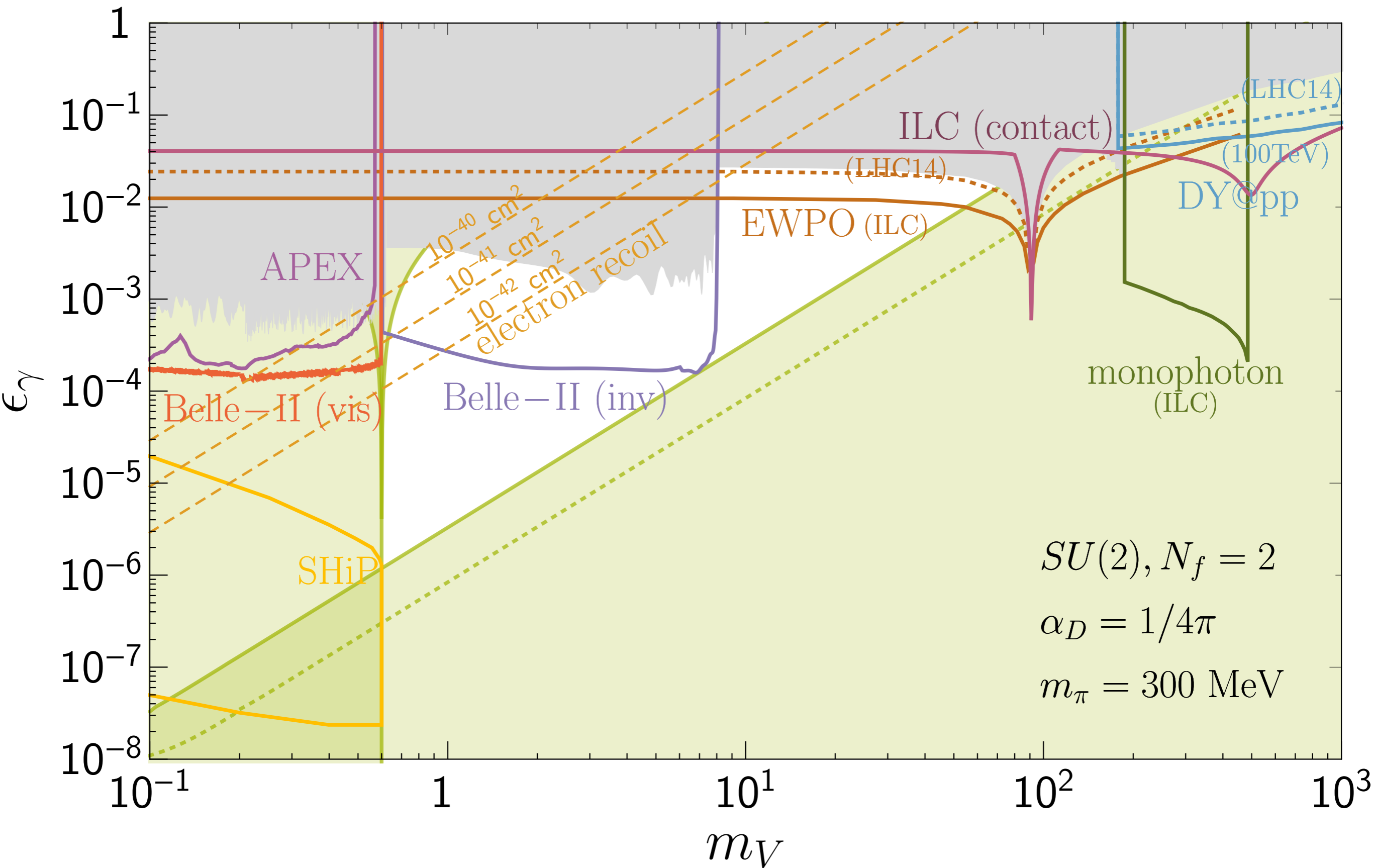




Holographic QCD



inspired by AdS/CFT from string theory



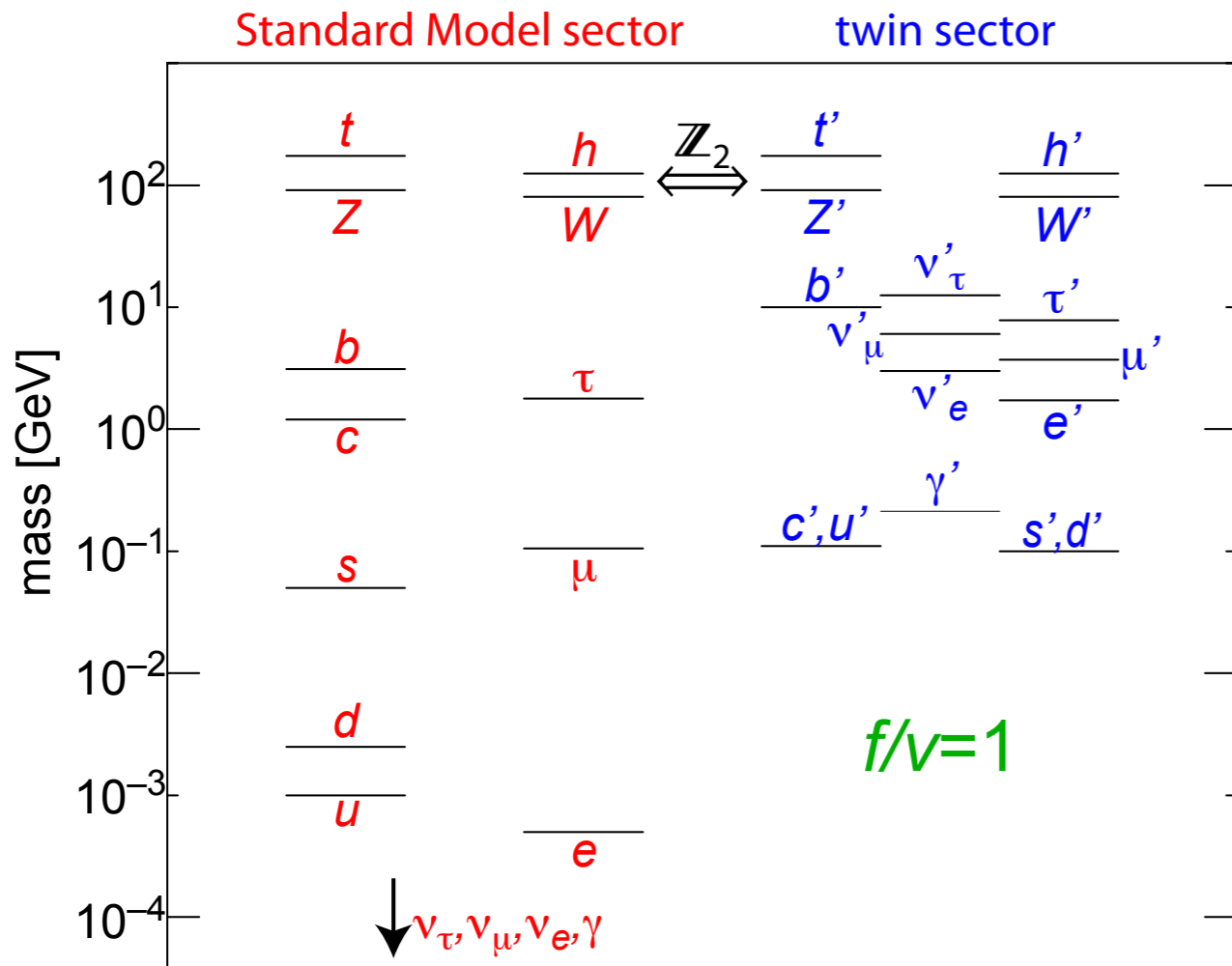


FIG. 1. A sample spectrum of twin particles. Here we use $f/v = 1$ to demonstrate the \mathbb{Z}_2 invariance between the visible and twin sectors for t, h, Z, W ; lighter particles are subject to \mathbb{Z}_2 -breaking effects without spoiling the solution to the hierarchy problem. In practice, twin sector masses are of course raised by a factor of $f/v \gtrsim 3$.

They are stable since they are the lightest particle with a conserved $SU(2)_f$ quantum number. (Here and below, we denote particles in the twin sector with a prime on the

meson M	particle content	$m_M^2 \propto$	m_M
$\theta^0(\mathbf{3}, \mathbf{1})$	$u'\bar{c}', c'\bar{u}', \frac{1}{\sqrt{2}}(u'\bar{u}' - c'\bar{c}')$	$2m_{u'}$	$m_\pi(1 + \Delta)$
