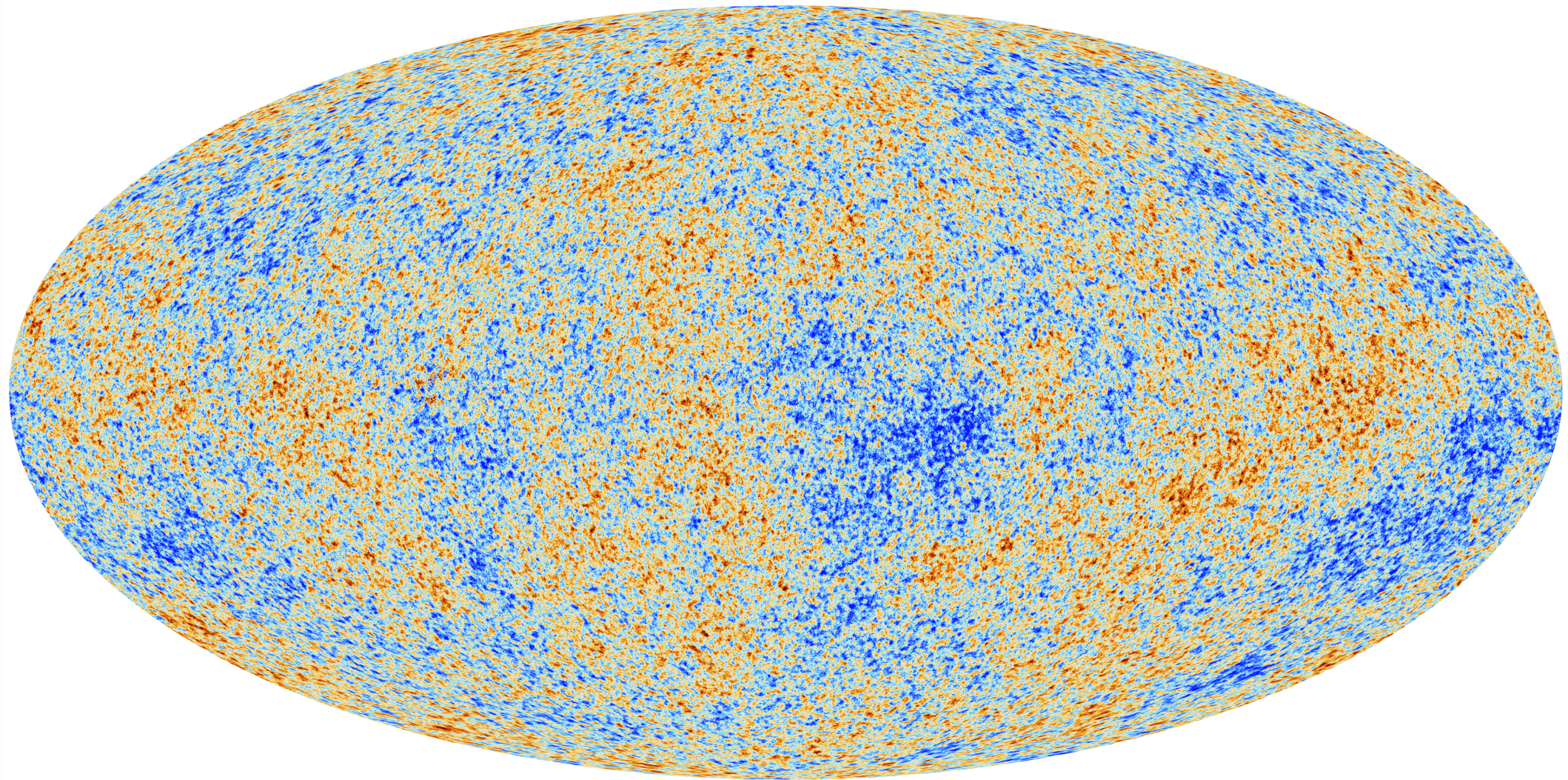


# An almost perfect Universe

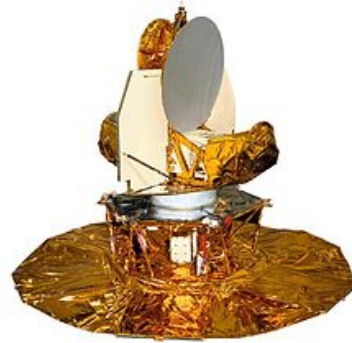
results from the Planck mission

Torsten Enßlin (MPI für Astrophysik)

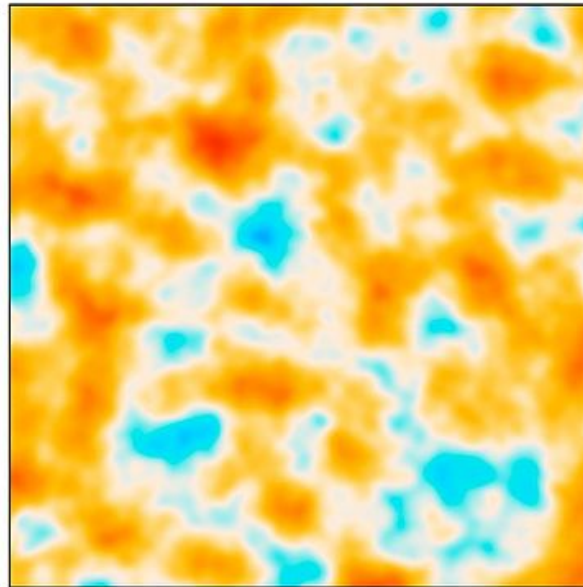
on behalf of the Planck Collaboration



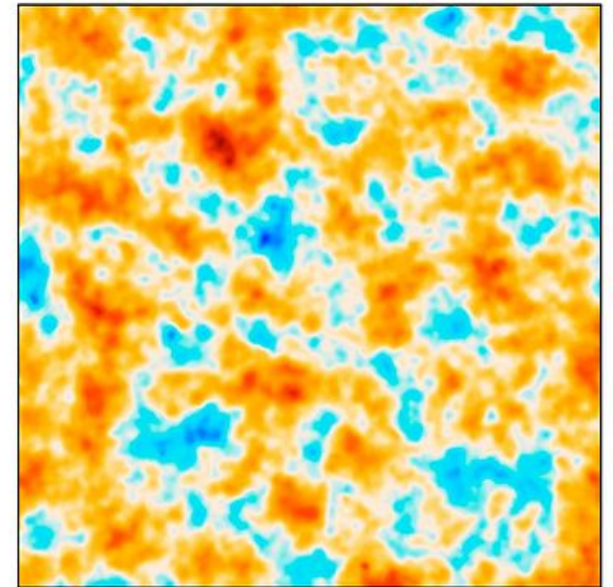
# Why Planck?



COBE



WMAP



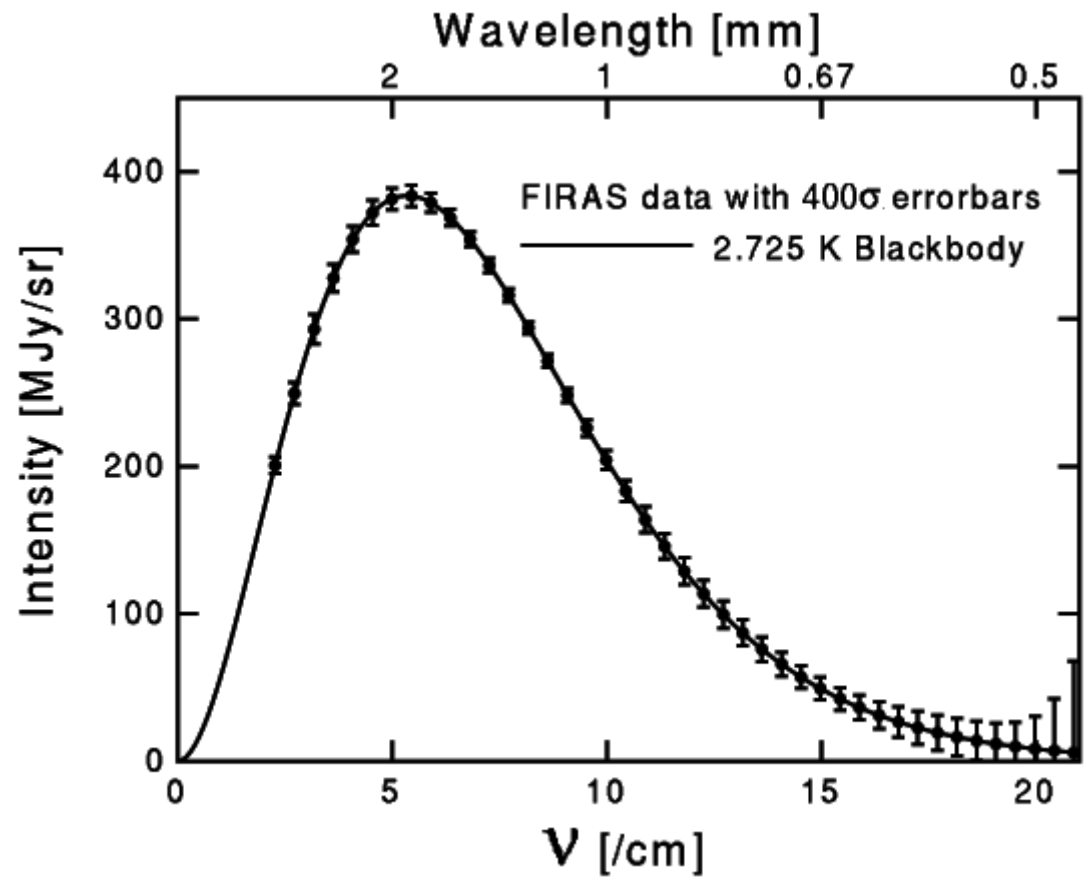
Planck

# Why Planck?

$$M_{\text{pl}} = \sqrt{\frac{\hbar c}{8\pi G}}$$




$$I(\nu) = \frac{2 h \nu^3}{c^2} \frac{1}{e^{\frac{h \nu}{k_B T}} - 1}$$





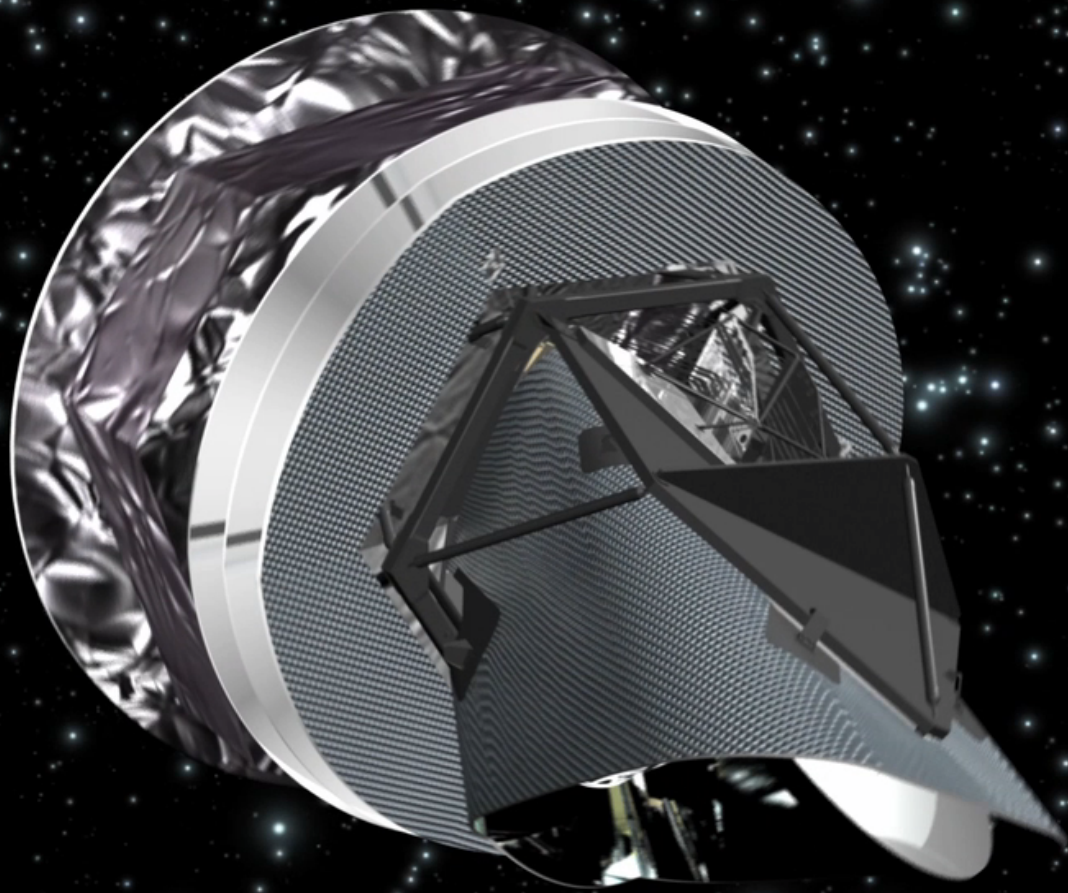
# Gaussian statistics



The background is a complex collage. On the left, there's a stylized illustration of a large, multi-story building with a central dome and spire. In the top center, there's a hexagonal logo with a stylized 'S' or 'B' inside. On the right, a large, textured, purple-toned face of a person is visible. In the center, there's a graph of a normal distribution curve. The curve is black, and the area under it is shaded in a light purple. The x-axis is labeled with  $\mu$  and  $\sigma$ . The y-axis is labeled with  $f(x)$ . The formula for the normal distribution is written in the center:  $f(x) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$ . Below the formula, the mean  $\mu$  is given as 2.725 Kelvin, and the standard deviation  $\sigma$  is given as 0.00001 Kelvin.
$$f(x) = \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$
$$\mu = 2.725 \text{ Kelvin}$$
$$\sigma = 0.00001 \text{ Kelvin}$$