$D^+(\mathbf{2}, \mathbf{2})$	$u'\bar{d}', c'\bar{d}', u'\bar{s}', c'\bar{s}'$	$m_{u'} + m_{d'}$	$m_\pi(1 + \frac{\Delta}{2})$
$D^-(\mathbf{2}, \mathbf{2})$	$d'\bar{u}', s'\bar{u}', d'\bar{c}', s'\bar{c}'$	$m_{u'} + m_{d'}$	$m_\pi(1 + \frac{\Delta}{2})$
$\eta^0(\mathbf{1}, \mathbf{1})$	$\frac{1}{2}(d'\bar{d}' + s'\bar{s}' - u'\bar{u}' - c'\bar{c}')$	$m_{u'} + m_{d'}$	$m_\pi(1 + \frac{\Delta}{2})$
$\pi^0(\mathbf{1}, \mathbf{3})$	$d'\bar{s}', s'\bar{d}', \frac{1}{\sqrt{2}}(d'\bar{d}' - s'\bar{s}')$	$2m_{d'}$	m_π

TABLE I. Decomposition of the meson $SU(4)_f$ 15-plet under $SU(2)_U \times SU(2)_D \times U(1)_{EM}$. The third column shows the linear combination of quark masses that determines the meson masses-squared. From top to bottom, the meson masses go from heaviest to lightest, assuming $m_{d'} = m_{s'} < m_{u'} = m_{c'} = m_{d',s'}(1 + \Delta)$.

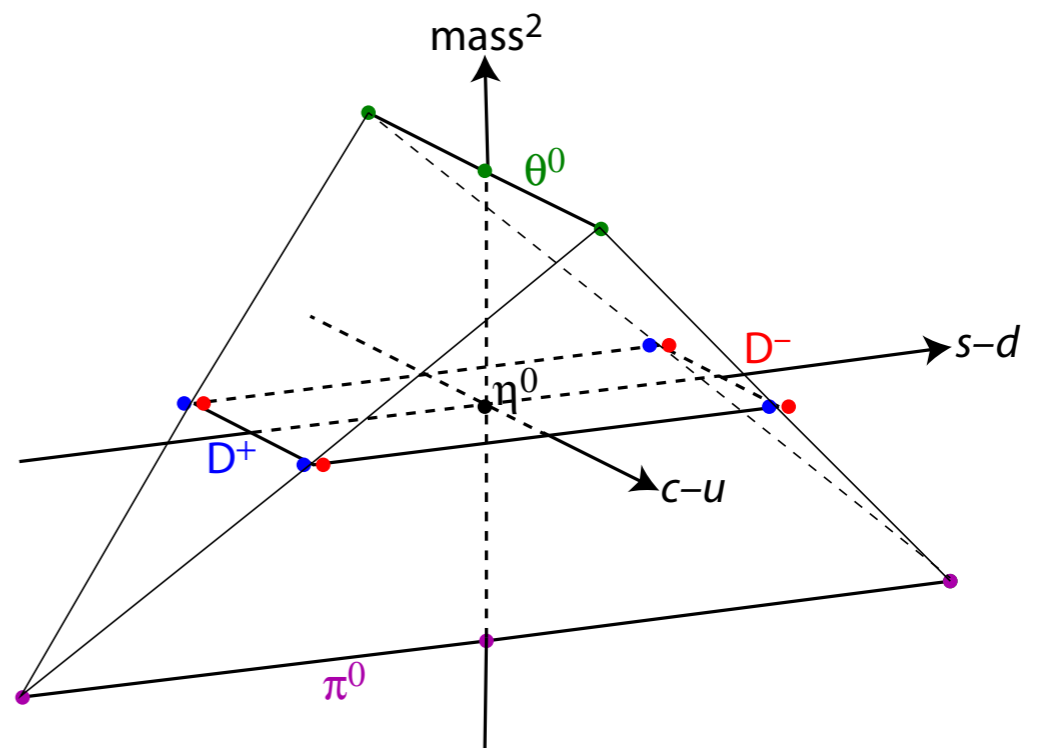


FIG. 2. A visual representation of the meson spectrum.