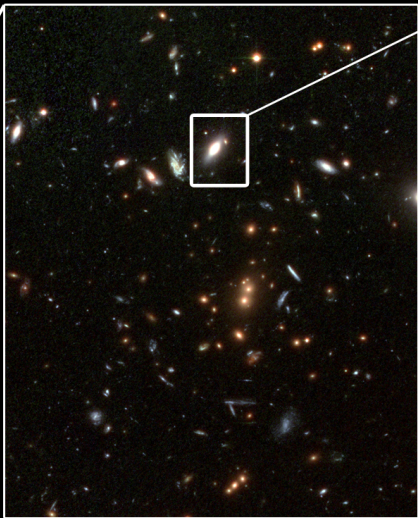
# Testing inflation







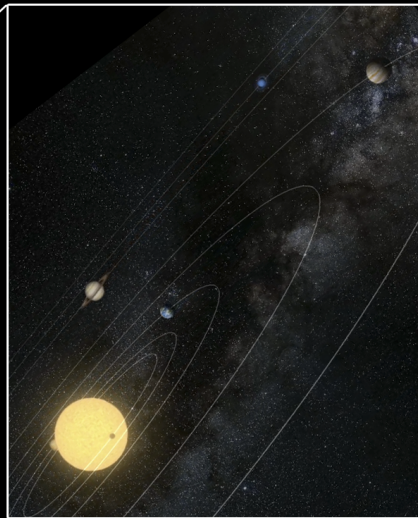
Clusters of galaxies

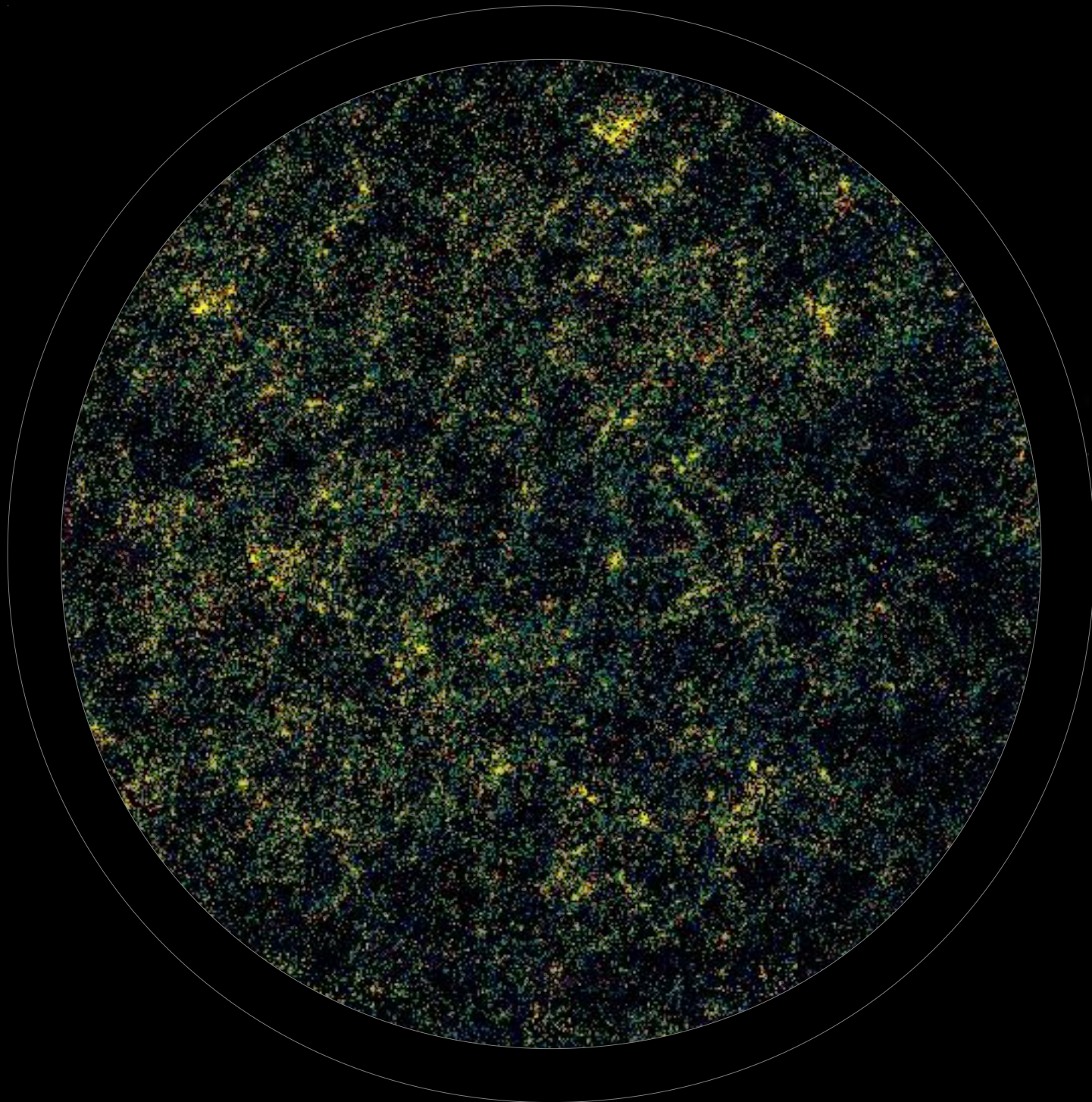


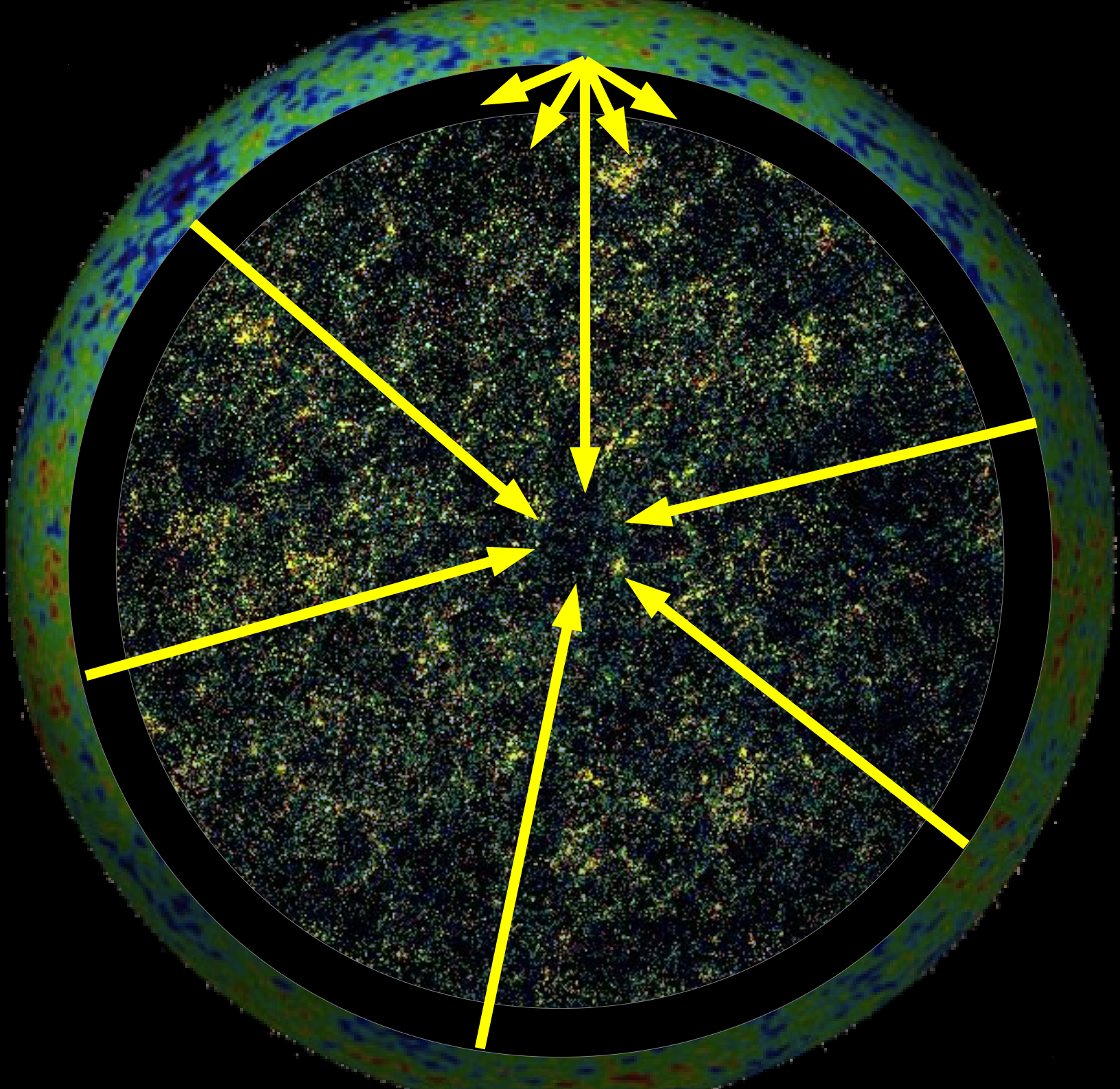
Galaxy



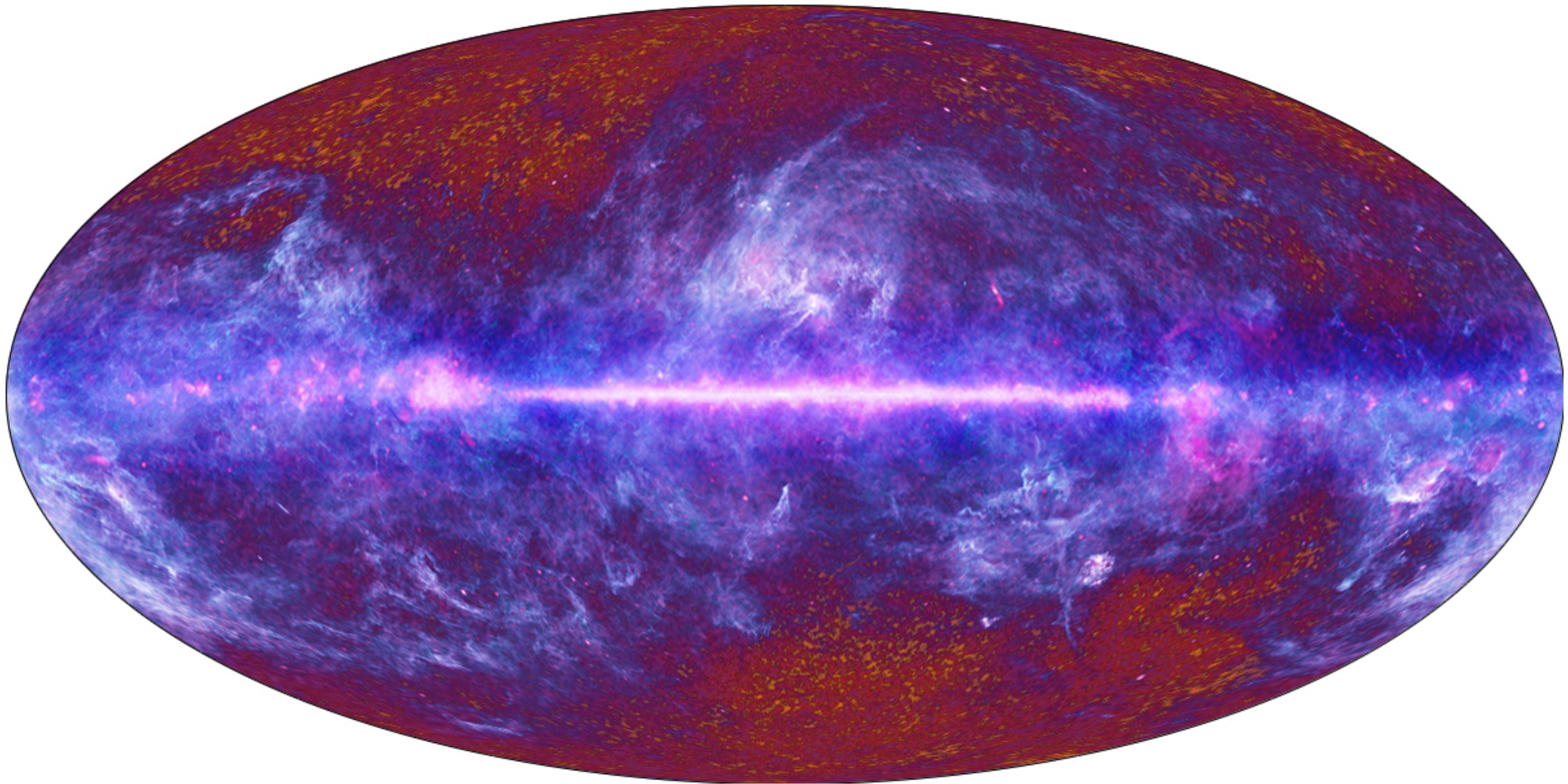
Solar System



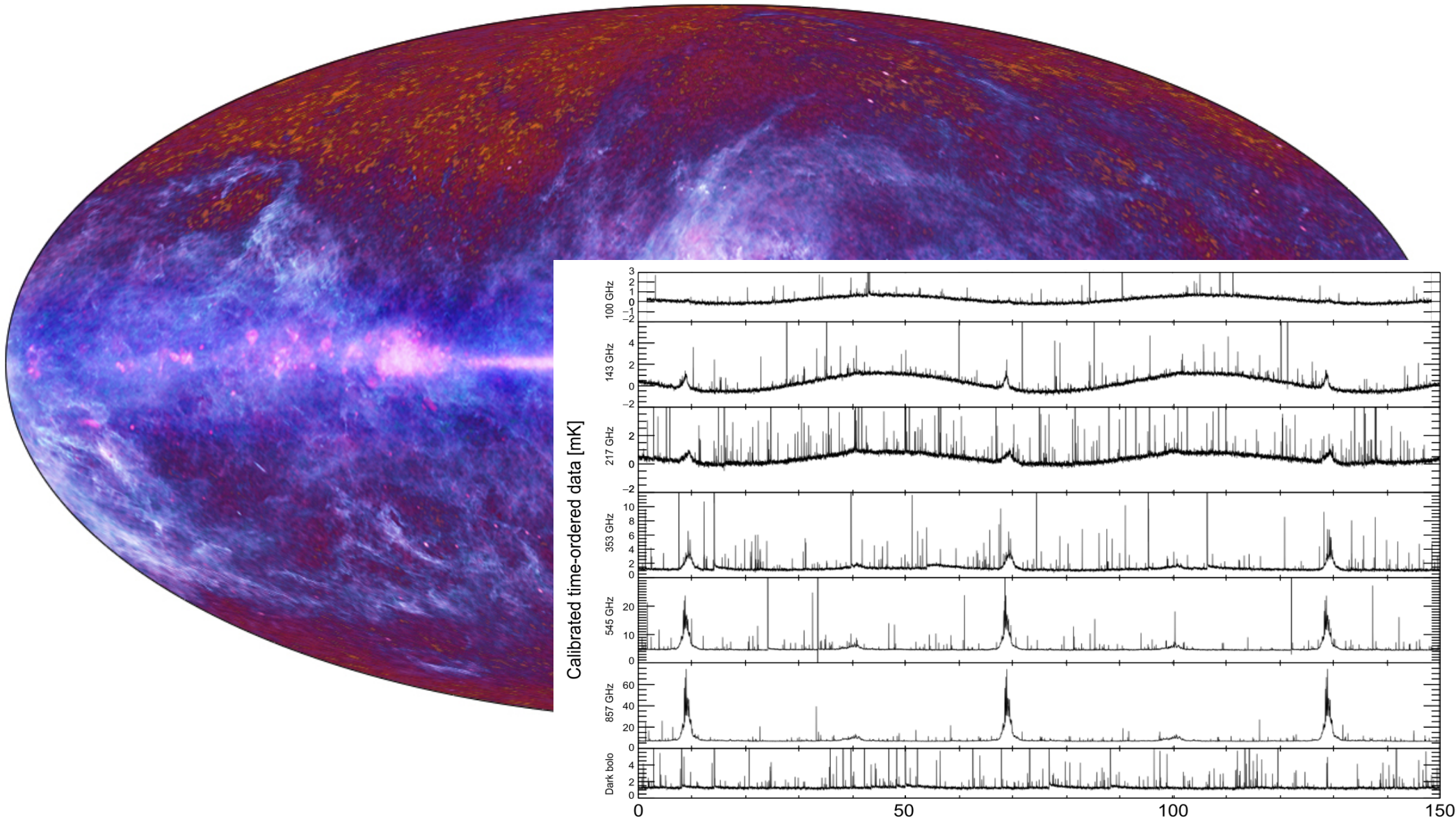




# The microwave sky



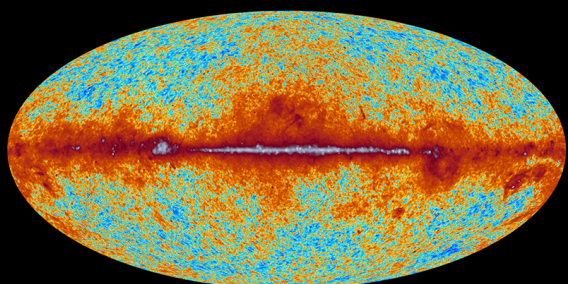
# The timelines



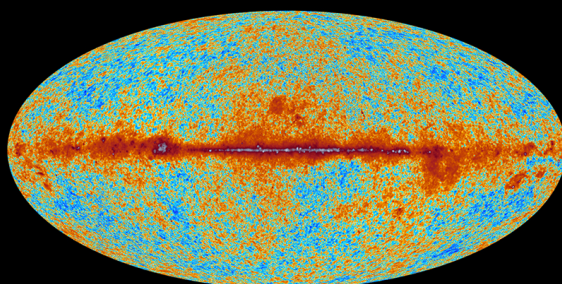


planck

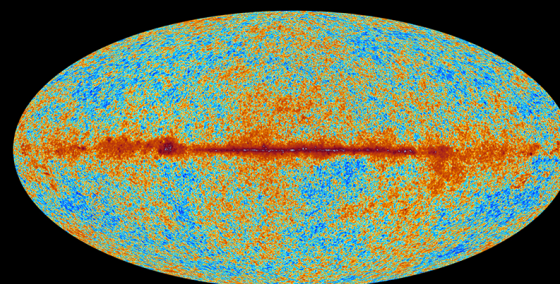
# The sky as seen by Planck



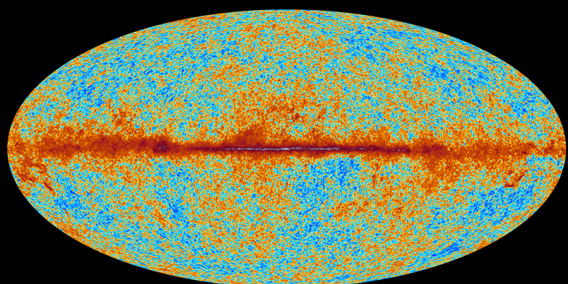
30 GHz



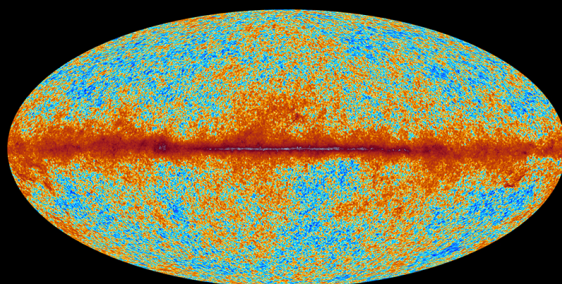
44 GHz



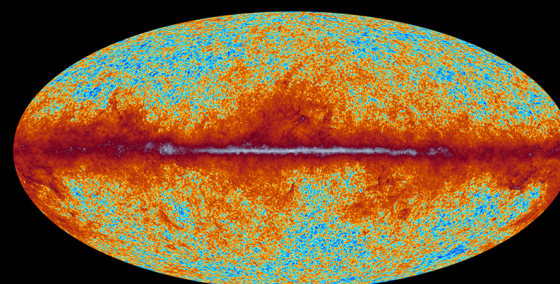
70 GHz



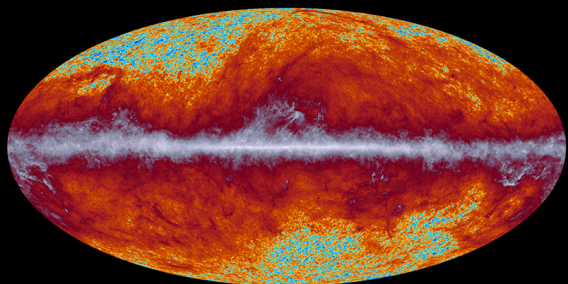
100 GHz



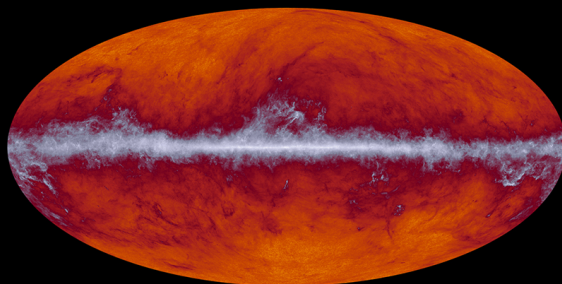
143 GHz



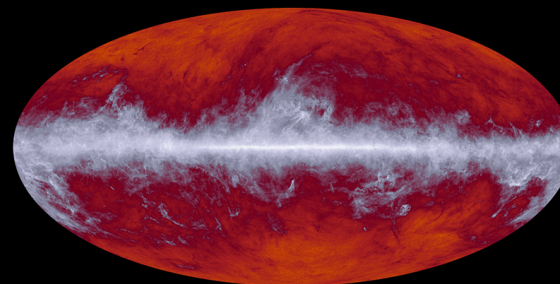
217 GHz



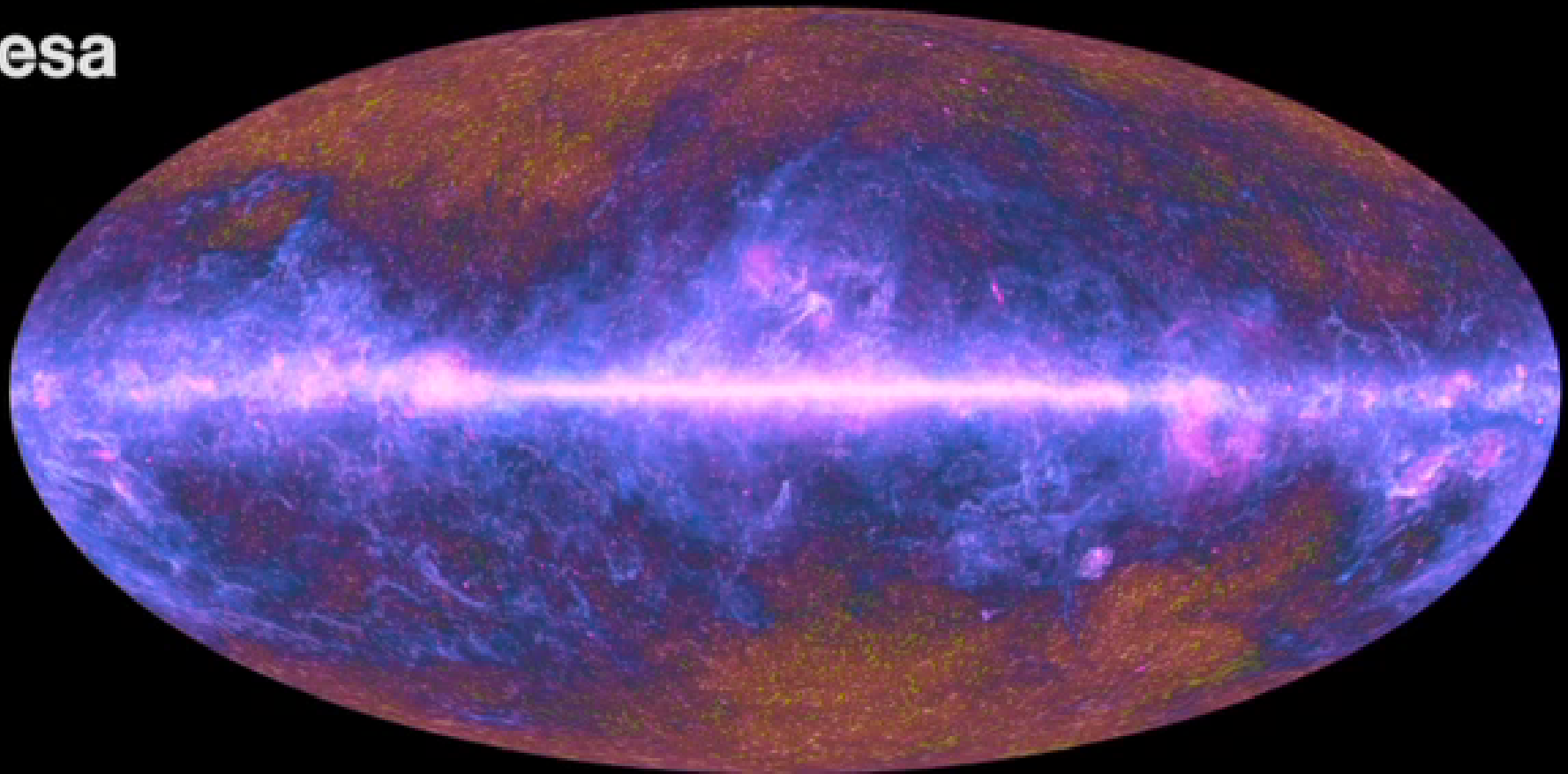
353 GHz



545 GHz



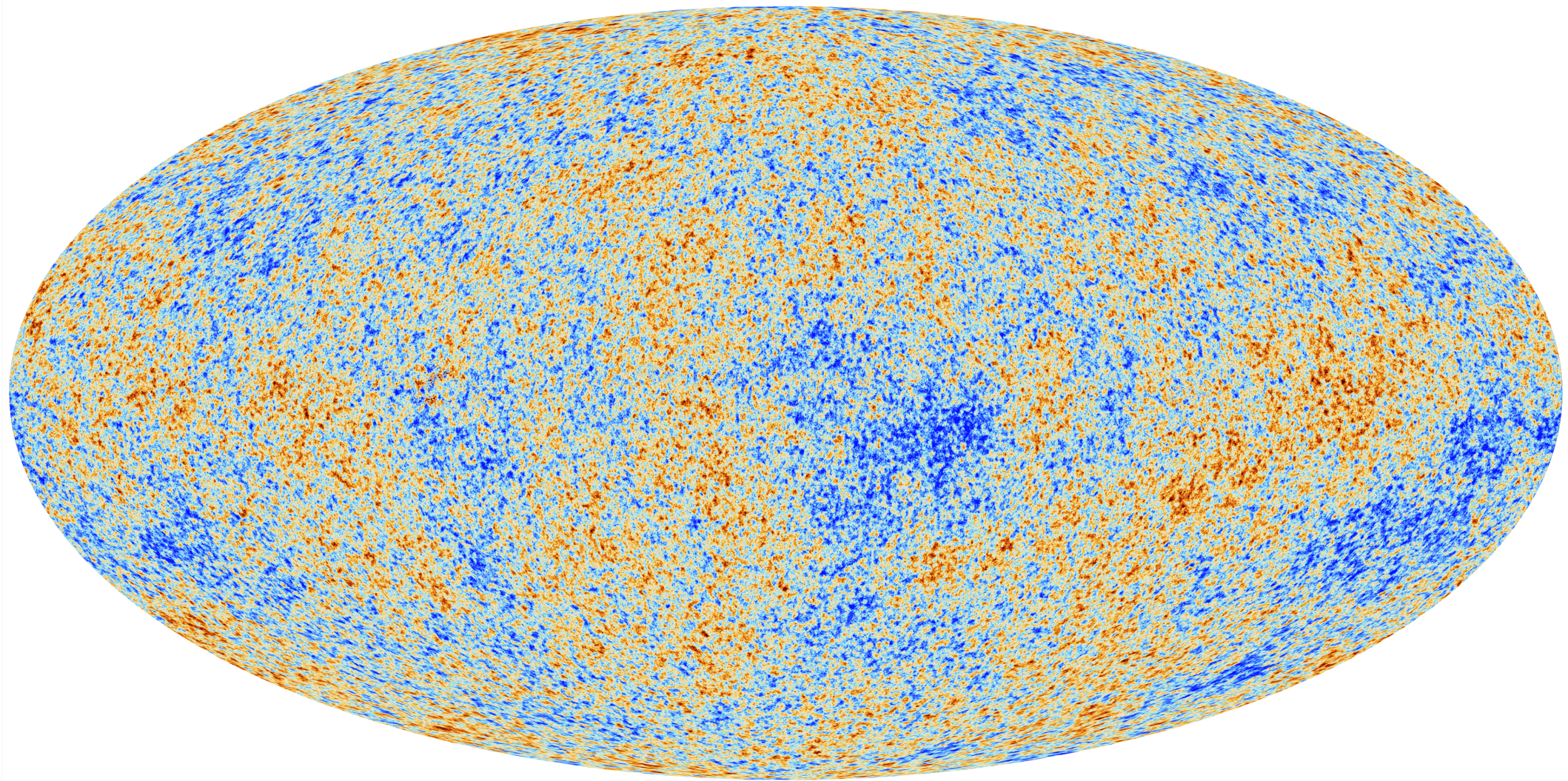
857 GHz



Individual sources + Radio emission from the Milky Way + Dust emission from the Milky Way + Cosmic Microwave Background

All emissions at microwave & submillimetre wavelengths

# The cosmic microwave background

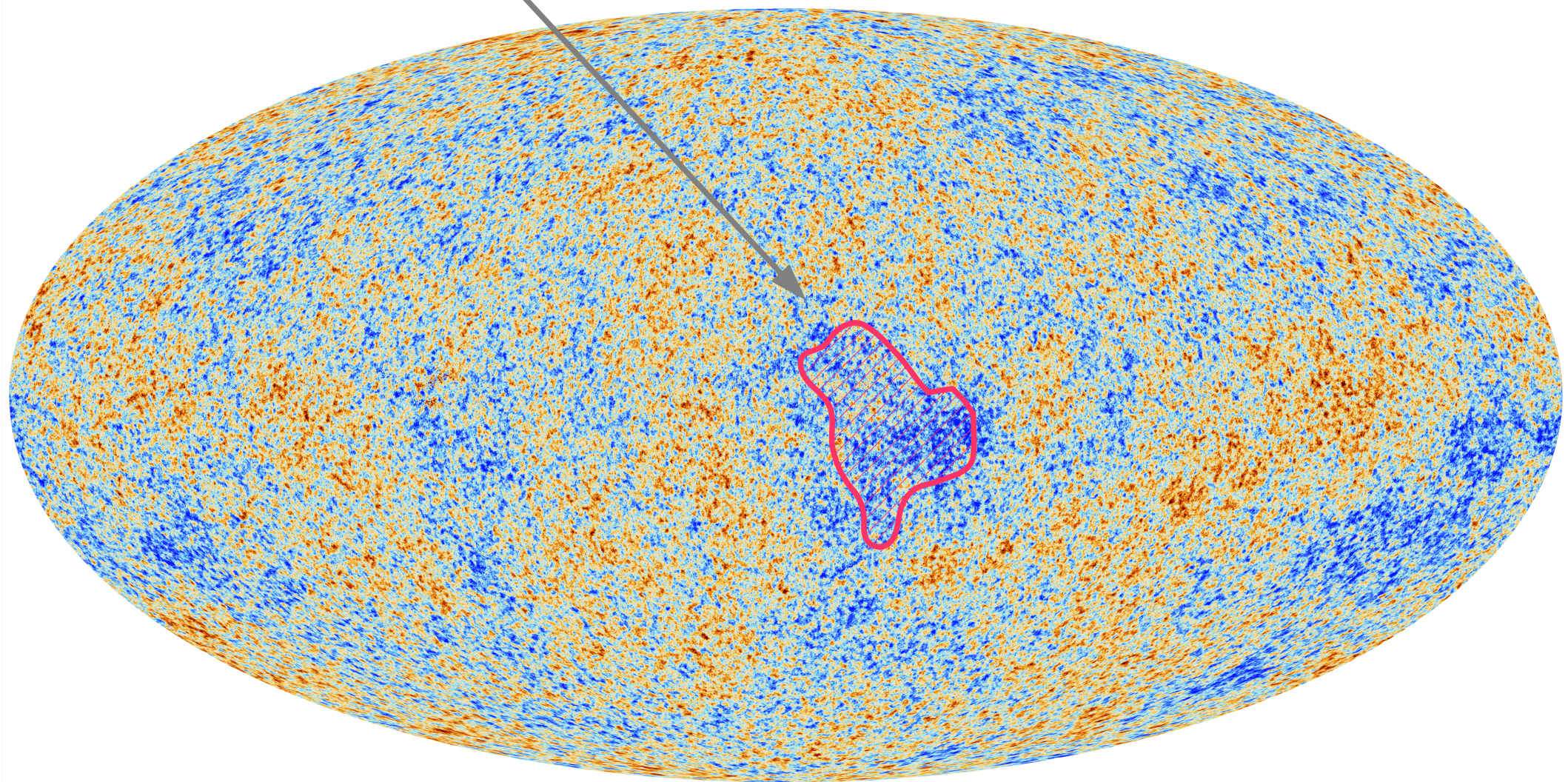


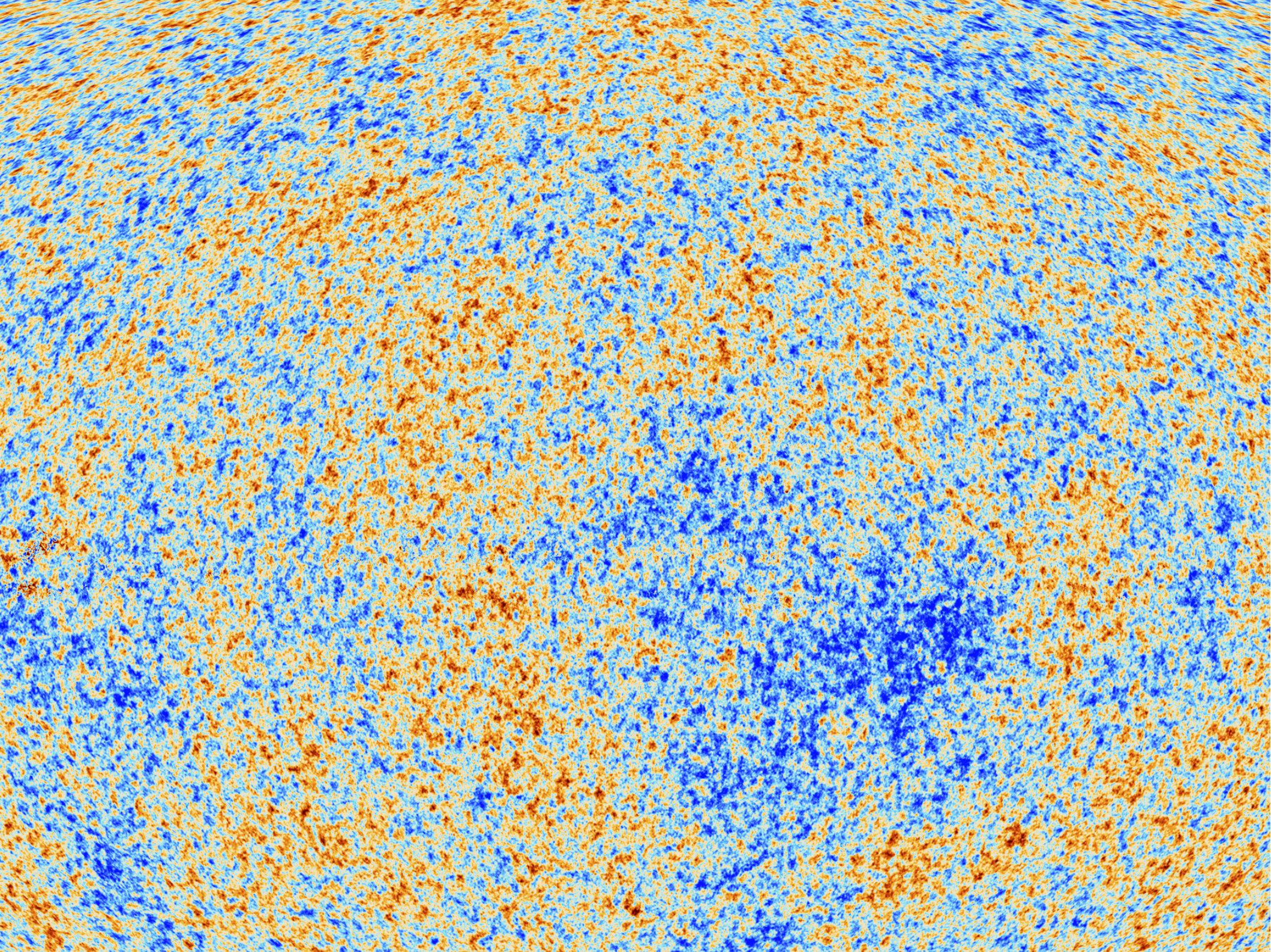


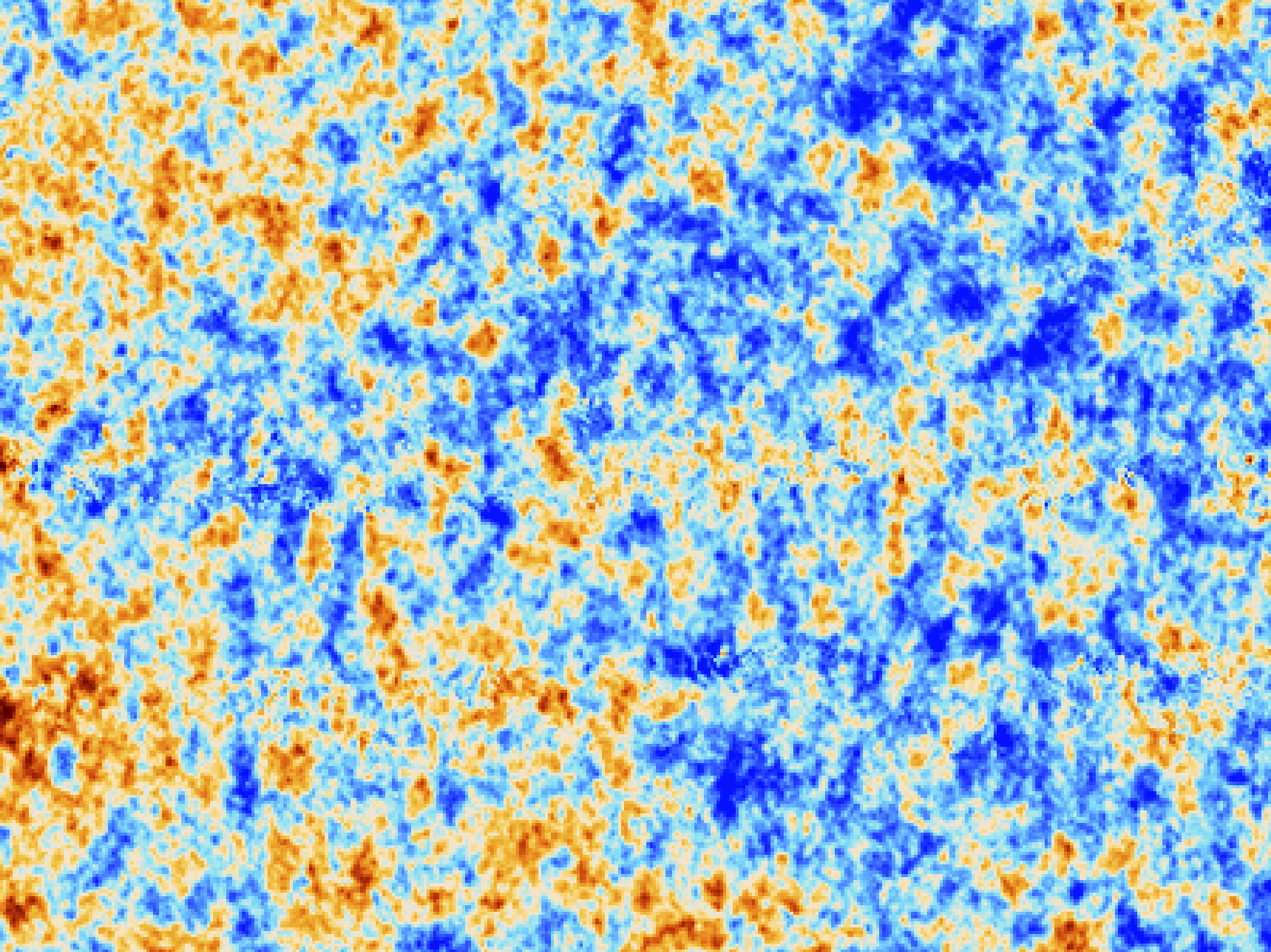
Effect of gravity on  
Photons

$\phi_{grav}$

$$\frac{\delta T}{T} = \frac{1}{3} \phi_{grav}$$

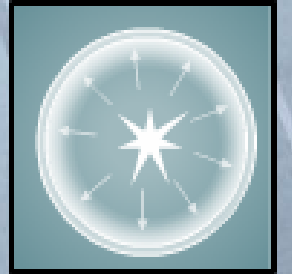
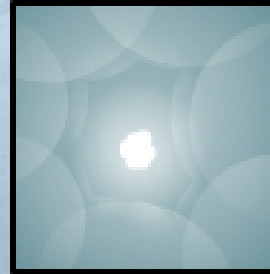
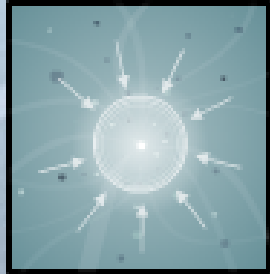
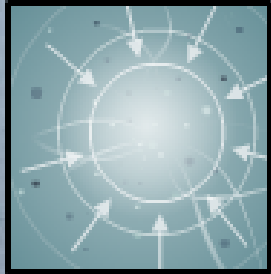
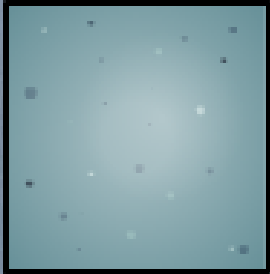






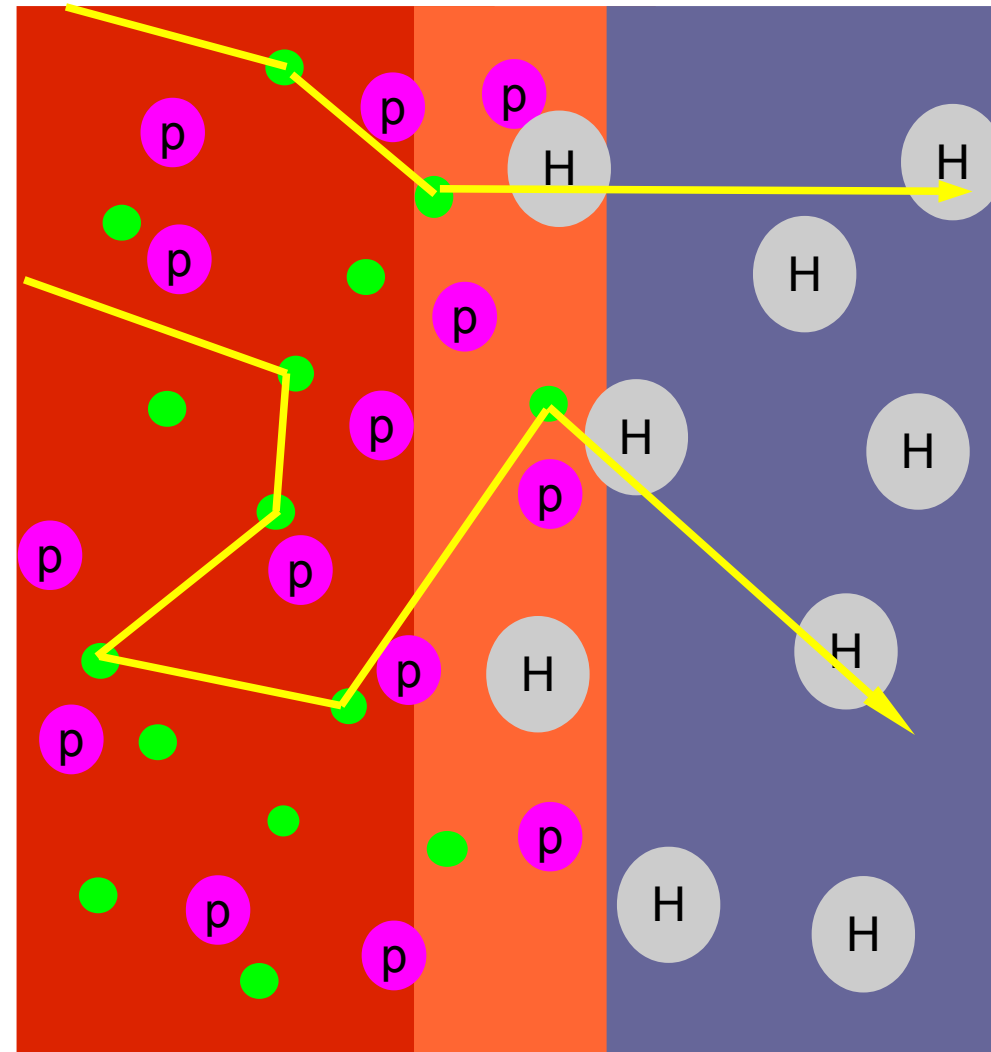


ca. 1 angular degree

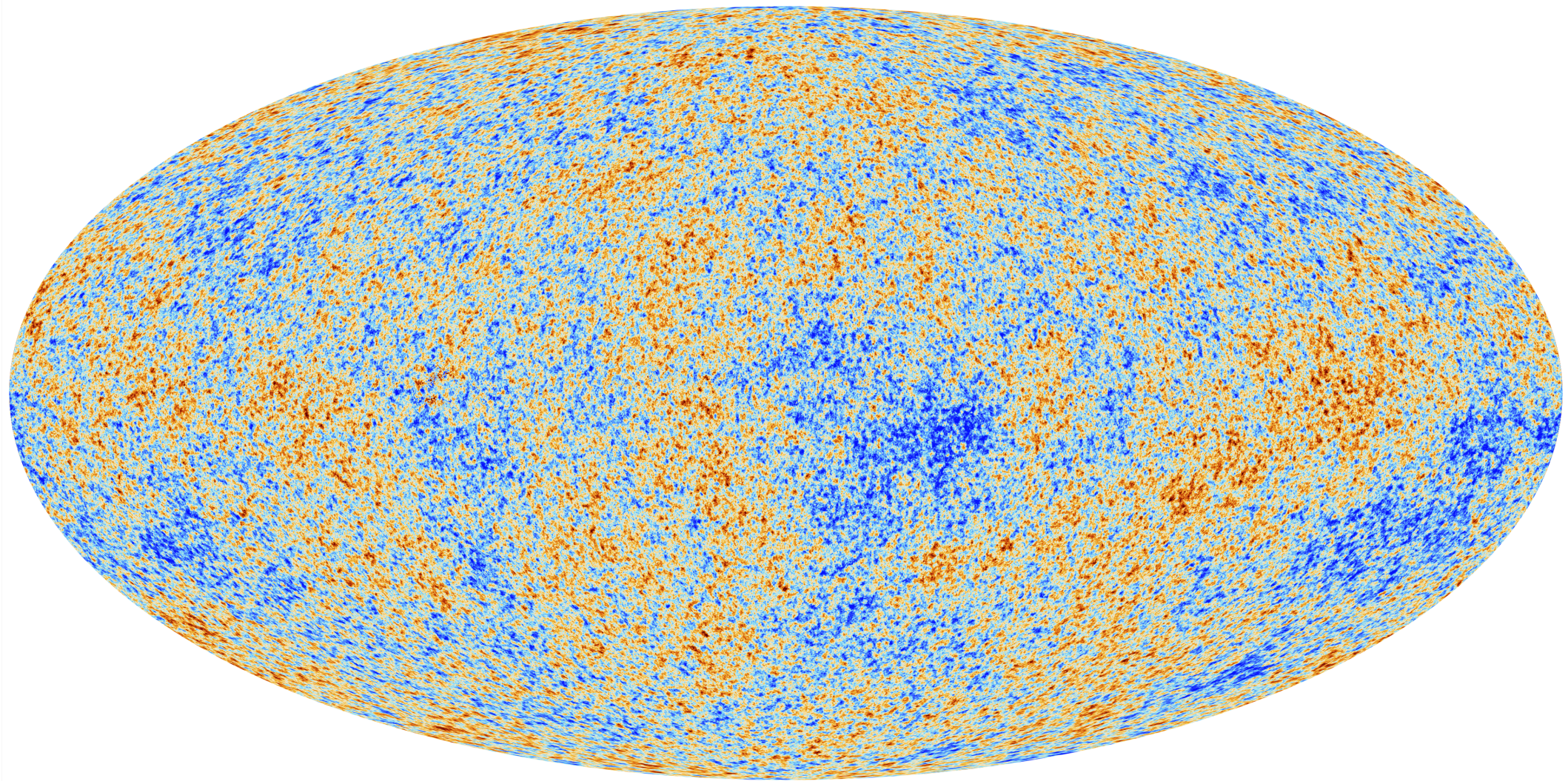


# Recombination & photon diffusion

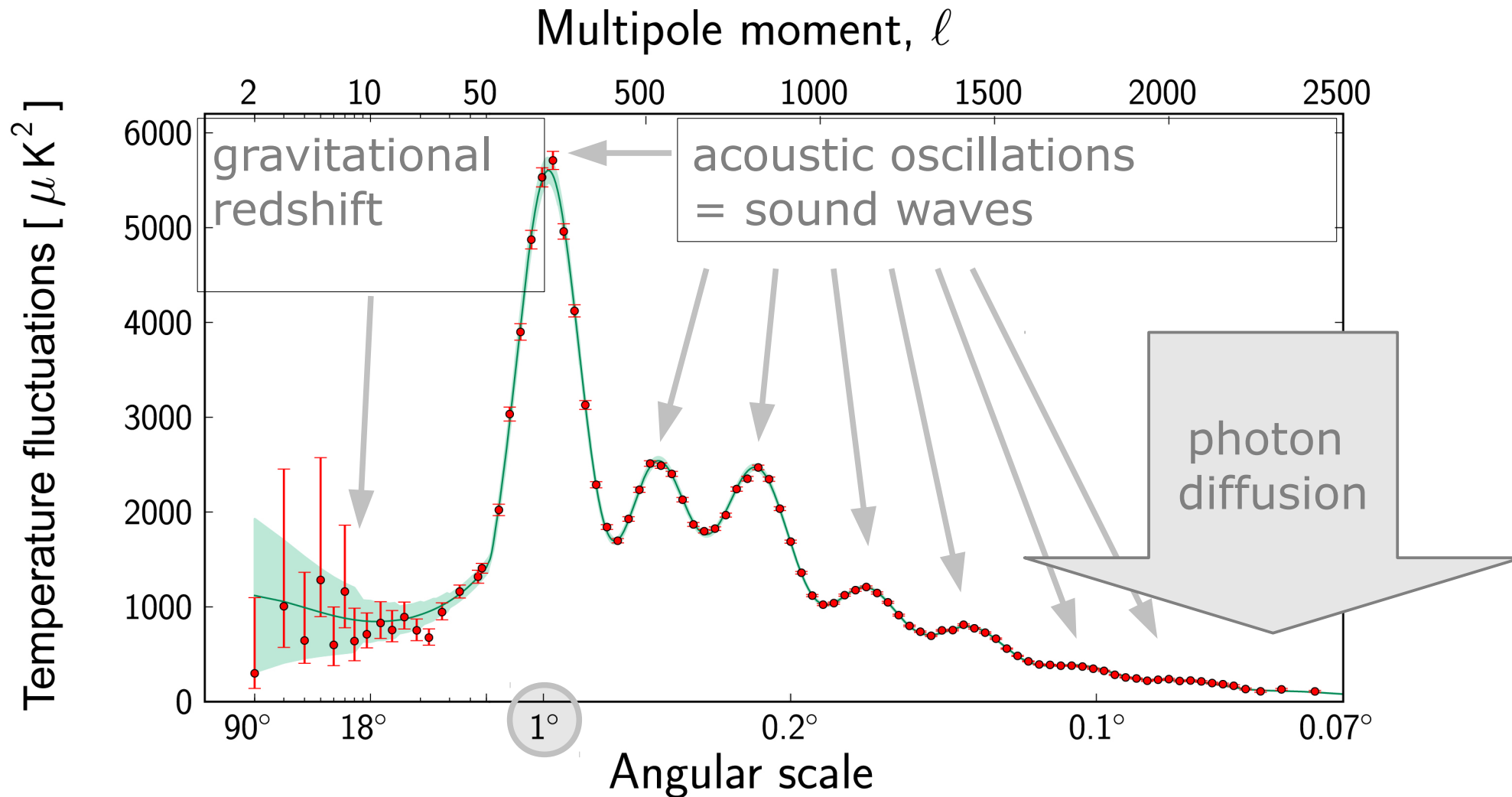
- Before recombination: photons scatter on free electrons
- After: photons free-stream through Universe
- Observed CMB light from surface of last scattering
- Finite duration of recombination permitted photons to diffuse
- Structures smaller than 5 arcmin are erased (1/6 of moon diameter)
- Planck resolution: 5 arcmin