Conclusion

- surprisingly an *old* theory for dark matter
- SIMP Miracle³
 - mass \sim QCD
 - coupling \sim QCD
 - theory \sim QCD
- can solve problem with DM profile
- very rich phenomenology
- can also be spin 1, axion mediation
- can be a part of twin Higgs
- Exciting *dark spectroscopy!*

Effective operators

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with Brian Henning, Xiaochuan Lu, Thomas Melia



no sign of new physics

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: March 2017

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [fb^{-1}]$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1-2 \tau$	2-10 jets/3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.85 TeV	$m(\tilde{q})=m(\tilde{g})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{q}	1.57 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(1^{st} \text{ gen. } \tilde{q})=m(2^{nd} \text{ gen. } \tilde{q})$	ATLAS-CONF-2017-022
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	3.2	\tilde{q}	608 GeV	$m(\tilde{q})-m(\tilde{\chi}_1^0) < 5$ GeV	1604.07773
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.02 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0 \rightarrow qqW^\pm\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.01 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}^\pm)=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\ell\ell/\nu\nu)\tilde{\chi}_1^0$	3 e, μ	4 jets	-	13.2	\tilde{g}	1.7 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2016-037
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	2 e, μ (SS)	0-3 jets	Yes	13.2	\tilde{g}	1.6 TeV	$m(\tilde{\chi}_1^0) < 500$ GeV	ATLAS-CONF-2016-037
	GMSB ($\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	\tilde{g}	2.0 TeV		1607.05979
	GGM (bino NLSP)	2 γ	-	Yes	3.2	\tilde{g}	1.65 TeV	$c\tau(\text{NLSP}) < 0.1$ mm	1606.09150
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3	\tilde{g}	37 TeV	$m(\tilde{\chi}_1^0) < 950$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu < 0$	1507.05493
3 rd gen. \tilde{g} med.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	36.1	\tilde{g}	1.92 TeV	$m(\tilde{\chi}_1^0) < 600$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	36.1	\tilde{g}	1.97 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^+$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	37 TeV	$m(\tilde{\chi}_1^0) < 300$ GeV	1407.0600
3 rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	3.2	\tilde{b}_1	840 GeV	$m(\tilde{\chi}_1^0) < 100$ GeV	1606.08772
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^\pm$	2 e, μ (SS)	1 b	Yes	13.2	\tilde{b}_1	325-685 GeV	$m(\tilde{\chi}_1^0) < 150$ GeV, $m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_1^0) + 100$ GeV	ATLAS-CONF-2016-037
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^\pm$	0-2 e, μ	1-2 b	Yes	4.7/13.3	\tilde{t}_1	117-170 GeV	$m(\tilde{\chi}_1^\pm) = 2m(\tilde{\chi}_1^0)$, $m(\tilde{\chi}_1^0) = 55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $\tilde{t}\tilde{\chi}_1^0$	0-2 e, μ	0-2 jets/1-2 b	Yes	20.3	\tilde{t}_1	90-198 GeV	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, ATLAS-CONF-2017-020
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	3.2	\tilde{t}_1	90-323 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0) = 5$ GeV	1604.07773
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-600 GeV	$m(\tilde{\chi}_1^0) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	36.1	\tilde{t}_2	290-790 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	ATLAS-CONF-2017-019
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	$\tilde{\ell}$	90-335 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	1403.5294
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\ell}\nu(\tilde{\ell}\bar{\nu})$	2 e, μ	0	Yes	13.3	$\tilde{\chi}_1^\pm$	640 GeV	$m(\tilde{\chi}_1^\pm) = 0$ GeV, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-096
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow \tilde{\tau}\nu(\tilde{\tau}\bar{\nu})$	2 τ	-	Yes	14.8	$\tilde{\chi}_1^\pm$	580 GeV	$m(\tilde{\chi}_1^\pm) = 0$ GeV, $m(\tilde{\tau}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-093
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm \rightarrow \tilde{\ell}_L\nu\tilde{\ell}_L, \tilde{\ell}(\bar{\nu})$	3 e, μ	0	Yes	13.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	1.0 TeV	$m(\tilde{\chi}_1^\pm) = 0$, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-096
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	2-3 e, μ	0-2 jets	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	425 GeV	$m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	1403.5294, 1402.7029
	$\tilde{\chi}_1^\pm\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^0$	270 GeV	$m(\tilde{\chi}_1^\pm) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0 \rightarrow \tilde{\ell}_R\tilde{\ell}_R$	4 e, μ	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	635 GeV	$m(\tilde{\chi}_2^0) = 0$, $m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	115-370 GeV	$c\tau < 1$ mm	1507.05493
	GGM (bino NLSP) weak prod.	2 γ	-	Yes	20.3	\tilde{W}	590 GeV	$c\tau < 1$ mm	1507.05493
	Long-lived particles	Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	36.1	$\tilde{\chi}_1^\pm$	430 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm) = 0.2$ ns
Direct $\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm$ prod., long-lived $\tilde{\chi}_1^\pm$		dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^\pm$	495 GeV	$m(\tilde{\chi}_1^\pm)-m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^\pm) < 15$ ns	1506.05332
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9	\tilde{g}	850 GeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $10 \mu\text{s} < \tau(\tilde{g}) < 1000$ s	1310.6584
Stable \tilde{g} R-hadron		trk	-	-	3.2	\tilde{g}	1.58 TeV		1606.05129
Metastable \tilde{g} R-hadron		dE/dx trk	-	-	3.2	\tilde{g}	1.57 TeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $\tau > 10$ ns	1604.04520
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{\ell}, \bar{\mu}) + \tau(e, \mu)$		1-2 μ	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$	1411.6795
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3	$\tilde{\chi}_1^0$	440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow ee\nu/e\nu\mu/\mu\nu$		displ. $ee/e\mu/\mu\mu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g}) = 1.3$ TeV	1504.05162
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$		displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480$ mm, $m(\tilde{g}) = 1.1$ TeV	1504.05162
RPV		LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\epsilon\tau/\mu\tau$	$e\mu, \epsilon\tau, \mu\tau$	-	-	3.2	$\tilde{\nu}_\tau$	1.9 TeV	$\lambda'_{311} = 0.11, \lambda_{132/133/233} = 0.07$
	Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{q}) = m(\tilde{g}), c\tau_{LS} P < 1$ mm	1404.2500
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow ee\nu, e\mu\nu, \mu\mu\nu$	4 e, μ	-	Yes	13.3	$\tilde{\chi}_1^\pm$	1.14 TeV	$m(\tilde{\chi}_1^0) > 400$ GeV, $\lambda_{12k} \neq 0$ ($k = 1, 2$)	ATLAS-CONF-2016-075
	$\tilde{\chi}_1^\pm\tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W\tilde{\chi}_1^0\tilde{\chi}_1^0 \rightarrow \tau\tau\nu_e, \epsilon\tau\nu_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^\pm$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^\pm), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.08 TeV	$BR(\tilde{g}) = BR(\tilde{b}) = BR(\tilde{c}) = 0\%$	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.55 TeV	$m(\tilde{\chi}_1^0) = 800$ GeV	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	2.1 TeV	$m(\tilde{\chi}_1^0) = 1$ TeV, $\lambda_{112} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	1.65 TeV	$m(\tilde{t}_1) = 1$ TeV, $\lambda_{323} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	15.4	\tilde{t}_1	410 GeV 450-510 GeV		ATLAS-CONF-2016-022, ATLAS-CONF-2016-084
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\ell}$	2 e, μ	2 b	-	20.3	\tilde{t}_1	0.4-1.0 TeV	$BR(\tilde{t}_1 \rightarrow b\tilde{\ell}/\mu) > 20\%$	ATLAS-CONF-2015-015
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	510 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹

1

Mass scale [TeV]

"The 2 TeV line has been reached for some scenarios"

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IN

SCIENTISTS NEED A NEW WAY

PHYSICS

TO EXPLAIN THE UNIVERSE

?