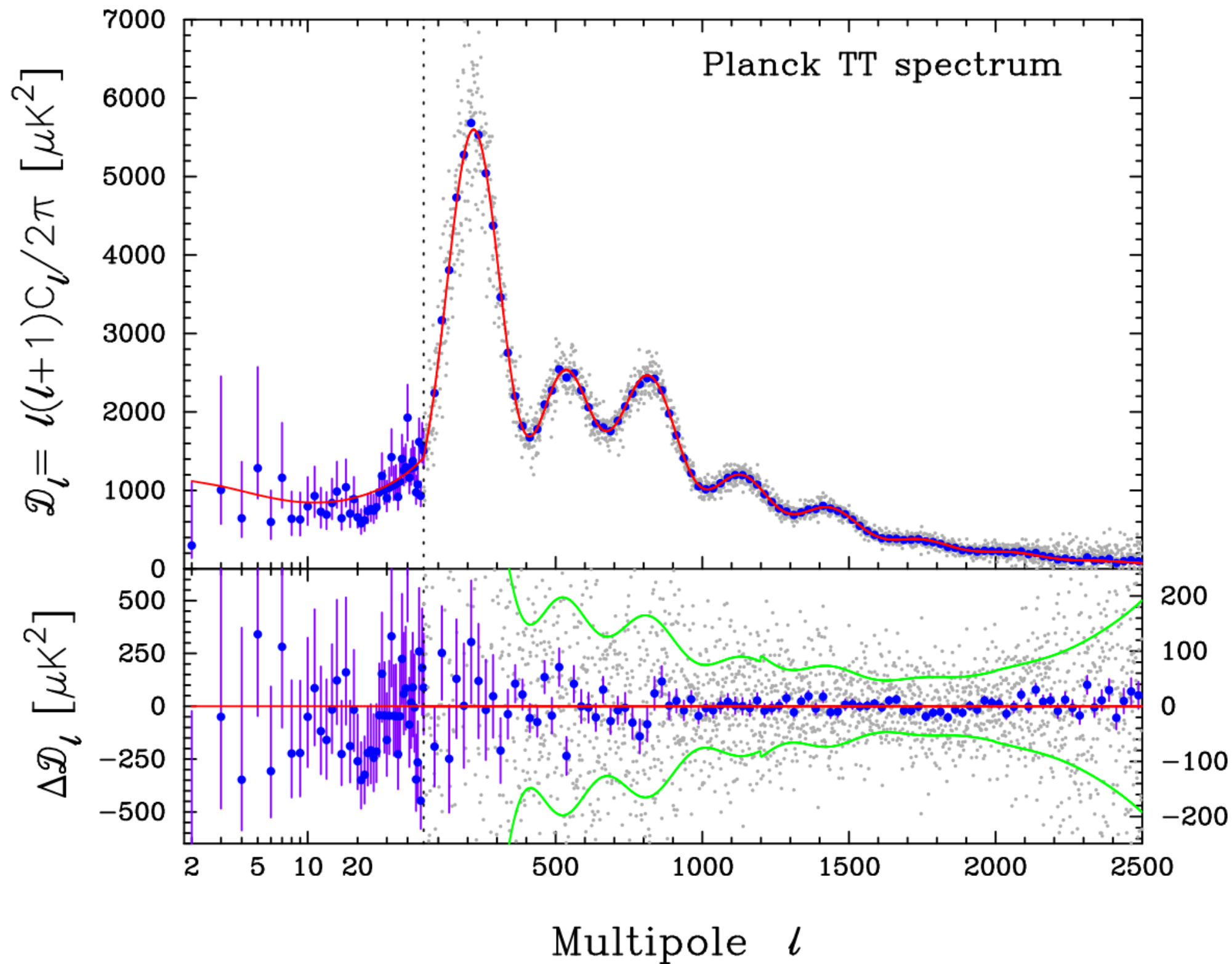


# The cosmic microwave background



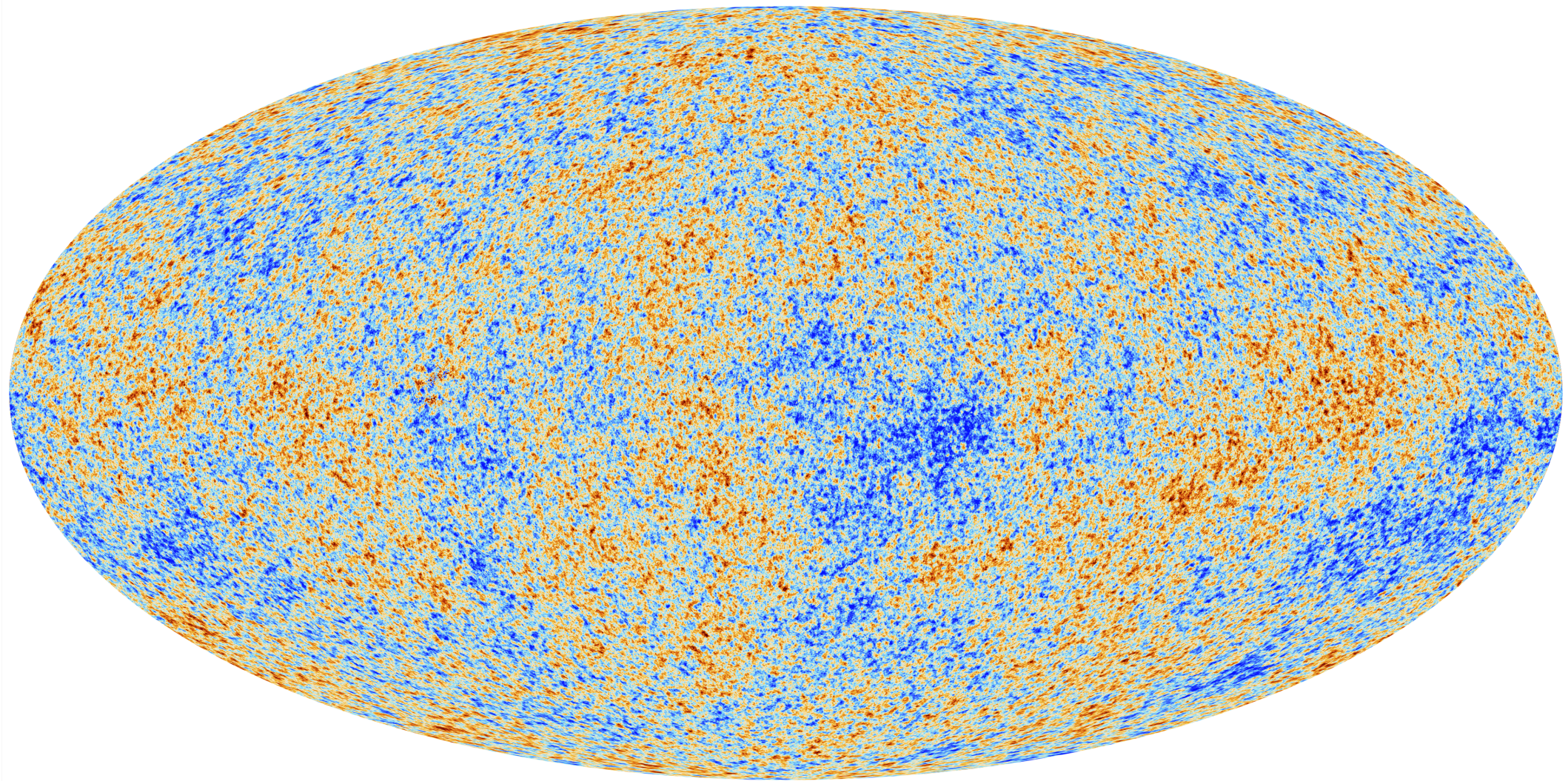
# Angular power spectrum



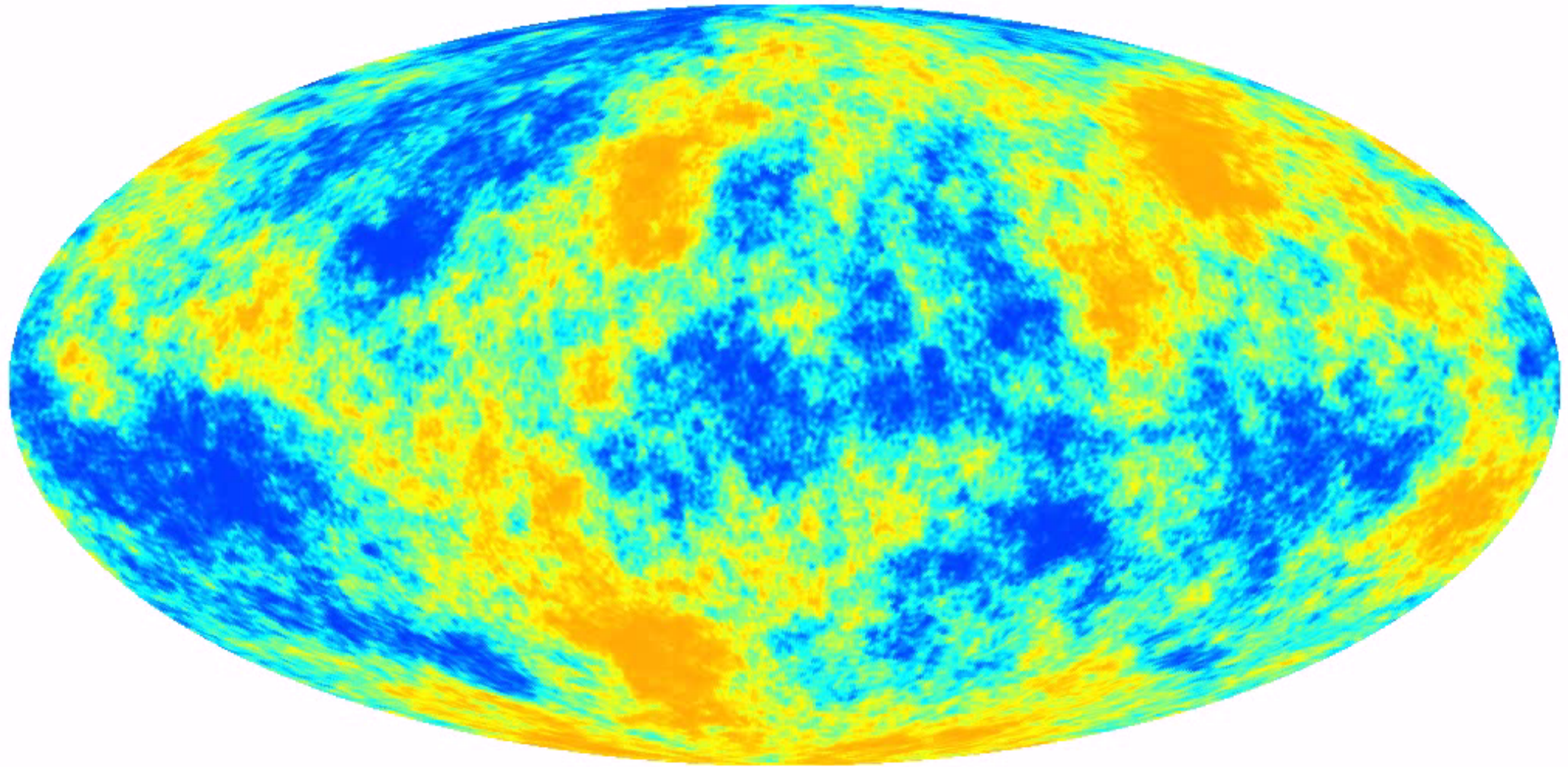




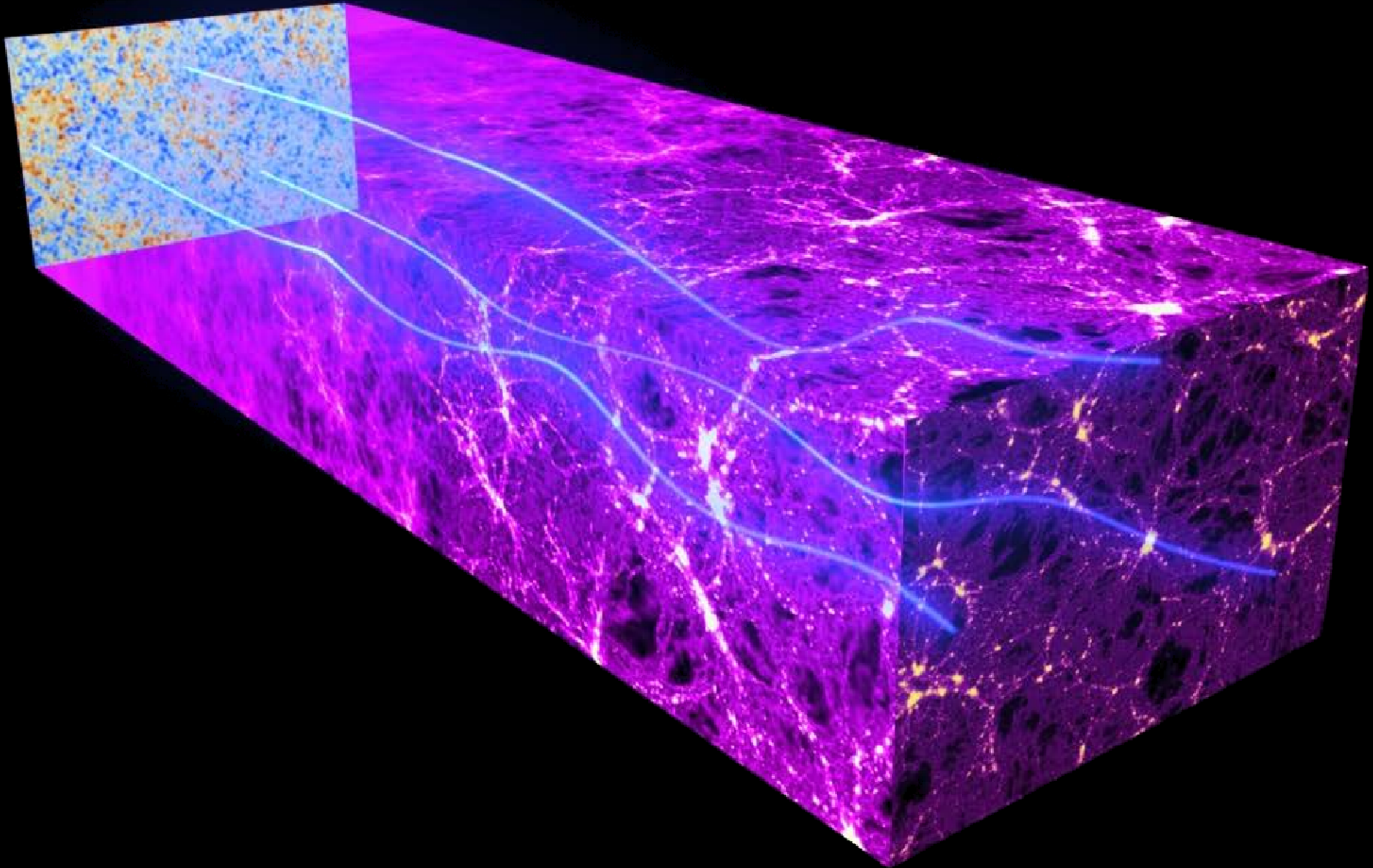
# The cosmic microwave background



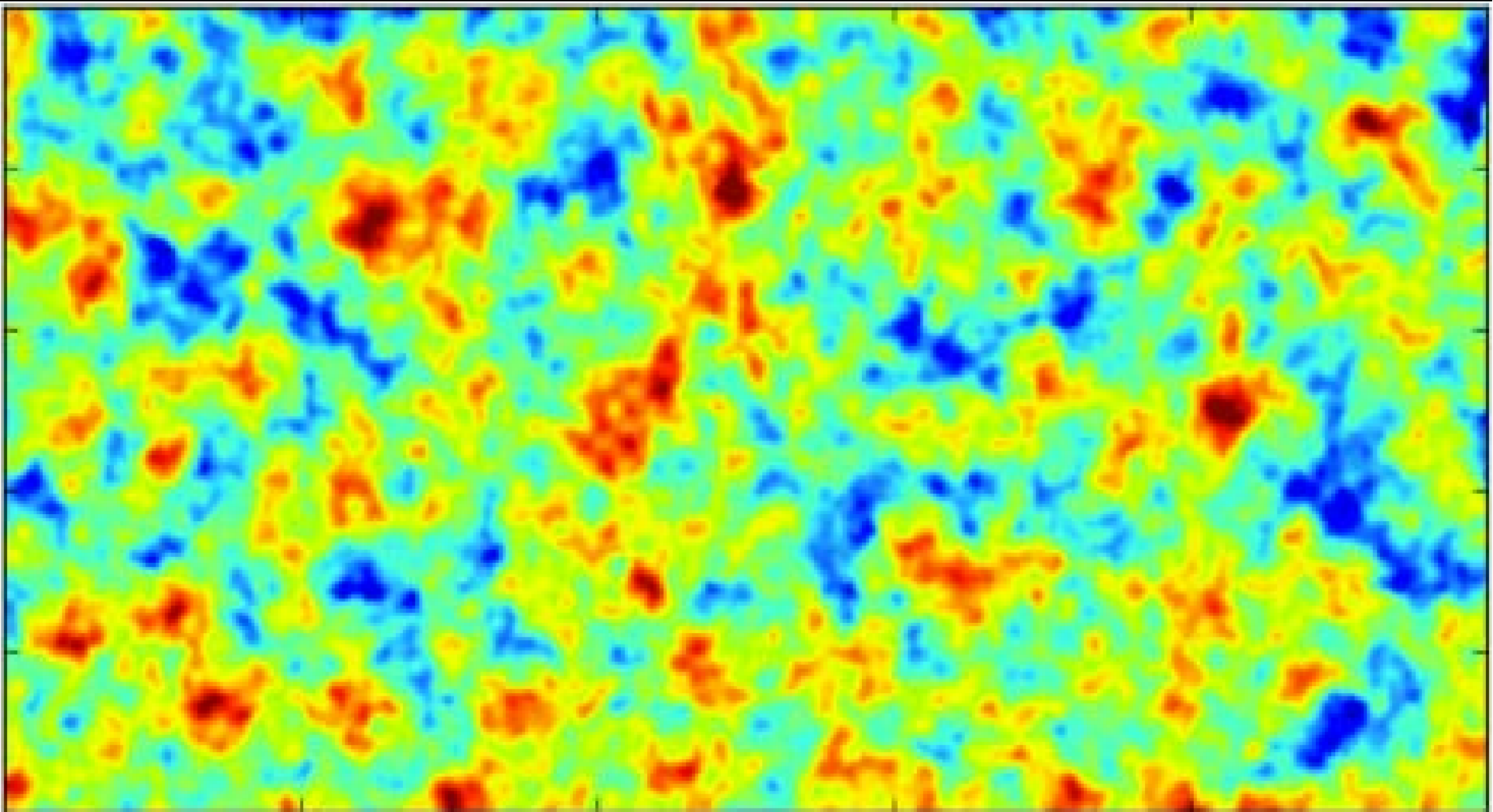
# Simulation



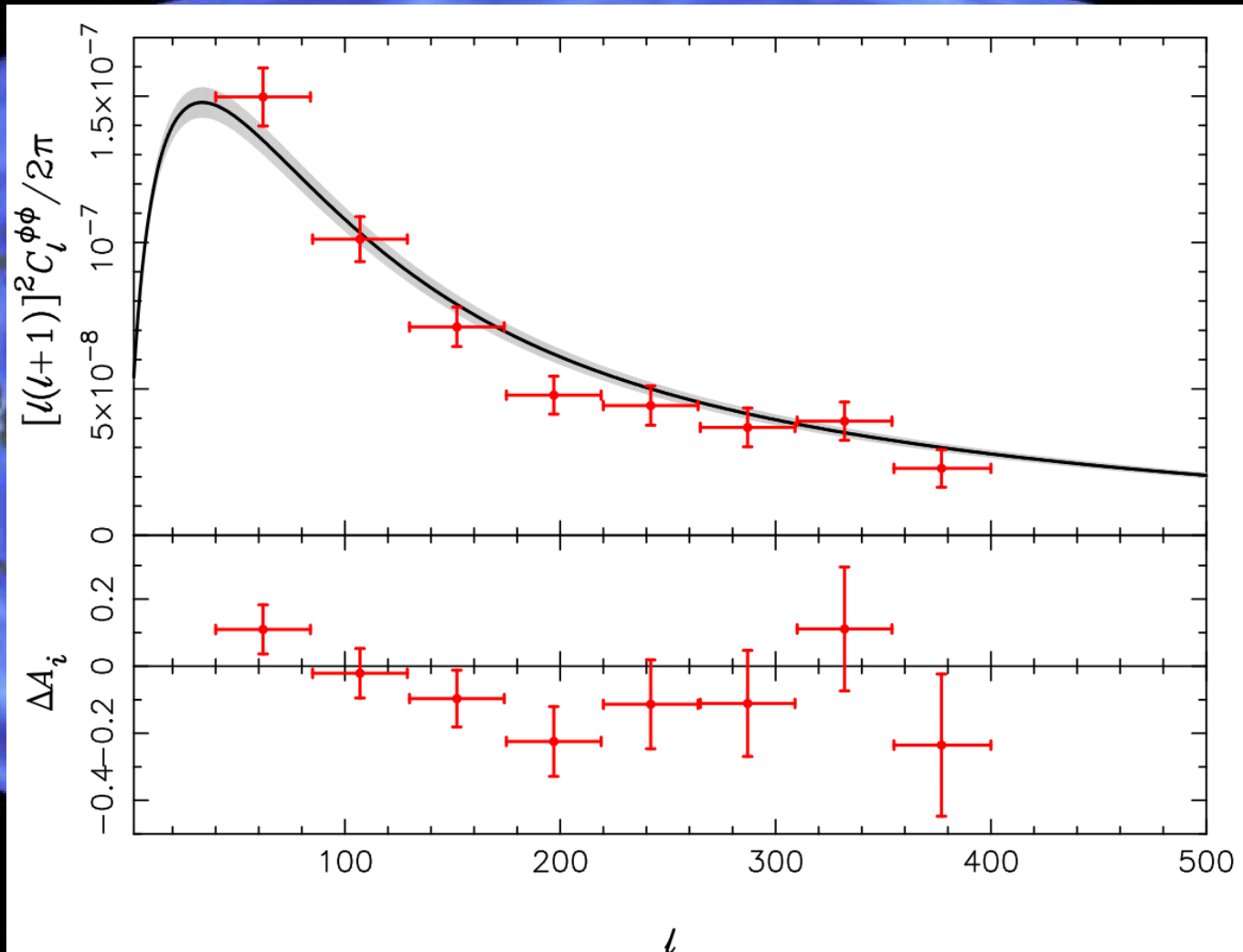
# gravitational lensing



# gravitational lensing



# projected mass distribution



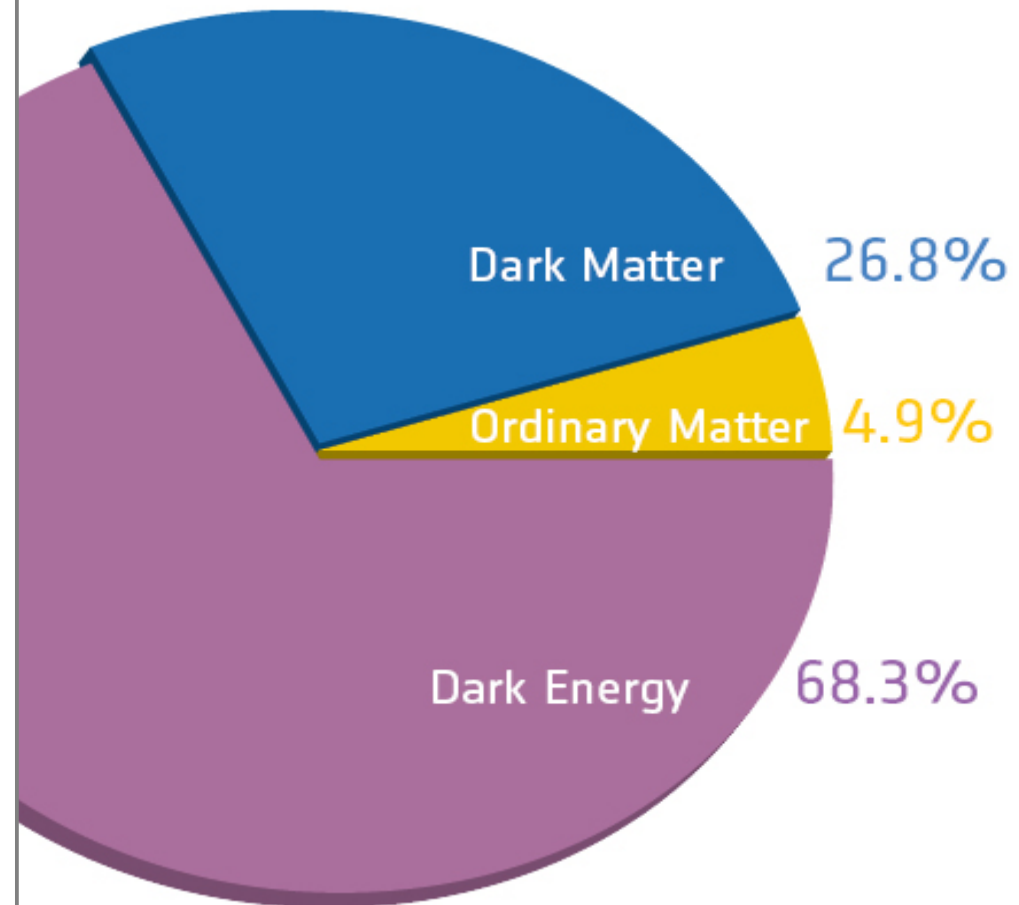
# Precision cosmology

Acoustic scale is most precisely known cosmic parameter:  
 $1.19345 \pm 0.00076^\circ$

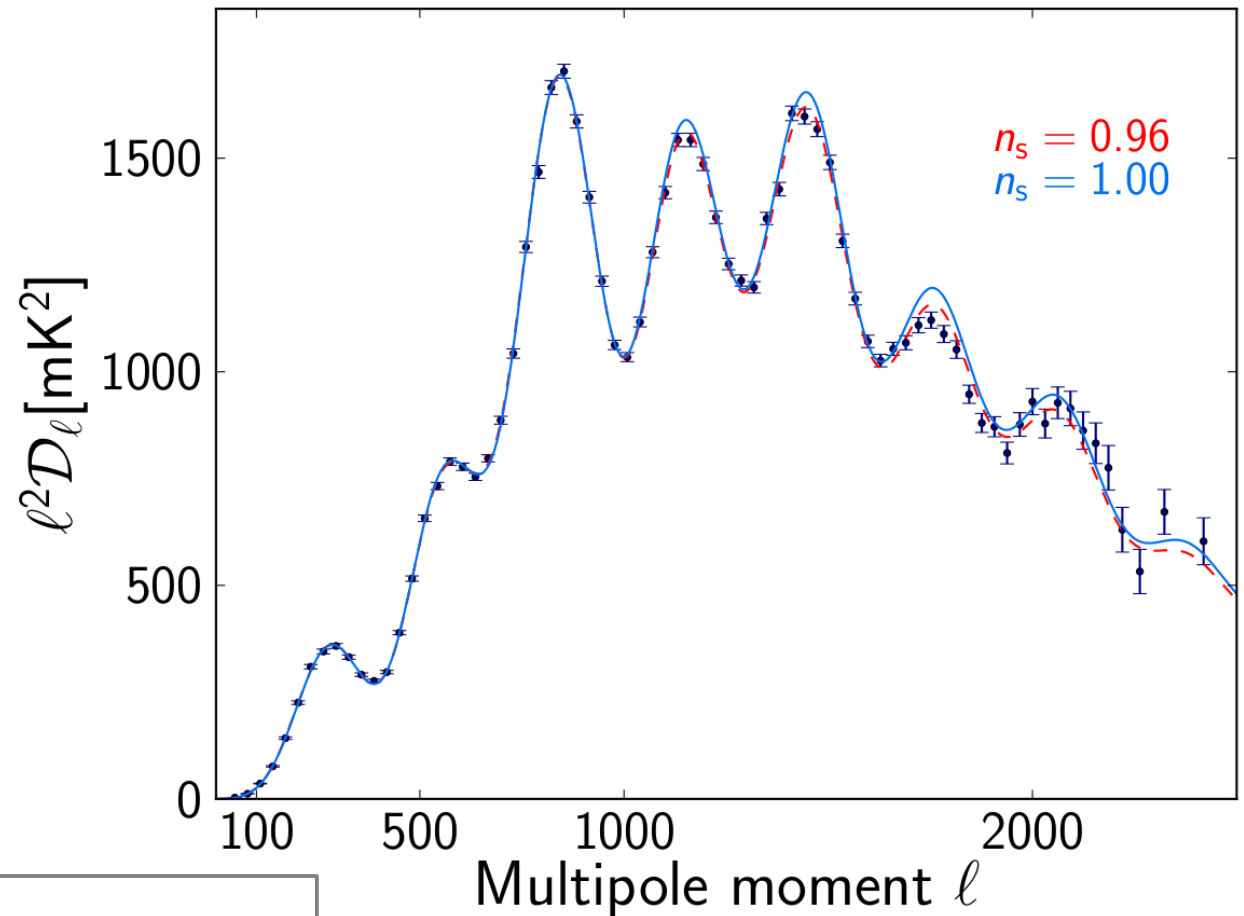
Expansion rate of the Universe, the Hubble constant:  
 $H_0 = 67.15 \text{ km/s/Mpc}$   
(previously 69.32 km/s/Mpc)

Age of the Universe:  
13.82 billion years  
(previously 13.77 billion years)

Primordial spectral index: 0.96  
Significant deviation from scale  
Invariance as predicted by inflation



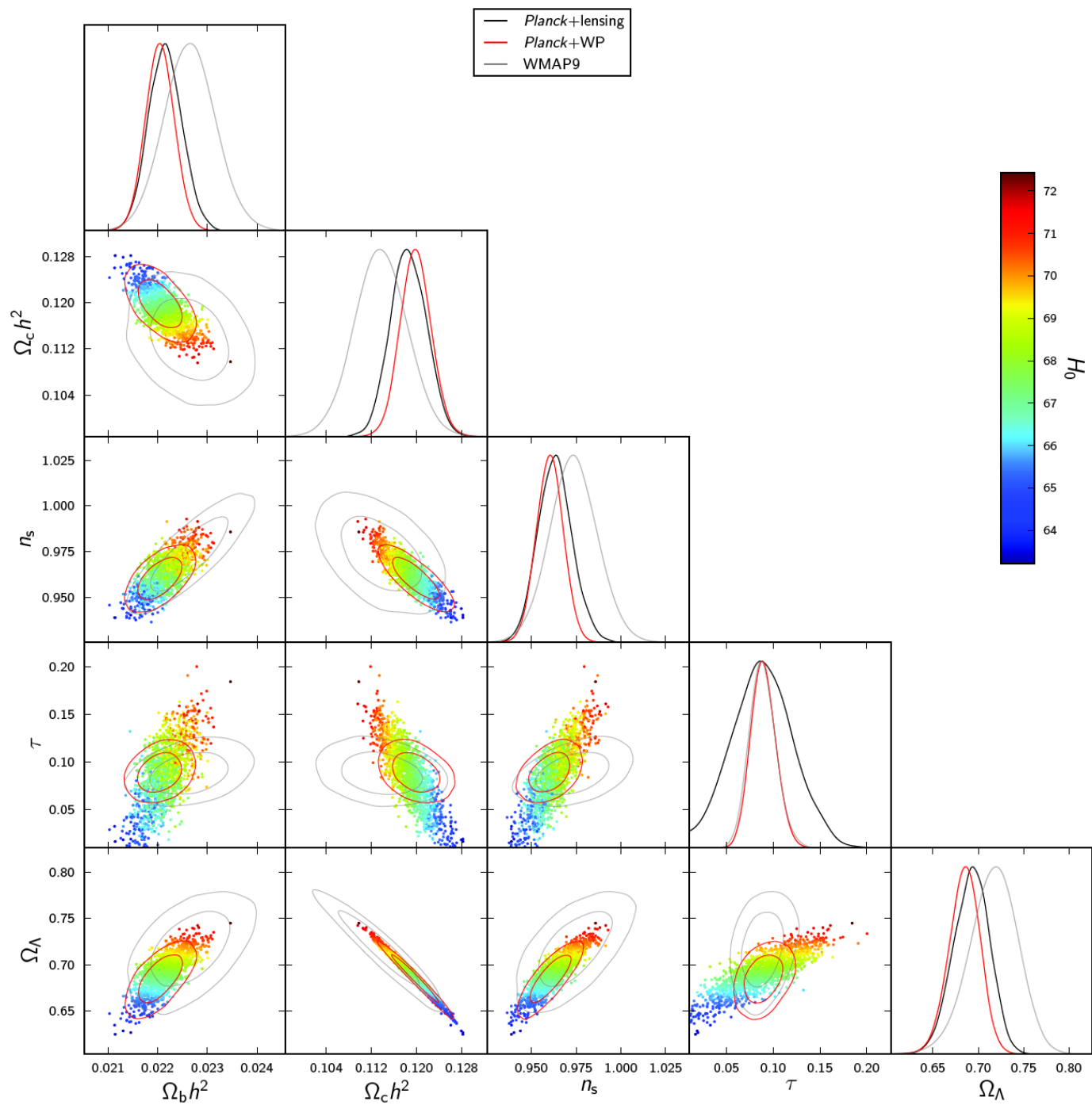
# Precision cosmology



Primordial spectral index: 0.96  
Significant deviation from scale  
Invariance as predicted by inflation

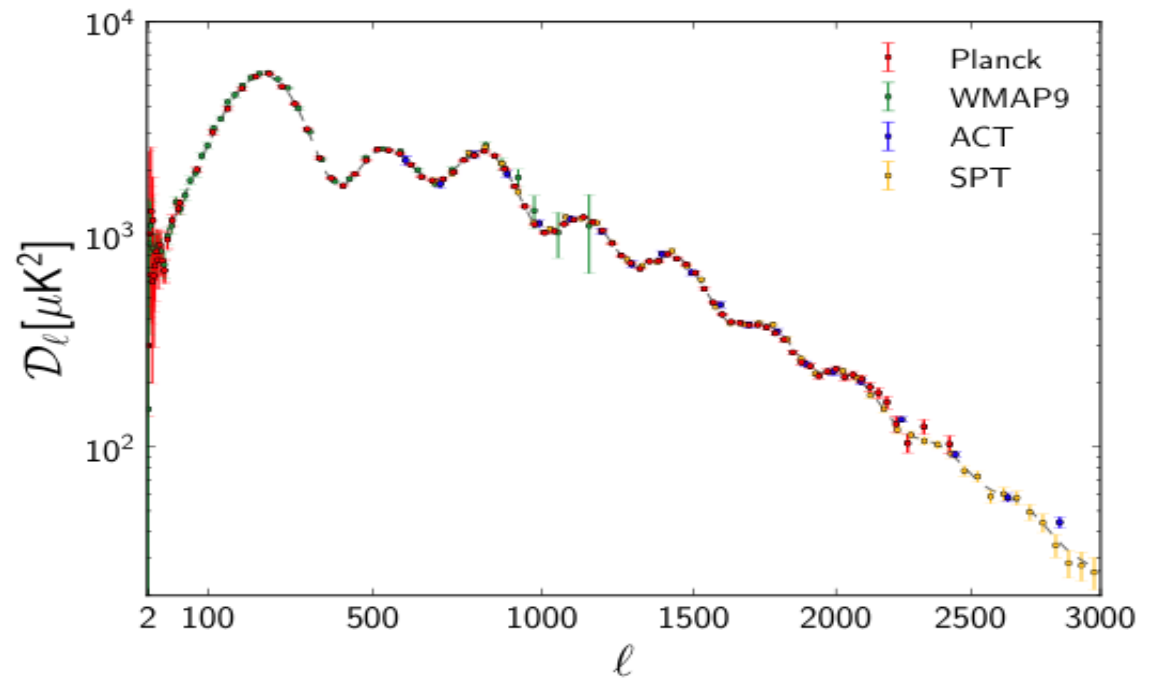
Parameter	<i>Planck</i> (CMB+lensing)	
	Best fit	68 % limits
$\Omega_b h^2$ . . . . .	0.022242	$0.02217 \pm 0.00033$
$\Omega_c h^2$ . . . . .	0.11805	$0.1186 \pm 0.0031$
$100\theta_{MC}$ . . . . .	1.04150	$1.04141 \pm 0.00067$
$\tau$ . . . . .	0.0949	$0.089 \pm 0.032$
$n_s$ . . . . .	0.9675	$0.9635 \pm 0.0094$
$\ln(10^{10} A_s)$ . . . . .	3.098	$3.085 \pm 0.057$



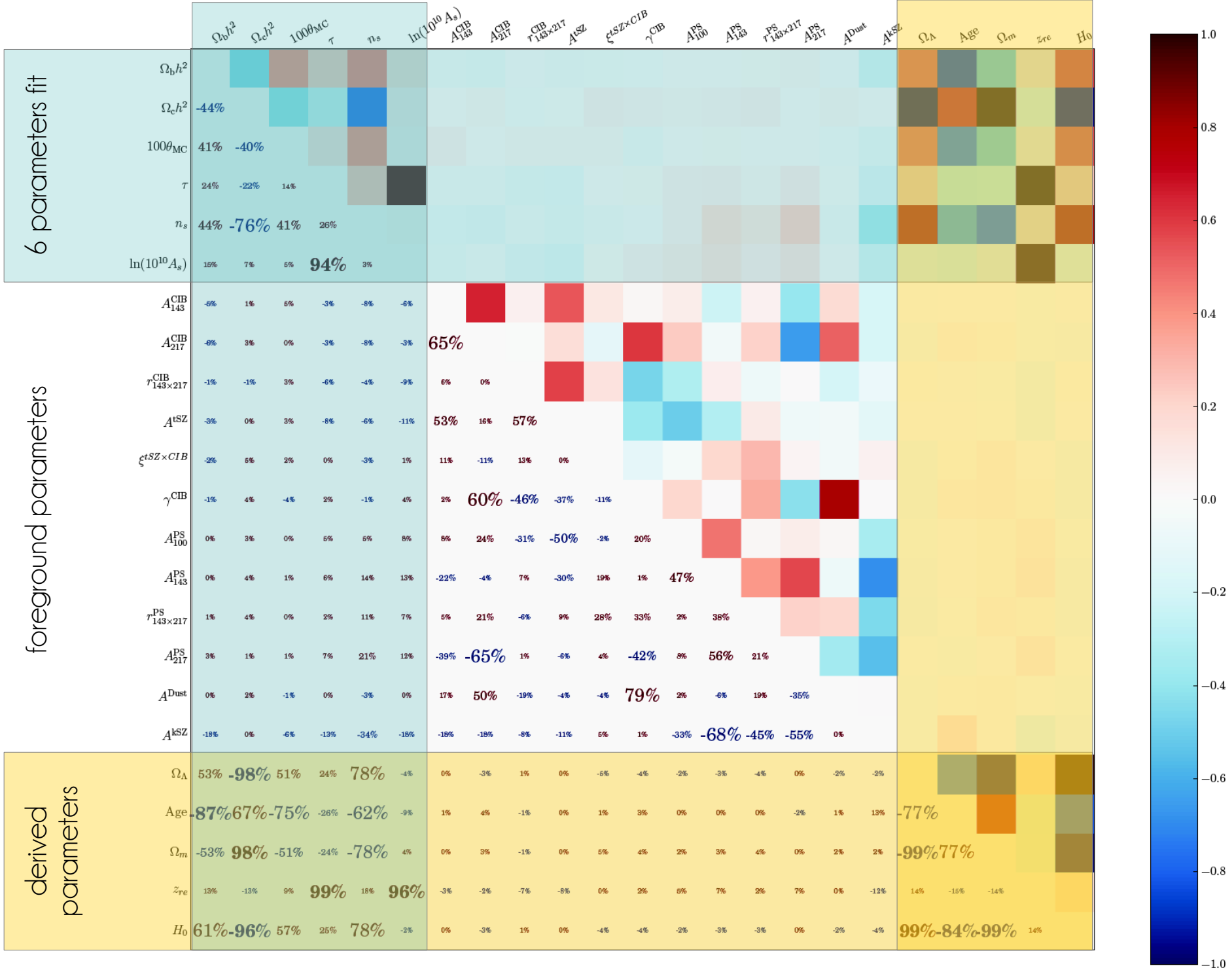


**Fig. 2.** Comparison of the base  $\Lambda$ CDM model parameters for *Planck*+lensing only (colour coded samples), and the 68% and 95% constraint contours adding WMAP low- $\ell$  polarization (WP, red contours), compared to textitWMAP9 (Bennett et al., 2012, grey contours).