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why effective operators

- No signal of BSM @ LHC so far
- use effective operators to parametrize physics at higher energies
 - precision electroweak
 - precision Higgs
 - precision flavor
 - B, L violation
 - coupling to the dark sector
- once deviation \Rightarrow BSM theory
- similar to four-fermion operators in weak interactions \Rightarrow Standard Model

Effective Operators

- Surprisingly difficult question
- In the case of the Standard Model
 - **Weinberg** (1980) on $D=6$ \not{B} , $D=5$ \not{L}
 - **Buchmüller-Wyler** (1986) on $D=6$ ops
 - 80 operators for $N_f=1$, B , L conserving
 - **Grzadkowski et al** (2010) removed redundancies and discovered one missed
 - 59 operators for $N_f=1$, B , L conserving
 - **Mahonar et al** (2013) general N_f
 - **Lehman-Martin** (2014,15) $D=7$ for general N_f , $D=8$ for $N_f=1$ (incomplete)

$$\begin{aligned}
\widehat{H}_6 = & H^3 H^\dagger{}^3 + u^\dagger Q^\dagger H H^\dagger{}^2 + 2Q^2 Q^\dagger{}^2 + Q^\dagger{}^3 L^\dagger + Q^3 L + 2QQ^\dagger LL^\dagger + L^2 L^\dagger{}^2 + uQH^2 H^\dagger \\
& + 2uu^\dagger QQ^\dagger + uu^\dagger LL^\dagger + u^2 u^\dagger{}^2 + e^\dagger u^\dagger Q^2 + e^\dagger L^\dagger H^2 H^\dagger + 2e^\dagger u^\dagger Q^\dagger L^\dagger + eLHH^\dagger{}^2 + euQ^\dagger{}^2 \\
& + 2euQL + ee^\dagger QQ^\dagger + ee^\dagger LL^\dagger + ee^\dagger uu^\dagger + e^2 e^\dagger{}^2 + d^\dagger Q^\dagger H^2 H^\dagger + 2d^\dagger u^\dagger Q^\dagger{}^2 + d^\dagger u^\dagger QL \\
& + d^\dagger e^\dagger u^\dagger{}^2 + d^\dagger e^\dagger Q^\dagger L + dQH H^\dagger{}^2 + 2duQ^2 + duQ^\dagger L^\dagger + de^\dagger QL^\dagger + deu^2 + 2dd^\dagger QQ^\dagger + dd^\dagger LL^\dagger \\
& + 2dd^\dagger uu^\dagger + dd^\dagger ee^\dagger + d^2 d^\dagger{}^2 + u^\dagger Q^\dagger H^\dagger G_R + d^\dagger Q^\dagger H G_R + HH^\dagger G_R^2 + G_R^3 + uQH G_L \\
& + dQH^\dagger G_L + HH^\dagger G_L^2 + G_L^3 + u^\dagger Q^\dagger H^\dagger W_R + e^\dagger L^\dagger H W_R + d^\dagger Q^\dagger H W_R + HH^\dagger W_R^2 + W_R^3 \\
& + uQH W_L + eLH^\dagger W_L + dQH^\dagger W_L + HH^\dagger W_L^2 + W_L^3 + u^\dagger Q^\dagger H^\dagger B_R + e^\dagger L^\dagger H B_R \\
& + d^\dagger Q^\dagger H B_R + HH^\dagger B_R W_R + HH^\dagger B_R^2 + uQH B_L + eLH^\dagger B_L + dQH^\dagger B_L + HH^\dagger B_L W_L \\
& + HH^\dagger B_L^2 + 2QQ^\dagger H H^\dagger \mathcal{D} + 2LL^\dagger H H^\dagger \mathcal{D} + uu^\dagger H H^\dagger \mathcal{D} + ee^\dagger H H^\dagger \mathcal{D} + d^\dagger u H^2 \mathcal{D} + du^\dagger H^\dagger{}^2 \mathcal{D} \\
& + dd^\dagger H H^\dagger \mathcal{D} + 2H^2 H^\dagger{}^2 \mathcal{D}^2 .
\end{aligned} \tag{3.16}$$

\mathcal{D} : space time derivative

redundancies

- effective operators are invariants under the gauge group, Lorentz group, etc
- their classifications go back to Hilbert, Weyl
- applied to superpotentials, Standard Model
- but so far **no general discussions on operators with derivatives**
- two sources of redundancies
 - **equation of motion (EOM)**
 - **integration by parts (IBP)**

Simplest Example

- scalars four-point at $O(p^2)$: $4(4+1)/2=10$

$$(\partial_\mu \partial_\mu \varphi_i) \varphi_j \varphi_k \varphi_l \quad (\partial_\mu \varphi_i) (\partial_\mu \varphi_j) \varphi_k \varphi_l$$

- $\partial^2 \varphi_i = m_i^2 \varphi_i$ removes the first class: 4

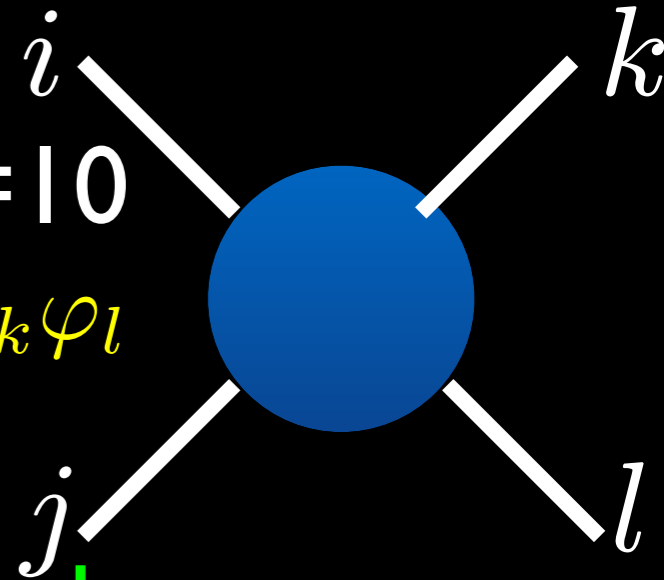
- We know only **2 out of 6 are independent**

- $s, t, u, s+t+u = m_1^2 + m_2^2 + m_3^2 + m_4^2$

$$(\partial_\mu \varphi_i) (\partial_\mu \varphi_j) \varphi_k \varphi_l - \varphi_i \varphi_j (\partial_\mu \varphi_k) (\partial_\mu \varphi_l) = \frac{1}{2} \partial^2 (\varphi_i \varphi_j) (\varphi_k \varphi_l) - \frac{1}{2} (\varphi_i \varphi_j) \partial^2 (\varphi_k \varphi_l) \approx 0$$

$$\partial_\mu \varphi_i \partial_\mu \varphi_j \varphi_k \varphi_l + \partial_\mu \varphi_i \varphi_j \partial_\mu \varphi_k \varphi_l + \partial_\mu \varphi_i \varphi_j \varphi_k \partial_\mu \varphi_l = \partial_\mu \varphi_i \partial_\mu (\varphi_j \varphi_k \varphi_l) \approx 0$$

- In addition, there are only d linearly independent momenta in d -dimensions for higher-point functions



Main idea

- Take kinetic terms as the zeroth order
Lagrangian $(\partial\phi)^2$, $\bar{\psi}i\not{\partial}\psi$, $(F_{\mu\nu})^2$
- Classically, it is conformally invariant under $SO(4,2) \simeq SO(6, \mathbb{C})$
- Operator-State correspondence in CFT tells us that operators fall into representations of the conformal group
 - **equation of motion**: short multiplets
 - **remove total derivatives**: primary states

Master formula

- Define a multi-variate Hilbert series

$$H(p, \phi_1, \dots, \phi_n) = \int d\mu_{\text{conformal}} d\mu_{\text{gauge}} \sum_{n=1}^{\infty} p^n \chi_{[n;0]}^* \prod_i PE[\phi_i \chi_i(q, \alpha, \beta)]$$

- integration over the gauge groups pick up gauge invariants
- integration over the conformal group picks only the primary states and Lorentz scalars
- expand it in power series in ϕ_i and p to find operators at given order in them
- Possible for any Lorentz-inv “free” QFT

*There are corrections for operators $d \leq 4$ due to lack of orthonormality among characters for short multiplets

Standard Model

```

χH[t_, α_, β_, x_, y_, z1_, z2_] := χscal[t, α, β] * u1[3, x] * su2f[y];
χHd[t_, α_, β_, x_, y_, z1_, z2_] := χscal[t, α, β] * u1[-3, x] * su2fb[y];
χQ[t_, α_, β_, x_, y_, z1_, z2_] := χfermL[t, α, β] * u1[1, x] * su2f[y] * su3f[z1, z2];
χQd[t_, α_, β_, x_, y_, z1_, z2_] :=
  χfermR[t, α, β] * u1[-1, x] * su2fb[y] * su3fb[z1, z2];
χu[t_, α_, β_, x_, y_, z1_, z2_] := χfermL[t, α, β] * u1[-4, x] * su3fb[z1, z2];
χud[t_, α_, β_, x_, y_, z1_, z2_] := χfermR[t, α, β] * u1[4, x] * su3f[z1, z2];
χd[t_, α_, β_, x_, y_, z1_, z2_] := χfermL[t, α, β] * u1[2, x] * su3fb[z1, z2];
χdd[t_, α_, β_, x_, y_, z1_, z2_] := χfermR[t, α, β] * u1[-2, x] * su3f[z1, z2];
χL[t_, α_, β_, x_, y_, z1_, z2_] := χfermL[t, α, β] * u1[-3, x] * su2f[y];
χLd[t_, α_, β_, x_, y_, z1_, z2_] := χfermR[t, α, β] * u1[3, x] * su2fb[y];
χe[t_, α_, β_, x_, y_, z1_, z2_] := χfermL[t, α, β] * u1[6, x];
χed[t_, α_, β_, x_, y_, z1_, z2_] := χfermR[t, α, β] * u1[-6, x];
χBl[t_, α_, β_, x_, y_, z1_, z2_] := χfsL[t, α, β];
χBr[t_, α_, β_, x_, y_, z1_, z2_] := χfsR[t, α, β];
χWl[t_, α_, β_, x_, y_, z1_, z2_] := χfsL[t, α, β] * su2ad[y];
χWr[t_, α_, β_, x_, y_, z1_, z2_] := χfsR[t, α, β] * su2ad[y];
χGl[t_, α_, β_, x_, y_, z1_, z2_] := χfsL[t, α, β] * su3ad[z1, z2];
χGr[t_, α_, β_, x_, y_, z1_, z2_] := χfsR[t, α, β] * su3ad[z1, z2];

```

$$H(p, \phi_1, \dots, \phi_n) = \int d\mu_{\text{conformal}} d\mu_{\text{gauge}} \sum_{n=1}^{\infty} p^n \chi_{[n;0]}^* \prod_i PE[\phi_i \chi_i(q, \alpha, \beta)]$$