Parameter	<i>Planck</i> (CMB+lensing)		<i>Planck</i> +WP+highL+BAO	
	Best fit	68 % limits	Best fit	68 % limits
$\Omega_b h^2$ . . . . .	0.022242	$0.02217 \pm 0.00033$	0.022161	$0.02214 \pm 0.00024$
$\Omega_c h^2$ . . . . .	0.11805	$0.1186 \pm 0.0031$	0.11889	$0.1187 \pm 0.0017$
$100\theta_{MC}$ . . . . .	1.04150	$1.04141 \pm 0.00067$	1.04148	$1.04147 \pm 0.00056$
$\tau$ . . . . .	0.0949	$0.089 \pm 0.032$	0.0952	$0.092 \pm 0.013$
$n_s$ . . . . .	0.9675	$0.9635 \pm 0.0094$	0.9611	$0.9608 \pm 0.0054$
$\ln(10^{10} A_s)$ . . . . .	3.098	$3.085 \pm 0.057$	3.0973	$3.091 \pm 0.025$



Parameter	<i>Planck</i> (CMB+lensing)		<i>Planck</i> +WP+highL+BAO	
	Best fit	68 % limits	Best fit	68 % limits
$\Omega_b h^2$ . . . . .	0.022242	$0.02217 \pm 0.00033$	0.022161	$0.02214 \pm 0.00024$
$\Omega_c h^2$ . . . . .	0.11805	$0.1186 \pm 0.0031$	0.11889	$0.1187 \pm 0.0017$
$100\theta_{\text{MC}}$ . . . . .	1.04150	$1.04141 \pm 0.00067$	1.04148	$1.04147 \pm 0.00056$
$\tau$ . . . . .	0.0949	$0.089 \pm 0.032$	0.0952	$0.092 \pm 0.013$
$n_s$ . . . . .	0.9675	$0.9635 \pm 0.0094$	0.9611	$0.9608 \pm 0.0054$
$\ln(10^{10} A_s)$ . . . . .	3.098	$3.085 \pm 0.057$	3.0973	$3.091 \pm 0.025$
$\Omega_\Lambda$ . . . . .	0.6964	$0.693 \pm 0.019$	0.6914	$0.692 \pm 0.010$
$\Omega_m$ . . . . .	0.3036	$0.307 \pm 0.019$		
$\sigma_8$ . . . . .	0.8285	$0.823 \pm 0.018$	0.8288	$0.826 \pm 0.012$
$z_{\text{re}}$ . . . . .	11.45	$10.8^{+3.1}_{-2.5}$	11.52	$11.3 \pm 1.1$
$H_0$ . . . . .	68.14	$67.9 \pm 1.5$	67.77	$67.80 \pm 0.77$
$10^9 A_s$ . . . . .	2.215	$2.19^{+0.12}_{-0.14}$		
$\Omega_m h^2$ . . . . .	0.14094	$0.1414 \pm 0.0029$		
$\Omega_m h^3$ . . . . .	0.09603	$0.09593 \pm 0.00058$		
$Y_{\text{P}}$ . . . . .	0.247785	$0.24775 \pm 0.00014$		
Age/Gyr . . . . .	13.784	$13.796 \pm 0.058$	13.7965	$13.798 \pm 0.037$
$z_*$ . . . . .	1090.01	$1090.16 \pm 0.65$		
$r_*$ . . . . .	144.58	$144.96 \pm 0.66$		
$100\theta_*$ . . . . .	1.04164	$1.04156 \pm 0.00066$	1.04163	$1.04162 \pm 0.00056$
$z_{\text{drag}}$ . . . . .	1059.59	$1059.43 \pm 0.64$		
$r_{\text{drag}}$ . . . . .	147.74	$147.70 \pm 0.63$	147.611	$147.68 \pm 0.45$
$k_{\text{D}}$ . . . . .	0.13998	$0.13996 \pm 0.00062$		
$100\theta_{\text{D}}$ . . . . .	0.161196	$0.16129 \pm 0.00036$		
$z_{\text{eq}}$ . . . . .	3352	$3362 \pm 69$		
$100\theta_{\text{eq}}$ . . . . .	0.8224	$0.821 \pm 0.013$		
$r_{\text{drag}}/D_{\text{V}}(0.57)$ . . . . .	0.07207	$0.0719 \pm 0.0011$		

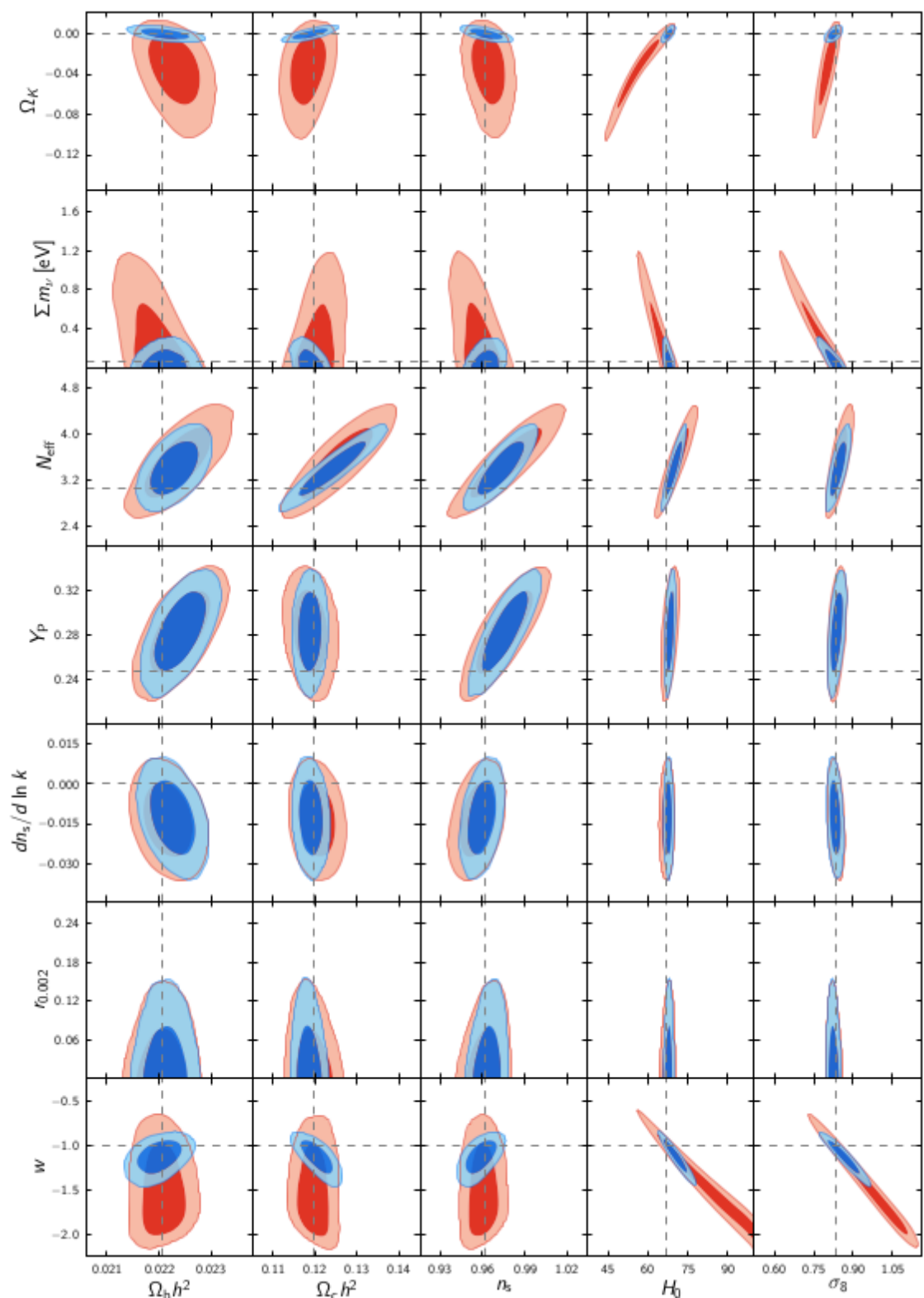


**Figure 21.** Correlation matrix between all the cosmological (top block), foreground (middle block), and derived (bottom block) parameters, estimated using the Plik likelihood.

# Extended $p$

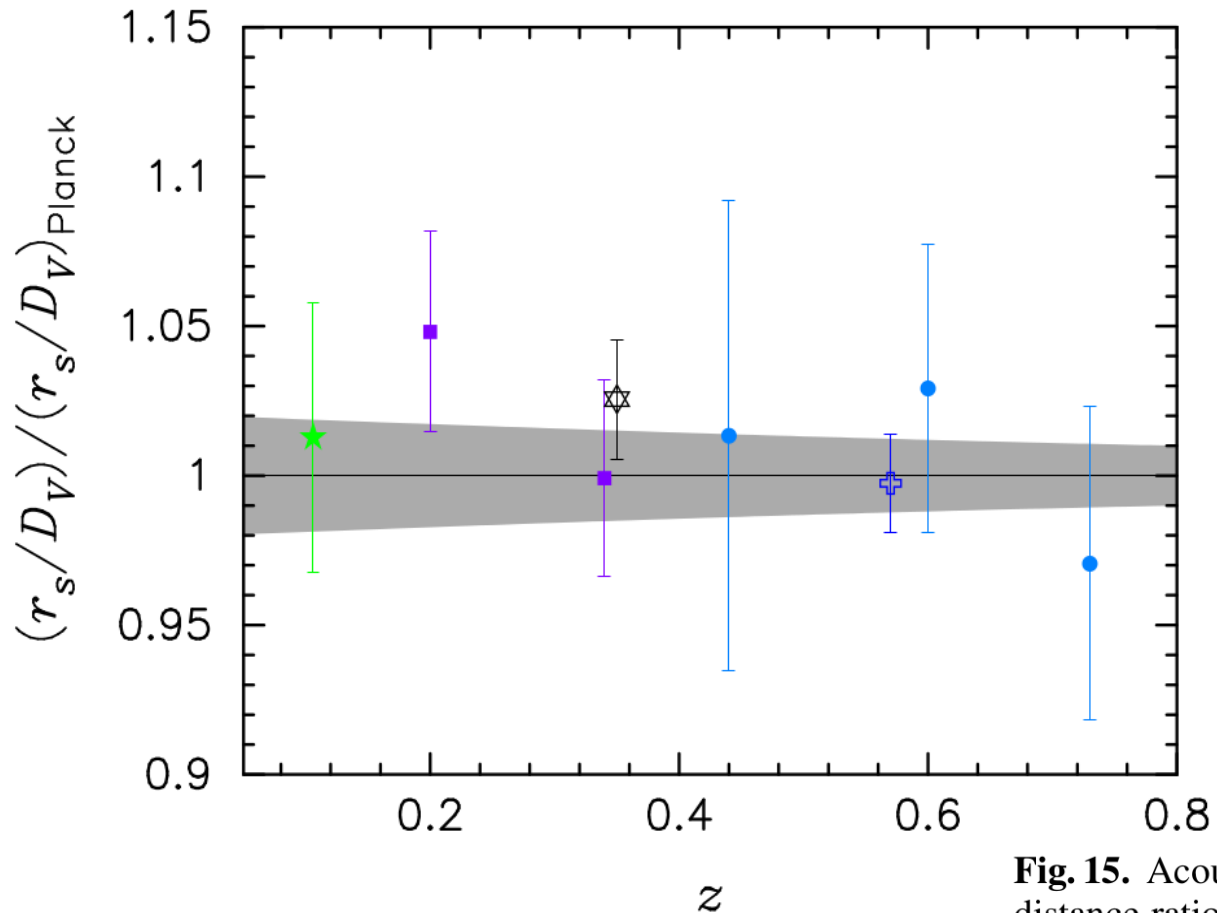
curvature of the Universe, mass  
spectral curvature, gra  
adiabatic in

Parameter	<i>Planck+WP</i>		<i>Planck+V</i>	
	Best fit	95% limits	Best fit	$\chi^2$
$\Omega_K$ . . . . .	-0.0105	-0.037 <sup>+0.043</sup> <sub>-0.049</sub>	0.0000	0.
$\Sigma m_\nu$ [eV] . . . . .	0.022	< 0.933	0.002	
$N_{\text{eff}}$ . . . . .	3.08	3.51 <sup>+0.80</sup> <sub>-0.74</sub>	3.08	
$Y_P$ . . . . .	0.2583	0.283 <sup>+0.045</sup> <sub>-0.048</sub>	0.2736	(
$dn_s/d \ln k$ . . . . .	-0.0090	-0.013 <sup>+0.018</sup> <sub>-0.018</sub>	-0.0102	-
$r_{0.002}$ . . . . .	0.000	< 0.120	0.000	
$w$ . . . . .	-1.20	-1.49 <sup>+0.65</sup> <sub>-0.57</sub>	-1.076	.



# Data consistency

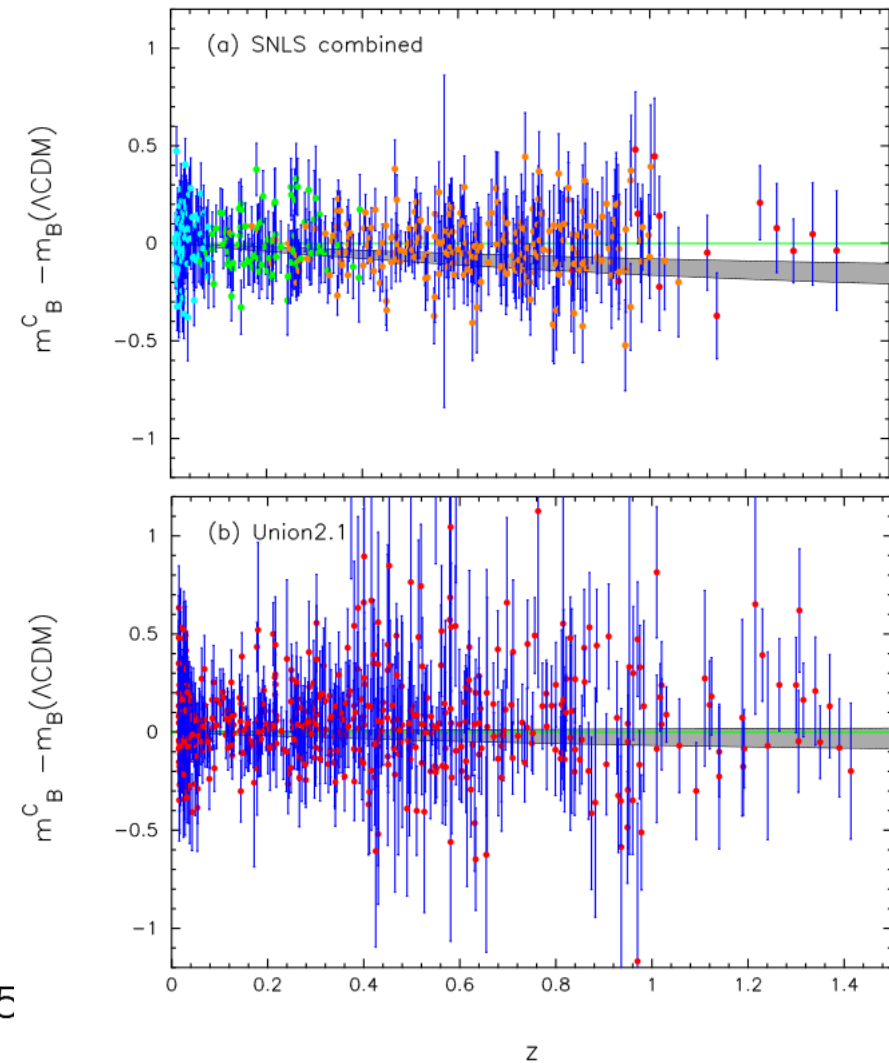
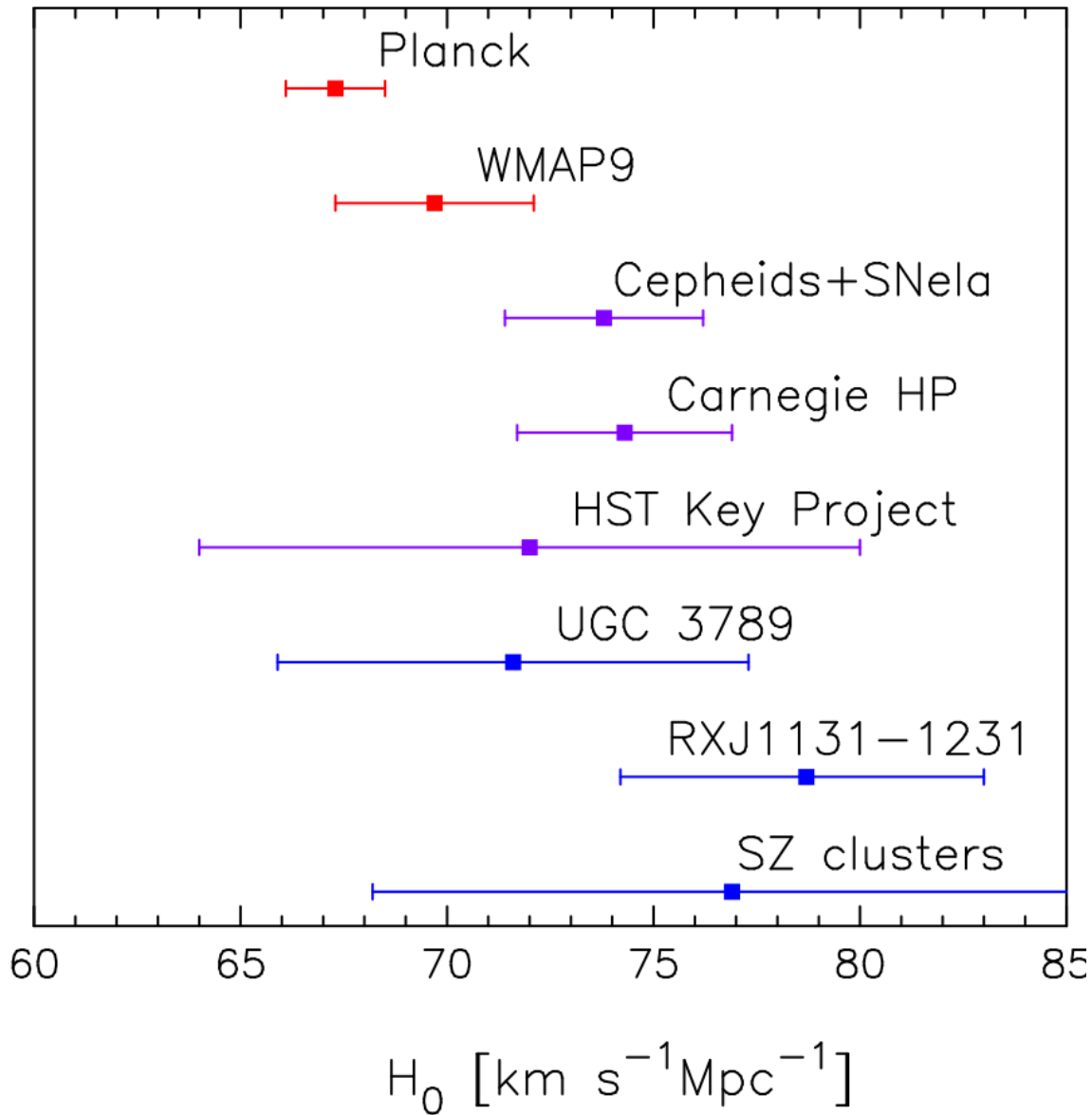
## acoustic-scale distance ratio