Hitoshi-no-MacBook-Pro.local 44: form hssm6.frm

Hitoshi-no-MacBook-Pro.local 49: form hssm8.frm

I

D=8 operators

$f =$
 $2*L^2*ld^2*t^2 + 2*ee*ed*L*ld*t^2 + ee^2*ed^2*t^2 + 2*d*dd*L*ld*t^2 + 2*d*dd*ee*ed*t^2 + 2*d^2*dd^2*t^2 + ud^2*dd*ed*t^2 + 2*u*ud*L*ld*t^2 + 2*u*ud*ee*ed*t^2 + 4*u*ud*d*dd*t^2 + u^2*d*ee*t^2 + 2*u^2*ud^2*t^2 + 2*Qd*dd*ee*L*t^2 + 3*Qd*ud*ed*L*ld*t^2 + 2*Qd*u*d*ld*t^2 + 3*Qd^2*ud*dd*t^2 + Qd^2*u*ee*t^2 + Qd^3*ld*t^2 + 2*Qd*d*ed*L*ld*t^2 + 2*Qd*ud*dd*L*t^2 + 3*Q*u*ee*L*t^2 + 4*Q*Qd*L*ld*t^2 + 2*Q*Qd*ee*ed*t^2 + 4*Q*Qd*d*dd*t^2 + 4*Q*Qd*u*ud*t^2 + Q^2*ud*ed*t^2 + 3*Q^2*u*d*t^2 + 4*Q^2*Qd^2*t^2 + Q^3*L*t^2 + Wr*L^2*ld^2 + Wr*ee*ed*L*ld + Wr*d*dd*L*ld + Wr*u*ud*L*ld + Wr*Qd*dd*ee*L + 3*Wr*Qd*ud*ed*L*ld + Wr*Qd*u*d*ld + 3*Wr*Qd^2*ud*dd + Wr*Qd^2*u*ee + 2*Wr*Qd^3*ld + Wr*Q*d*ed*L*ld + Wr*Q*ud*dd*L + 3*Wr*Q*Qd*L*ld + Wr*Q*Qd*ee*ed + 2*Wr*Q*Qd*d*dd + 2*Wr*Q*Qd*u*ud + 2*Wr*Q^2*Qd^2 + Wr^2*L*ld*t + Wr^2*Q*Qd*t + 2*Wr^4 + Wl*L^2*ld^2 + Wl*ee*ed*L*ld + Wl*d*dd*L*ld + Wl*u*ud*L*ld + Wl*Qd*dd*ee*L + Wl*Qd*u*d*ld + Wl*Q*d*ed*L*ld + Wl*Q*ud*dd*L + 3*Wl*Q*u*ee*L + 3*Wl*Q*Qd*L*ld + Wl*Q*Qd*ee*ed + 2*Wl*Q*Qd*d*dd + 2*Wl*Q*Qd*u*ud + Wl*Q^2*ud*ed + 3*Wl*Q^2*u*d + 2*Wl*Q^2*Qd^2 + 2*Wl*Q^3*L + 2*Wl*Wr*L*ld*t + Wl*Wr*ee*ed*t + Wl*Wr*d*dd*t + Wl*Wr*u*ud*t + 2*Wl*Wr*Q*Qd*t + Wl^2*L*ld*t + Wl^2*Q*Qd*t + 2*Wl^2*Wr^2 + 2*Wl^4 + Gr*d*dd*L*ld + Gr*d*dd*ee*ed + Gr*d^2*dd^2 + 3*Gr*ud^2*dd*ed + Gr*u*ud*L*ld + Gr*u*ud*ee*ed + 4*Gr*u*ud*d*dd + Gr*u^2*ud^2 + Gr*Qd*dd*ee*L + 3*Gr*Qd*ud*ed*L*ld + 2*Gr*Qd*u*d*ld + 6*Gr*Qd^2*ud*dd + Gr*Qd^2*u*ee + 2*Gr*Qd^3*ld + Gr*Q*d*ed*L*ld + 2*Gr*Q*ud*dd*L + 2*Gr*Q*Qd*L*ld + Gr*Q*Qd*ee*ed + 4*Gr*Q*Qd*d*dd + 4*Gr*Q*Qd*u*ud + Gr*Q^2*ud*ed + 2*Gr*Q^2*Qd^2 + Gr*Wr*Q*Qd*t + Gr*Wl*Q*Qd*t + Gr^2*d*dd*t + Gr^2*u*ud*t + Gr^2*Q*Qd*t + 2*Gr^2*Wr^2 + Gr^2*Wl^2 + 3*Gr^4 + Gl*d*dd*L*ld + Gl*d*dd*ee*ed + Gl*d^2*dd^2 + Gl*u*ud*L*ld + Gl*u*ud*ee*ed + 4*Gl*u*ud*d*dd + 3*Gl*u^2*d*ee + Gl*u^2*ud^2 + Gl*Qd*dd*ee*L + 2*Gl*Qd*u*d*ld + Gl*Qd^2*u*ee + Gl*Q*d*ed*L*ld + 2*Gl*Q*ud*dd*L + 3*Gl*Q*u*ee*L + 2*Gl*Q*Qd*L*ld + Gl*Q*Qd*ee*ed + 4*Gl*Q*Qd*d*dd + 4*Gl*Q*Qd*u*ud + Gl*Q^2*ud*ed + 6*Gl*Q^2*u*d + 2*Gl*Q^2*Qd^2 + 2*Gl*Q^3*L + Gl*Wr*Q*Qd*t + Gl*Wl*Q*Qd*t + Gl*Gr*L*ld*t + Gl*Gr*ee*ed*t + 3*Gl*Gr*d*dd*t + 3*Gl*Gr*u*ud*t + 3*Gl*Gr*Q*Qd*t + Gl*Gr*Wl*Wr + Gl^2*d*dd*t + Gl^2*u*ud*t + Gl^2*Q*Qd*t + Gl^2*Wr^2 + 2*Gl^2*Wl^2 + 3*Gl^2*Gr^2 + 3*Gl^4 + Br*ee*ed*L*ld + Br*d*dd*L*ld + Br*d*dd*ee*ed + 2*Br*ud^2*dd*ed + Br*u*ud*L*ld + Br*u*ud*ee*ed + 2*Br*u*ud*d*dd + Br*Qd*dd*ee*L + 3*Br*Qd*ud*ed*L*ld + Br*Qd*u*d*ld + 3*Br*Qd^2*ud*dd + Br*Qd^3*ld + Br*Q*d*ed*L*ld + Br*Q*ud*dd*L + 2*Br*Q*Qd*L*ld + Br*Q*Qd*ee*ed + 2*Br*Q*Qd*d*dd + 2*Br*Q*Qd*u*ud + Br*Q^2*ud*ed + Br*Wr*L*ld*t + Br*Wr*Q*Qd*t + Br*Wl*L*ld*t + Br*Wl*Q*Qd*t + Br*Gr*d*dd*t + Br*Gr*u*ud*t + Br*Gr*Q*Qd*t + Br*Gr^3 + Br*Gl*d*dd*t + Br*Gl*u*ud*t + Br*Gl*Q*Qd*t + Br*Gl^2*Gr + 2*Br^2*Wr^2 + Br^2*Wl^2 + 2*Br^2*Gr^2 + Br^2*G1^2 + Br^2*G1^2 + Br^4 + Bl*ee*ed*L*ld + Bl*d*dd*L*ld + Bl*d*dd*ee*ed + Bl*u*ud*L*ld + Bl*u*ud*ee*ed + 2*Bl*u*ud*d*dd + 2*Bl*u^2*d*ee + Bl*Qd*dd*ee*L + Bl*Qd*u*d*ld + Bl*Qd^2*u*ee + Bl*Q*d*ed*L*ld + Bl*Q*ud*dd*L + 3*Bl*Q*u*ee*L + 2*Bl*Q*Qd*L*ld + Bl*Q*Qd*ee*ed + 2*Bl*Q*Qd*d*dd + 2*Bl*Q*Qd*u*ud + 3*Bl*Q^2*u*d + Bl*Q^3*L + Bl*Wr*L*ld*t + Bl*Wr*Q*Qd*t + Bl*Wl*L*ld*t + Bl*Wl*Q*Qd*t + Bl*Gr*d*dd*t + Bl*Gr*u*ud*t + Bl*Gr*Q*Qd*t + Bl*Gl*d*dd*t + Bl*Gl*u*ud*t + Bl*Gl*Q*Qd*t + Bl*Gl*Gr^2 + Bl*Gl^3 + Bl*Br*L*ld*t + Bl*Br*ee*ed*t + Bl*Br*d*dd*t + Bl*Br*u*ud*t + Bl*Br*Q*Qd*t + Bl*Br*Wl*Wr + Bl*Br*G1*Gr + Bl^2*Wr^2 + 2*Bl^2*Wl^2 + Bl^2*Gr^2 + 2*Bl^2*G1^2 + Bl^2*Br^2 + Bl^4 + 3*Hd*ee*L^2*ld*t + Hd*ee^2*ed*L*t + 3*Hd*d*dd*ee*L*t + 3*Hd*ud*d*ed*L*ld*t + 2*Hd*ud^2*dd*L*t + 2*Hd*u*d^2*ld*t + 3*Hd*u*ud*ee*L*t + 6*Hd*Qd*ud*L*ld*t + 3*Hd*Qd*ud*ee*ed*t + 6*Hd*Qd*ud*d*dd*t + 3*Hd*Qd*u*d*ee*t + 3*Hd*Qd*u*ud^2*t + 3*Hd*Qd^2*d*ld*t + Hd*Qd^3*ee*t + 6*Hd*Q*d*L*ld*t + 3*Hd*Q*d*ee*ed*t + 3*Hd*Q*d^2*dd*t + 2*Hd*Q*d*ee*ed*t + 3*Hd*Q*d^2*dd*t + 2*Hd*Q*ud^2*ed*t + 6*Hd*Q*u*ud*d*t + 6*Hd*Q*Qd*ee*L*t + 6*Hd*Q*Qd^2*ud*t + 3*Hd*Q^2*ud*L*t + 6*Hd*Q^2*Qd*d*t + Hd*Wr*ee*L*t^2 + 2*Hd*Wr*Qd*ud*t^2 + Hd*Wr*Q*d*t^2 + Hd*Wr^2*ee*L + 2*Hd*Wr^2*Qd*ud + Hd*Wr^2*Q*d + 2*Hd*Wl*ee*L*t^2 + Hd*Wl*Qd*ud*t^2 + 2*Hd*Wl*Q*d*t^2 + 2*Hd*Wl^2*ee*L + Hd*Wl^2*Qd*ud + 2*Hd*Wl^2*Q*d + 2*Hd*Gr*Qd*ud*t^2 + Hd*Gr*Q*d*t^2 + 2*Hd*Gr*Wr*Qd*ud + Hd*Gr*Wr*Q*d + Hd*Gr^2*ee*L + 3*Hd*Gr^2*Qd*ud + 2*Hd*Gr^2*Q*d + Hd*G1*Qd*ud*t^2 + 2*Hd*G1*Q*d*t^2 + Hd*G1*Wl*Qd*ud + 2*Hd*G1*Wl*Q*d + Hd*G1^2*ee*L + 2*Hd*G1^2*Qd*ud + 3*Hd*G1^2*Q*d + Hd*Br*ee*L*t^2 + 2*Hd*Br*Qd*ud*t^2 + Hd*Br*Q*d*t^2 + Hd*Br*Wr*ee*L + 2*Hd*Br*$