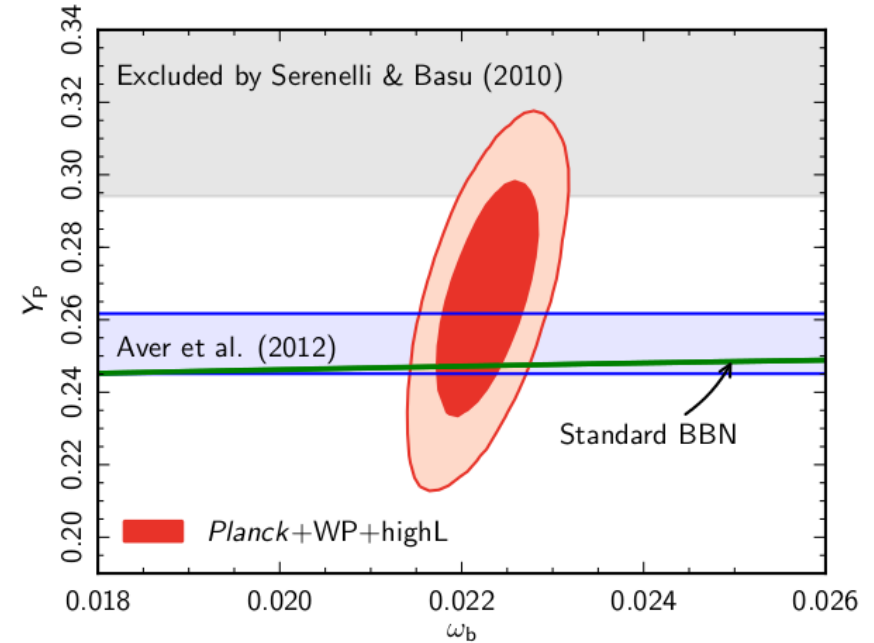
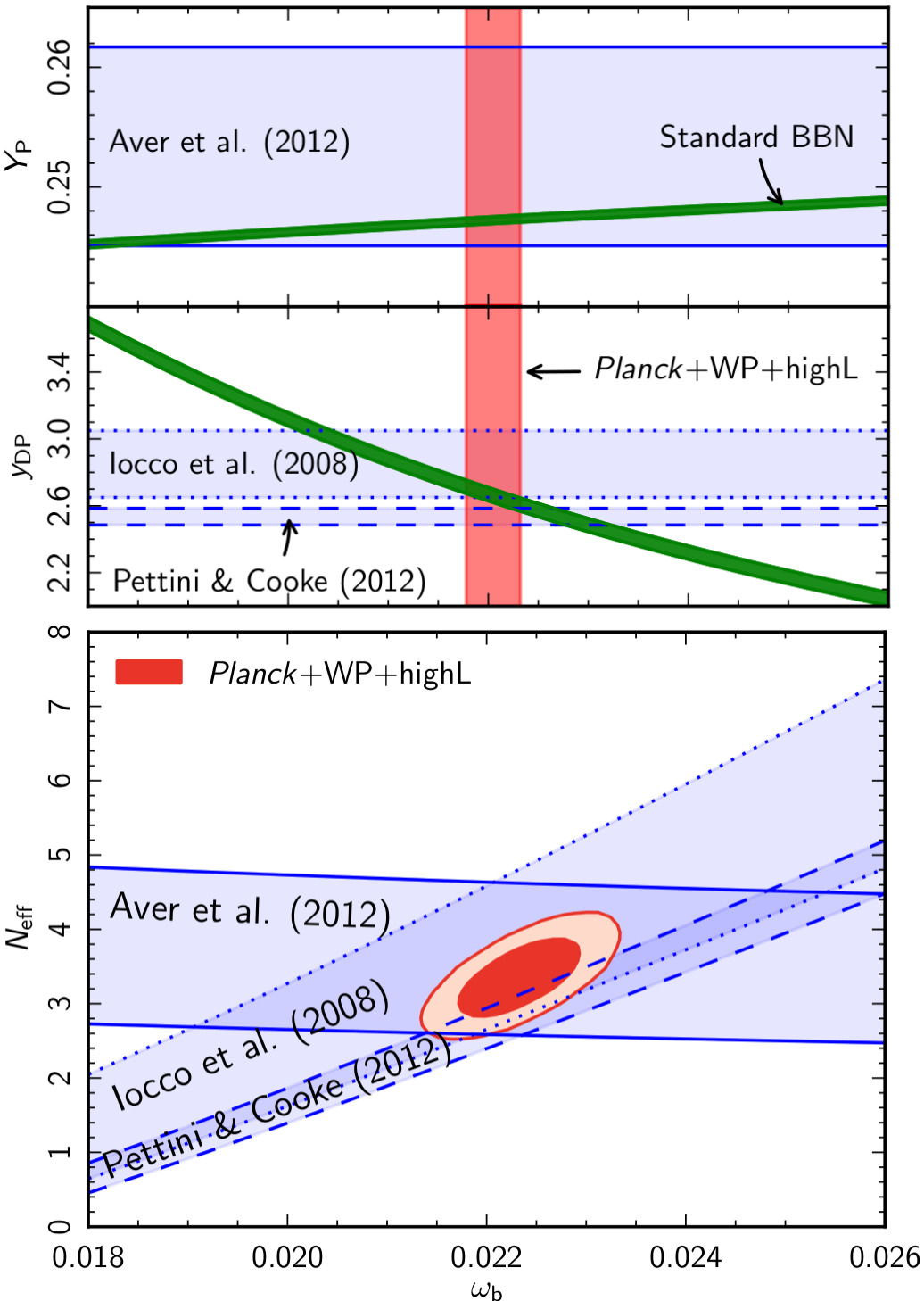
**Fig. 15.** Acoustic-scale distance ratio  $r_s/D_V(z)$  divided by the distance ratio of the *Planck* base  $\Lambda$ CDM model. The points are colour coded as follows: green star (6dF); purple squares (SDSS DR7 as analyzed by Percival et al., 2010); black star (SDSS DR7 as analyzed by Padmanabhan et al., 2012); blue cross (BOSS DR9); and blue circles (WiggleZ). The grey band shows the approximate  $\pm 1\sigma$  range allowed by *Planck* (computed from the CosmoMC chains).

# Data consistency

## Hubble constant & Lambda



# Data consistency Big Bang Nucleosynthesis



**Fig. 30.** The green stripe shows the predictions of standard BBN for the primordial abundance of helium 4 as a function of the baryon density (with 68% CL errors on nuclear reaction rates). The horizontal band shows observational bounds on helium 4 compiled by Aver et al. (2012) (68% CL), while the grey region in the upper part of the figure delineates the conservative 95% upper bound inferred from solar helium abundance by (Serenelli & Basu, 2010). Finally, we show the 68% and 95% joint contours on  $(\omega_b, Y_P)$  inferred from *Planck*+WP+highL, when  $Y_P$  is left as a free parameter in the CMB analysis. Both BBN predictions and CMB results assume  $N_{\text{eff}} = 3.046$  and no significant lepton asymmetry.



# Data consistency

## Tensions & Agreements

### Tensions:

#### Hubble constant:

with WMAP

with SN measurements (might go away)

#### primordial power spectrum amplitude:

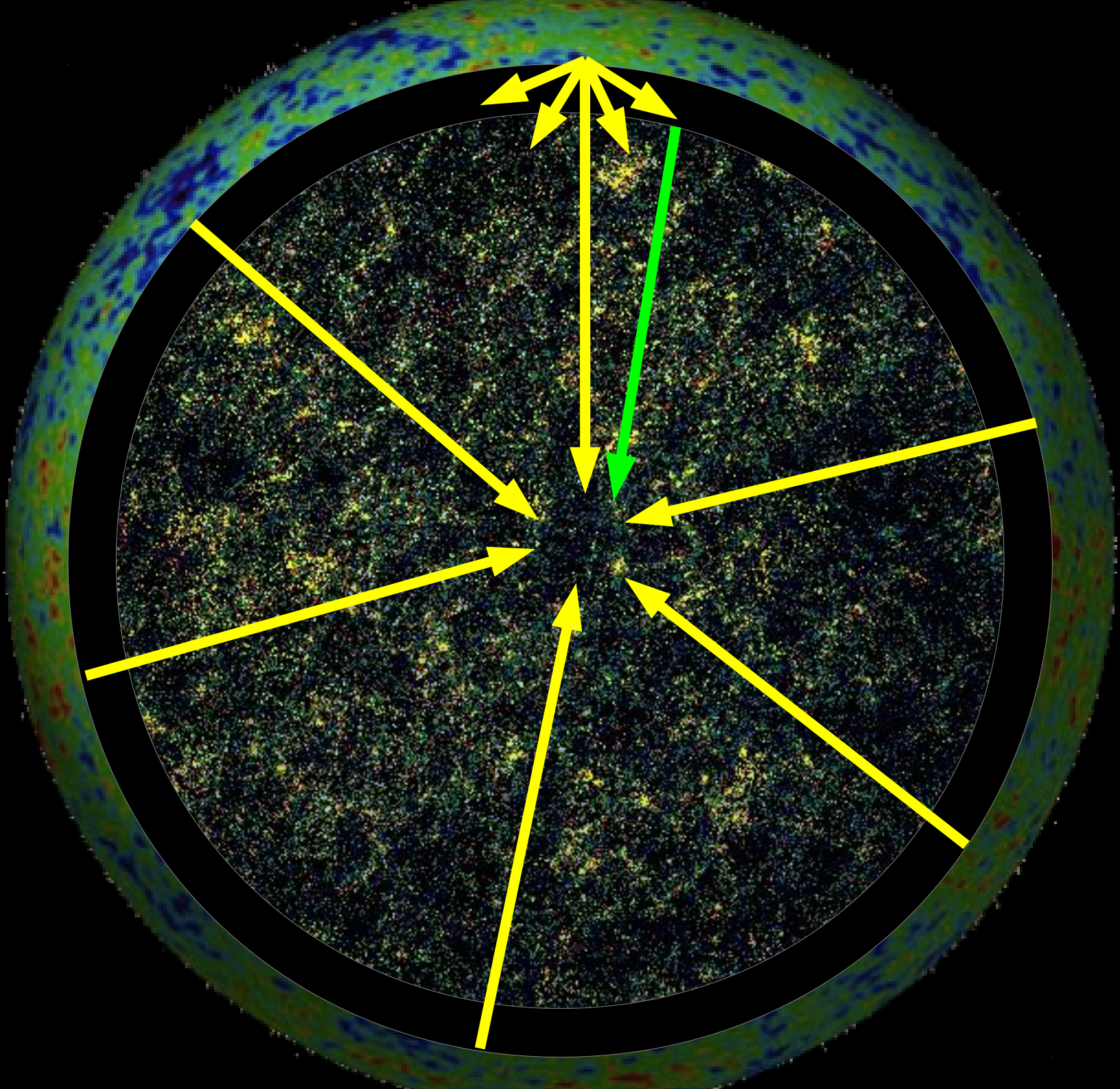
with WMAP (2%)

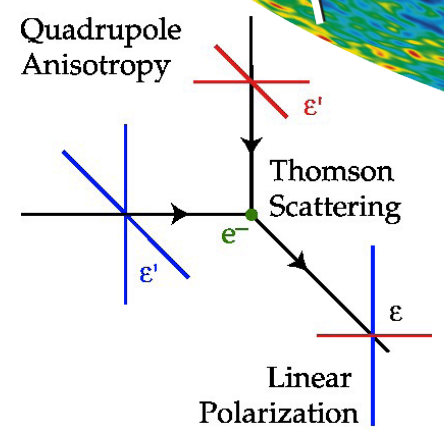
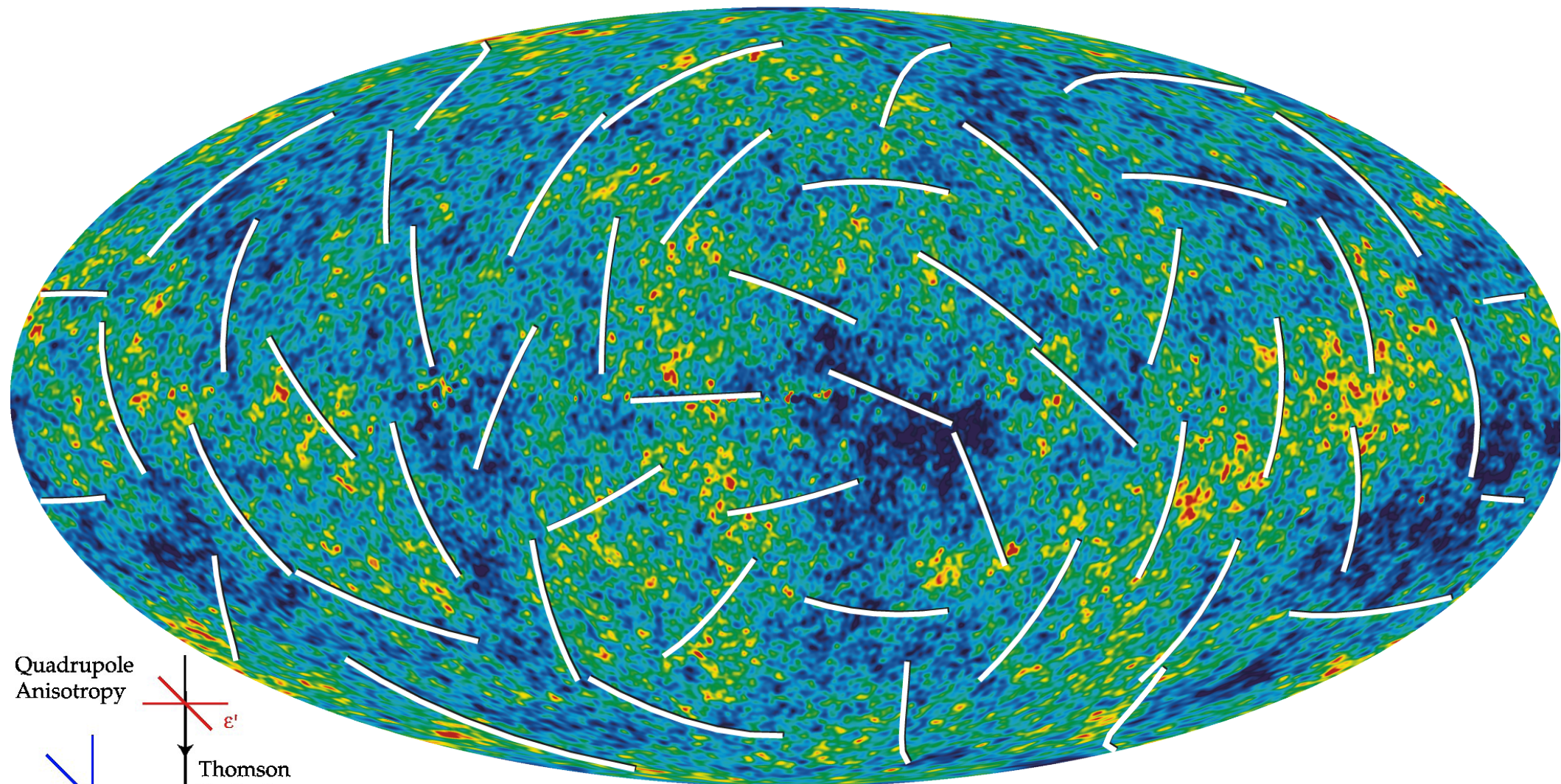
with Planck SZ cluster counts

All tensions only one to a few sigmas

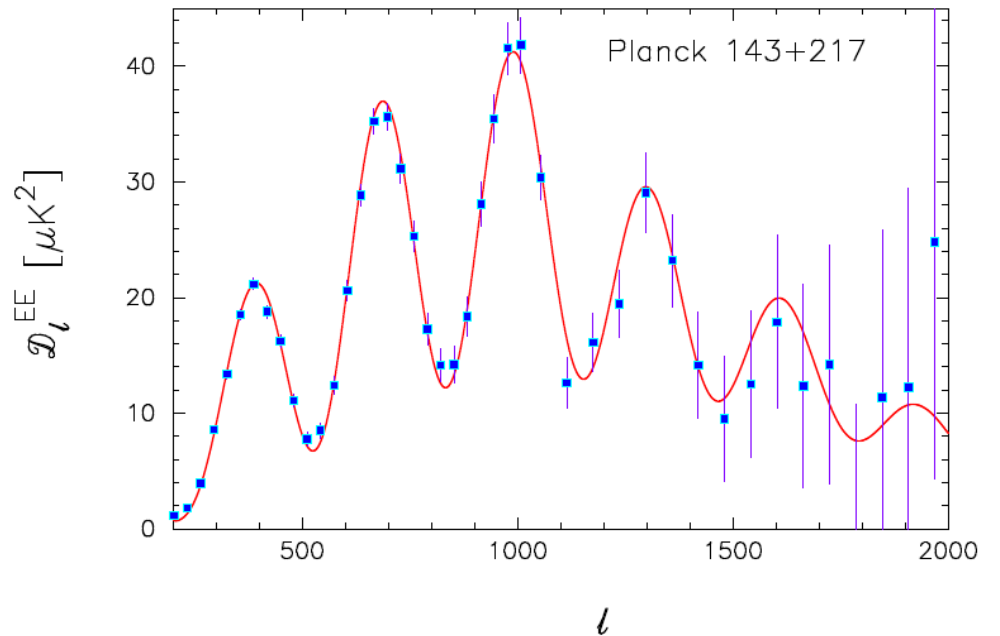
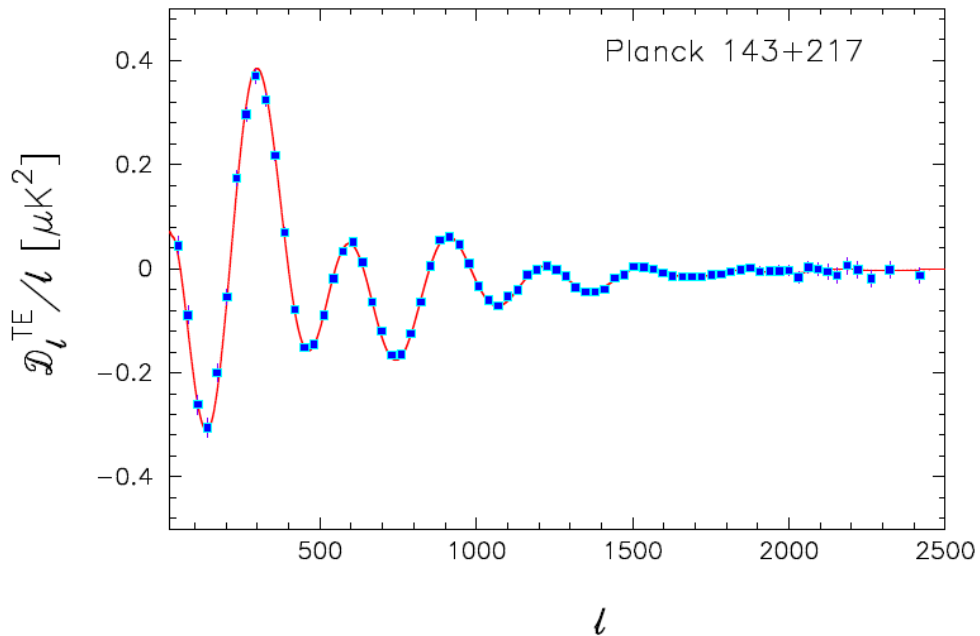
### Agreements:

With LSS power spectrum, with BAOs, with BBN





# Polarization teaser



**Fig. 11.** *Planck TE (left) and EE spectra (right) computed as described in the text. The red lines show the polarization spectra from the base  $\Lambda$ CDM Planck+WP+highL model, which is fitted to the TT data only.*