$Wr*Qd*ud + Hd*Br*Wr*Q*d + 2*Hd*Br*Gr*Qd*ud + Hd*Br*Gr*Q*d + Hd*Br^2*ee*L + Hd*Br^2*Qd*ud + Hd*Br^2*Q*d + 2*Hd*Br^2*ee*L*t^2 + Hd*Br^2*Qd*ud*t^2 + 2*Hd*Br^2*Qd*d*t^2 + 2*Hd*Br^2*Wl*ee*L + Hd*Br^2*Wl*Qd*ud + 2*Hd*Br^2*Wl*Q*d + Hd*Br^2*G1*Qd*ud + 2*Hd*Br^2*G1*Q*d + Hd*Br^2*ee*L + Hd*Br^2*Qd*ud + Hd*Br^2*Q*d + Hd^2*ee^2*L^2 + Hd^2*ud*d*t^3 + Hd^2*ud*d*L*ld + Hd^2*Qd*ud*ee*L + 2*Hd^2*Qd^2*ud^2 + 2*Hd^2*Q*d*ee*L + 2*Hd^2*Q*Qd*ud*d + 2*Hd^2*Q^2*d^2 + Hd^2*Wr*ud*d*t + Hd^2*Wl*ud*d*t + Hd^2*Gr*ud*d*t + Hd^2*G1*ud*d*t + Hd^2*Br*ud*d*t + Hd^2*Bl*ud*d*t + 3*H*ed*L*ld^2*t + H*ee*ed^2*ld*t + 3*H*d*dd*ed*ld*t + 2*H*ud*dd^2*L*t + 3*H*u*dd*ee*L*t + 3*H*u*ud*ed*ld*t + 2*H*u^2*d*ld*t + 6*H*Qd*dd*L*ld*t + 3*H*Qd*dd*ee*ed*t + 3*H*Qd*d*dd^2*t + 6*H*Qd*u*ud*dd*t + 2*H*Qd*u^2*ee*t + 3*H*Qd^2*u*ld*t + 3*H*Q*ud*dd*ed*t + 6*H*Q*u*L*ld*t + 3*H*Q*u*ee*ed*t + 6*H*Q*u*d*dd*t + 3*H*Q*u^2*ud*t + 6*H*Q*Qd*ed*ld*t + 6*H*Q*Qd^2*dd*t + 3*H*Q^2*dd*L*t + 6*H*Q^2*Qd*u*t + H*Q^3*ed*t + 2*H*Wr*ed*ld*t^2 + 2*H*Wr*Qd*dd*t^2 + H*Wr*Q*u*t^2 + 2*H*Wr^2*ed*ld + 2*H*Wr^2*Qd*dd + H*Wr^2*Q*u + H*Wl*ed*ld*t^2 + H*Wl*Qd*dd*t^2 + 2*H*Wl*Q*u*t^2 + 2*H*Gr*Wr*Qd*dd + H*Gr*Wr*Q*ud + H*Gr^2*ed*ld + 3*H*Gr^2*Qd*dd + 2*H*Gr^2*Q*u + H*G1*Qd*dd*t^2 + 2*H*G1*Q*u*t^2 + H*G1*Wl*Qd*dd + 2*H*G1*Wl*Q*u + H*G1^2*ed*ld + 2*H*G1^2*Qd*dd + 3*H*G1^2*Q*u + 2*H*Br*ed*ld*t^2 + 2*H*Br*Qd*dd*t^2 + H*Br*Q*u*t^2 + 2*H*Br*Wr*ed*ld + 2*H*Br*Wr*Qd*dd + H*Br*Wr*Q*u + 2*H*Br*Gr*Qd*dd + H*Br*Gr*Q*u + H*Br^2*ed*ld + H*Br^2*Qd*dd + H*Br^2*Q*u + H*Bl*ed*ld*t^2 + H*Bl*Qd*dd*t^2 + 2*H*Bl*Q*u*t^2 + H*Bl*Wl*ed*ld + H*Bl*Wl*Qd*dd + 2*H*Bl*Wl*Q*u + H*Bl*G1*Qd*dd + 2*H*Bl*G1*Q*u + H*Bl^2*ed*ld + H*Bl^2*Qd*dd + H*Bl^2*Q*u + 4*H*Hd*L*ld*t^3 + 2*H*Hd*L^2*ld^2 + 2*H*Hd*ee*ed*t^3 + 2*H*Hd*ee*ed*L*ld + H*Hd*ee^2*ed^2 + 2*H*Hd*d*dd*t^3 + 2*H*Hd*d*dd*L*ld + H*Hd*d*dd*ee*ed + H*Hd*d^2*dd^2 + H*Hd*ud^2*dd*ed + 2*H*Hd*u*ud*t^3 + 2*H*Hd*u*ud*L*ld + H*Hd*u*ud*ee*ed + 2*H*Hd*u*ud*d*dd + H*Hd*u^2*d*ee + H*Hd*u^2*ud^2 + 2*H*Hd*Qd*dd*ee*L + 4*H*Hd*Qd*ud*ed*ld + 2*H*Hd*Qd*u*d*ld + 4*H*Hd*Qd^2*ud*dd + H*Hd*Qd^2*u*ee + 2*H*Hd*Qd^3*ld + 2*H*Hd*Q*d*ed*ld + 2*H*Hd*Q*ud*dd*L + 4*H*Hd*Q*u*ee*L + 4*H*Hd*Q*Qd*t^3 + 5*H*Hd*Q*Qd*L*ld + 2*H*Hd*Q*Qd*ee*ed + 4*H*Hd*Q*Qd*d*dd + 4*H*Hd*Q*Qd*u*ud + H*Hd*Q^2*ud*ed + 4*H*Hd*Q^2*u*d + 3*H*Hd*Q^2*Qd^2 + 2*H*Hd*Q^3*L + 6*H*Hd*Wr*L*ld*t + 2*H*Hd*Wr*ee*ed*t + 2*H*Hd*Wr*d*dd*t + 2*H*Hd*Wr*u*ud*t + 6*H*Hd*Wr*Q*Qd*t + 2*H*Hd*Wr^2*t^2 + H*Hd*Wr^3 + 6*H*Hd*Wl*L*ld*t + 2*H*Hd*Wl*ee*ed*t + 2*H*Hd*Wl*d*dd*t + 2*H*Hd*Wl*u*ud*t + 6*H*Hd*Wl*Q*Qd*t + 2*H*Hd*Wl*Wr*t^2 + 2*H*Hd*Wl^2*t^2 + H*Hd*Wl^3 + 2*H*Hd*Gr*d*dd*t + 2*H*Hd*Gr*u*ud*t + 4*H*Hd*Gr*Q*Qd*t + H*Hd*Gr^2*t^2 + H*Hd*Gr^3 + 2*H*Hd*G1*d*dd*t + 2*H*Hd*G1*u*ud*t + 4*H*Hd*G1*Q*Qd*t + H*Hd*G1*Gr*t^2 + H*Hd*G1^2*t^2 + H*Hd*G1^2*t^2 + 4*H*Hd*Br*L*ld*t + 2*H*Hd*Br*ee*ed*t + 2*H*Hd*Br*d*dd*t + 2*H*Hd*Br*u*ud*t + 4*H*Hd*Br*Q*Qd*t + 2*H*Hd*Br*Wr*t^2 + H*Hd*Br*Wr^2 + H*Hd*Br*Wl*t^2 + H*Hd*Br^2*t^2 + 4*H*Hd*Bl*L*ld*t + 2*H*Hd*Bl*ee*ed*t + 2*H*Hd*Bl*d*dd*t + 2*H*Hd*Bl*u*ud*t + 4*H*Hd*Bl*Q*Qd*t + H*Hd*Bl*Wr*t^2 + 2*H*Hd*Bl*Wl*t^2 + H*Hd*Bl*Wl^2 + H*Hd*Bl*Br*t^2 + H*Hd*Bl^2*t^2 + 6*H*Hd^2*ee*L*t^2 + 6*H*Hd^2*Qd*ud*t^2 + 6*H*Hd^2*Q*d*t^2 + 2*H*Hd^2*Wr*Qd*ud + 2*H*Hd^2*Wl*ee*L + 2*H*Hd^2*Wl*Q*d + H*Hd^2*Gr*Qd*ud + H*Hd^2*G1*Q*d + H*Hd^2*Br*Qd*ud + H*Hd^2*Bl*ee*L + H*Hd^2*Bl*Q*d + H*Hd^3*ud*d*t + H^2*ed^2*ld^2 + H^2*u*dd*t^3 + H^2*u*dd*L*ld + 2*H^2*Qd*dd*ed*ld + 2*H^2*Qd^2*dd^2 + H^2*Q*u*ed*ld + 2*H^2*Q*Qd*u*dd + 2*H^2*Q^2*u^2 + H^2*Wr*u*dd*t + H^2*Wl*u*dd*t + H^2*Gr*u*dd*t + H^2*G1*u*dd*t + H^2*Br*u*dd*t + H^2*Bl*u*dd*t + 6*H^2*Hd*ed*ld*t^2 + 6*H^2*Hd*Qd*dd*t^2 + 6*H^2*Hd*Q*u*t^2 + 2*H^2*Hd*Wr*ed*ld + 2*H^2*Hd*Wr*Qd*dd + 2*H^2*Hd*Wl*Q*u + H^2*Hd*Gr*Qd*dd + H^2*Hd*G1*Q*u + H^2*Hd*Br*ed*ld + H^2*Hd*Br*Qd*dd + H^2*Hd*Bl*Q*u + 3*H^2*Hd^2*t^4 + 4*H^2*Hd^2*L*ld*t + H^2*Hd^2*ee*ed*t + H^2*Hd^2*d*dd*t + H^2*Hd^2*u*ud*t + 4*H^2*Hd^2*Q*Qd*t + 2*H^2*Hd^2*Wr*t^2 + 2*H^2*Hd^2*Wr^2 + 2*H^2*Hd^2*Wl*t^2 + 2*H^2*Hd^2*Wl^2 + H^2*Hd^2*Gr^2 + H^2*Hd^2*G1^2 + H^2*Hd^2*Br*t^2 + H^2*Hd^2*Br*Wr + H^2*Hd^2*Br^2 + H^2*Hd^2*Bl*t^2 + H^2*Hd^2*Bl*Wl + H^2*Hd^2*Bl^2 + H^2*Hd^3*ee*L + H^2*Hd^3*Qd*ud + H^2*Hd^3*Q*d + H^3*Hd*u*dd*t + H^3*Hd^2*ed*L*ld + H^3*Hd^2*Qd*dd + H^3*Hd^2*Q*u + 2*H^3*Hd^3*t^2 + H^4*Hd^4;$

993 of them for $N_f=1$

Conclusions

- Nailed the question of classifying effective operators in any given Lorentz-inv theory
- Also for chiral Lagrangians
- useful techniques for matching
- careful mapping to observables
- hope for deviations from Standard Model
- inverse problem to identify BSM physics