# non-Gaussian statistics



$$f(x)$$

=

$$\frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

$$\mu$$

=

2.725 Kelvin

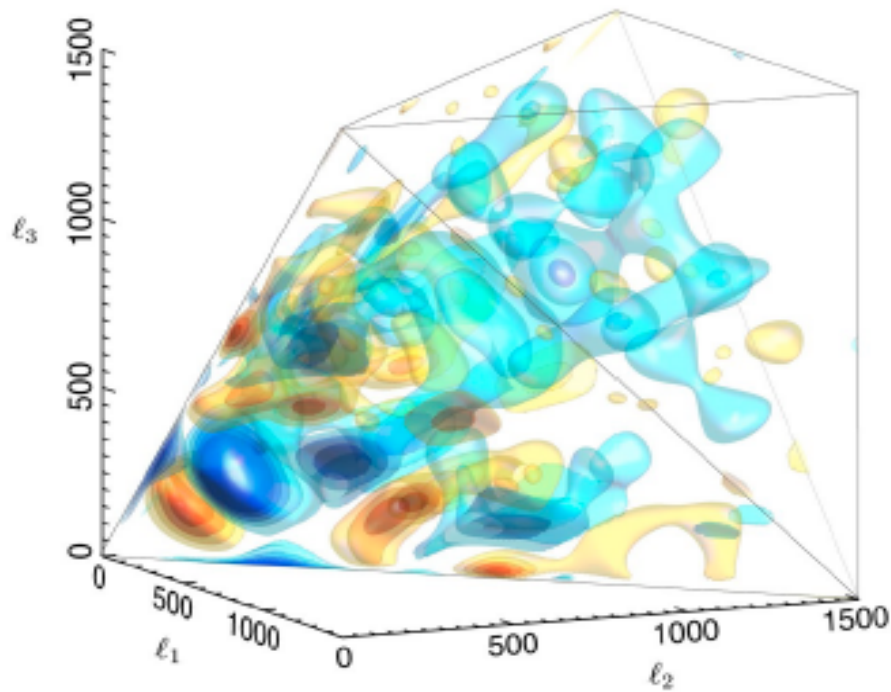
$$\sigma$$

=

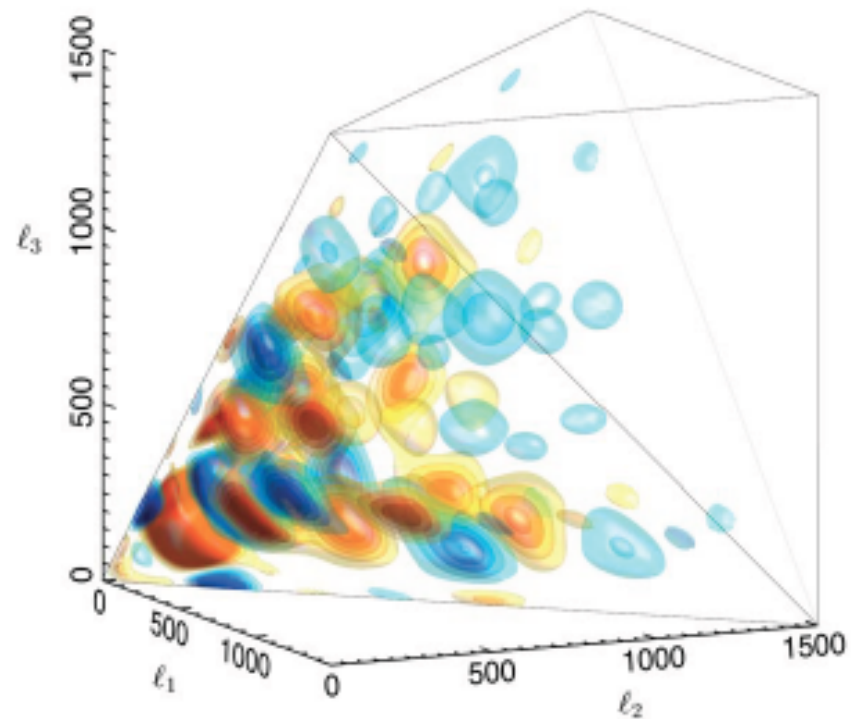
0.000 01 Kelvin

# Bispektrum

## non-Gaussian statistics

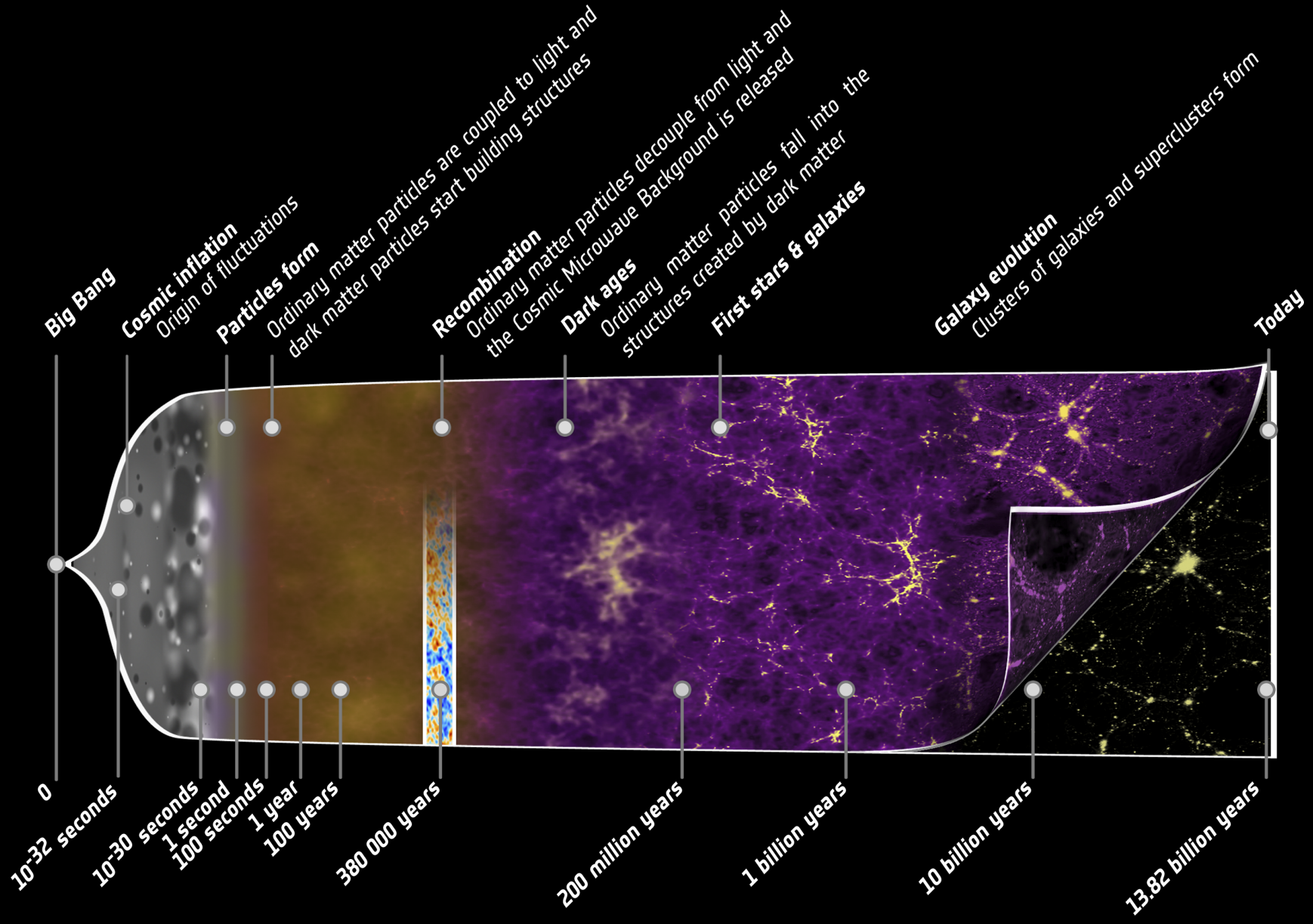


Data



Best Fit Model

	Independent KSW	ISW-lensing subtracted KSW
<b>SMICA</b>		
Local .....	$9.8 \pm 5.8$	$27 \pm 5.8$
Equilateral .....	$-37 \pm 75$	$-42 \pm 75$
Orthogonal .....	$-46 \pm 39$	$-25 \pm 39$



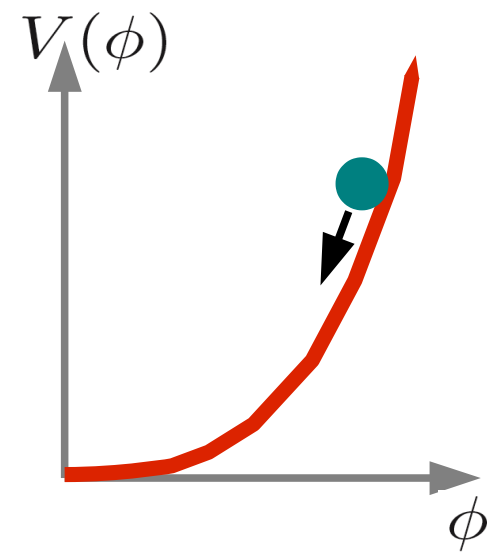
# Inflation

$$\ddot{\phi}(t) + 3 H(t) \dot{\phi}(t) + V_{\phi} = 0$$

$$H^2 = \frac{1}{3 M_{\text{pl}}^2} \left( V(\phi) + \frac{1}{2} \dot{\phi}^2 \right)$$

$$\epsilon_V = \frac{M_{\text{pl}}^2 V_{\phi}^2}{2 V^2}$$

$$\eta_V = \frac{M_{\text{pl}}^2 V_{\phi\phi}}{V}$$

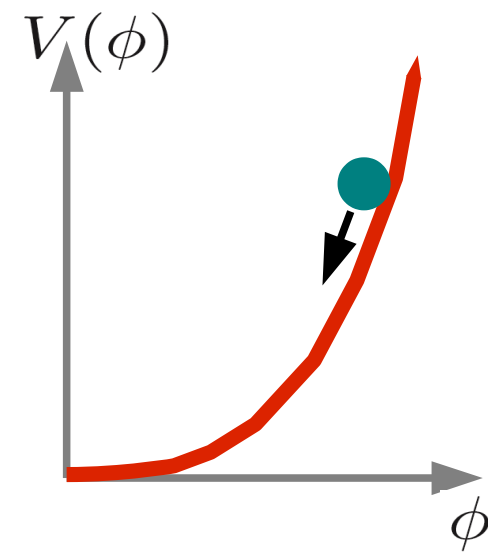


slow role inflation

slow role parameters



# Inflation



$$N_* = \int_{t_*}^{t_e} dt H \approx \frac{1}{M_{\text{pl}}^2} \int_{\phi_*}^{\phi_e} d\phi \frac{V}{V_\phi}$$

$$A_s \approx \frac{V}{24\pi^2 M_{\text{pl}}^4 \epsilon_V}$$

$$A_t \approx \frac{2V}{3\pi^2 M_{\text{pl}}^4} \quad r \equiv \frac{\mathcal{P}_t(k_*)}{\mathcal{P}_R(k_*)} \approx 16\epsilon \approx -8n_t$$

$$n_s - 1 \approx 2\eta_V - 6\epsilon_V$$

$$n_t \approx -2\epsilon_V$$

$$dn_s/d \ln k \approx -16\epsilon_V \eta_V + 24\epsilon_V^2 + 2\xi_V^2$$

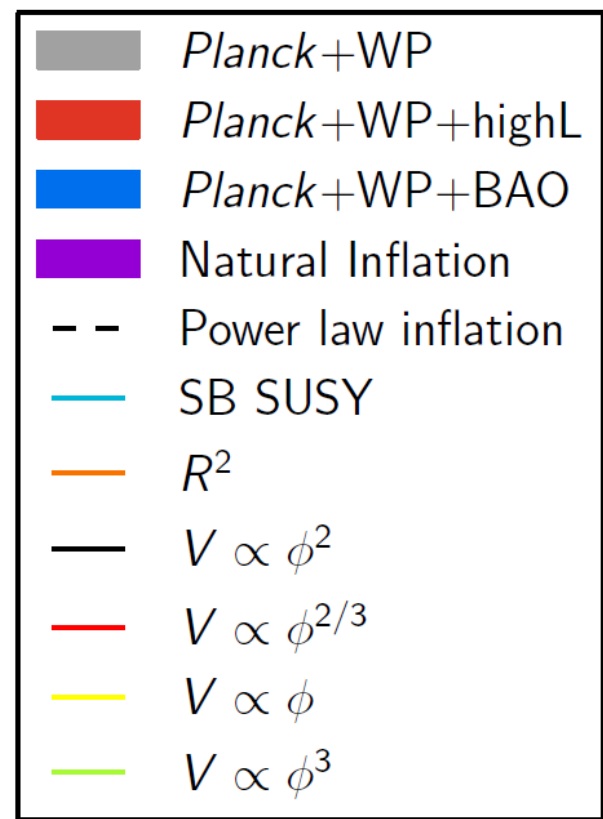
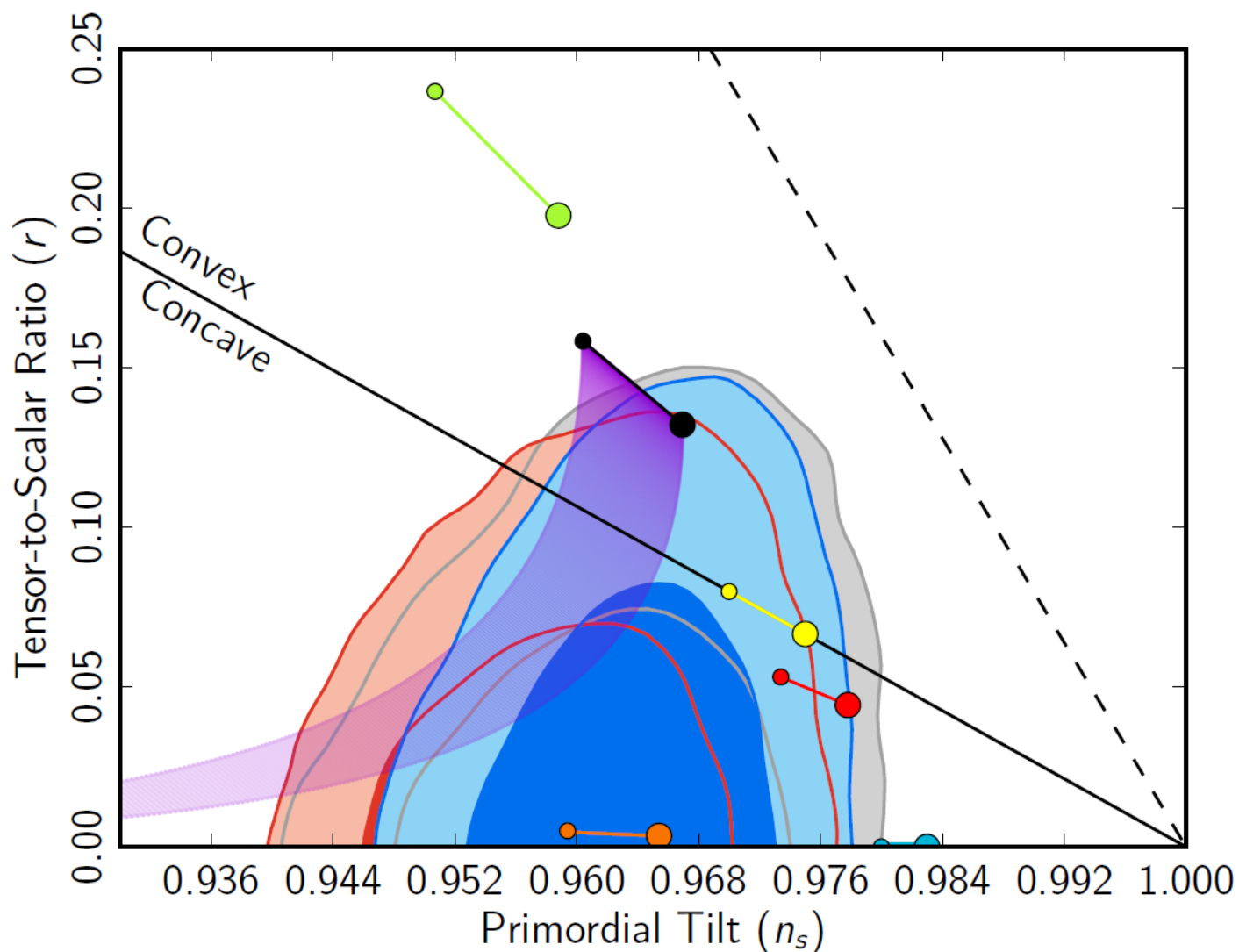
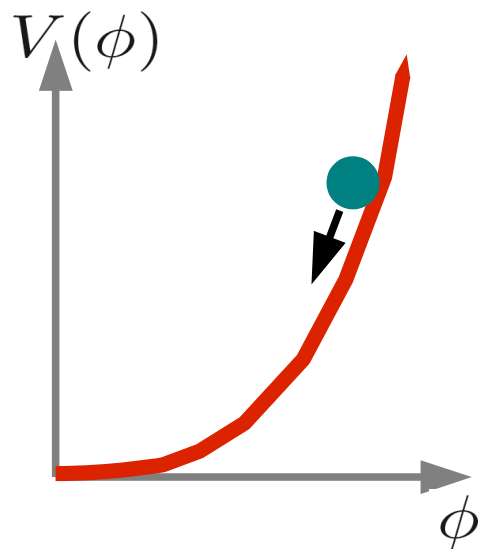
$$dn_t/d \ln k \approx -4\epsilon_V \eta_V + 8\epsilon_V^2$$

$$d^2 n_s/d \ln k^2 \approx -192\epsilon_V^3 + 192\epsilon_V^2 \eta_V - 32\epsilon_V \eta_V^2 \\ - 24\epsilon_V \xi_V^2 + 2\eta_V \xi_V^2 + 2\omega_V^3,$$

$$\epsilon_V = \frac{M_{\text{pl}}^2 V_\phi^2}{2V^2}$$

$$\eta_V = \frac{M_{\text{pl}}^2 V_{\phi\phi}}{V}$$

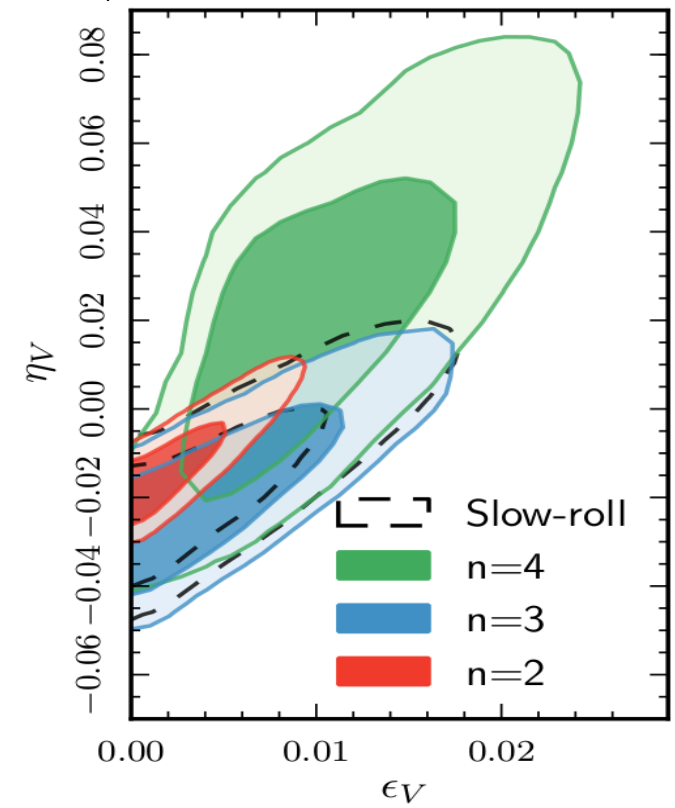
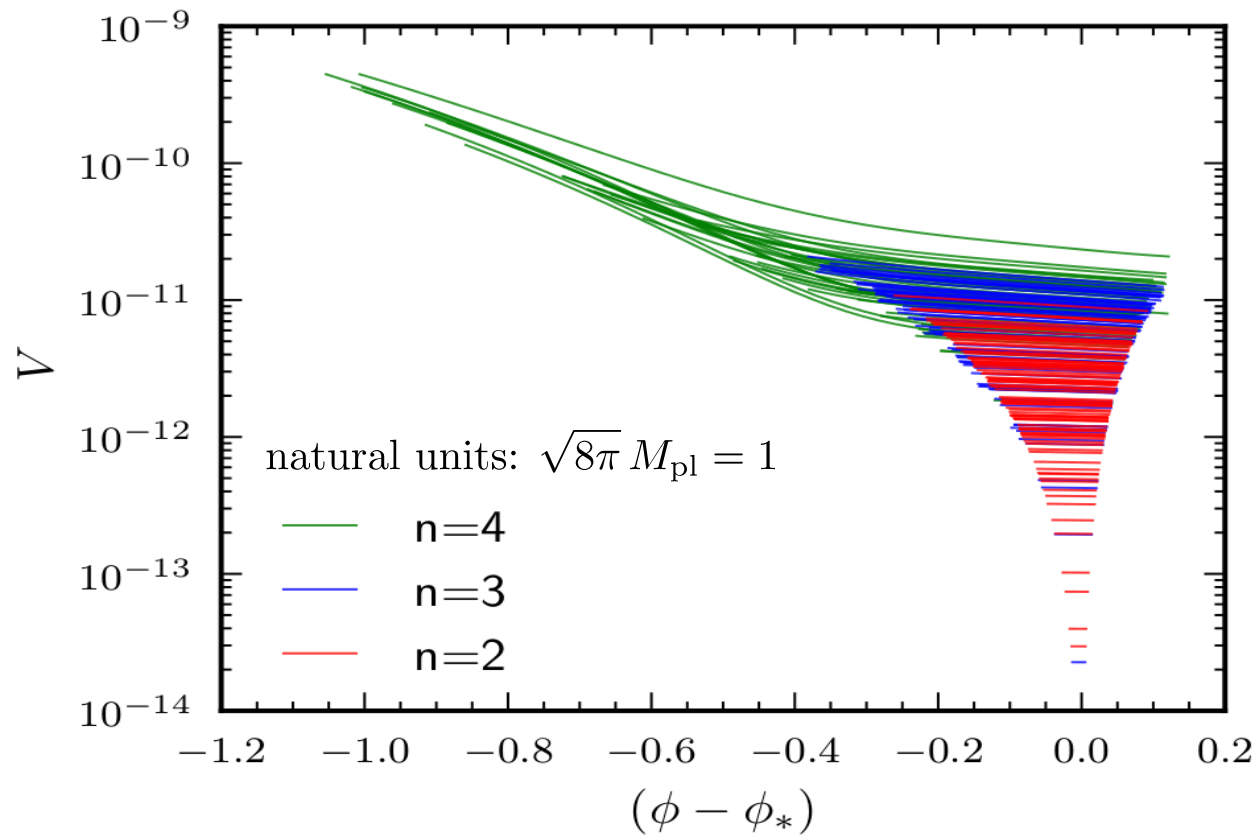
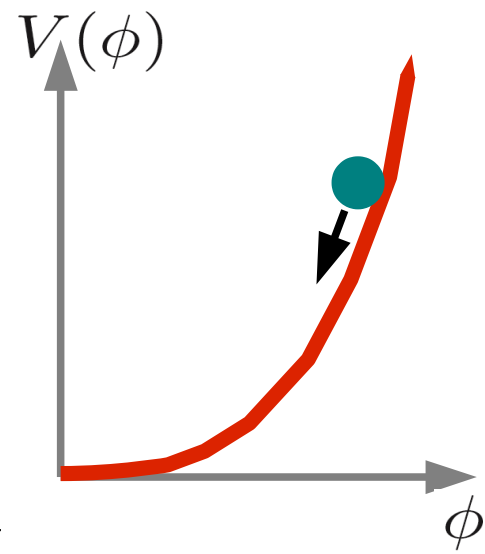
# Inflation



# Inflation

$$\epsilon_V = \frac{M_{\text{pl}}^2 V_\phi^2}{2V^2}$$

$$\eta_V = \frac{M_{\text{pl}}^2 V_{\phi\phi}}{V}$$



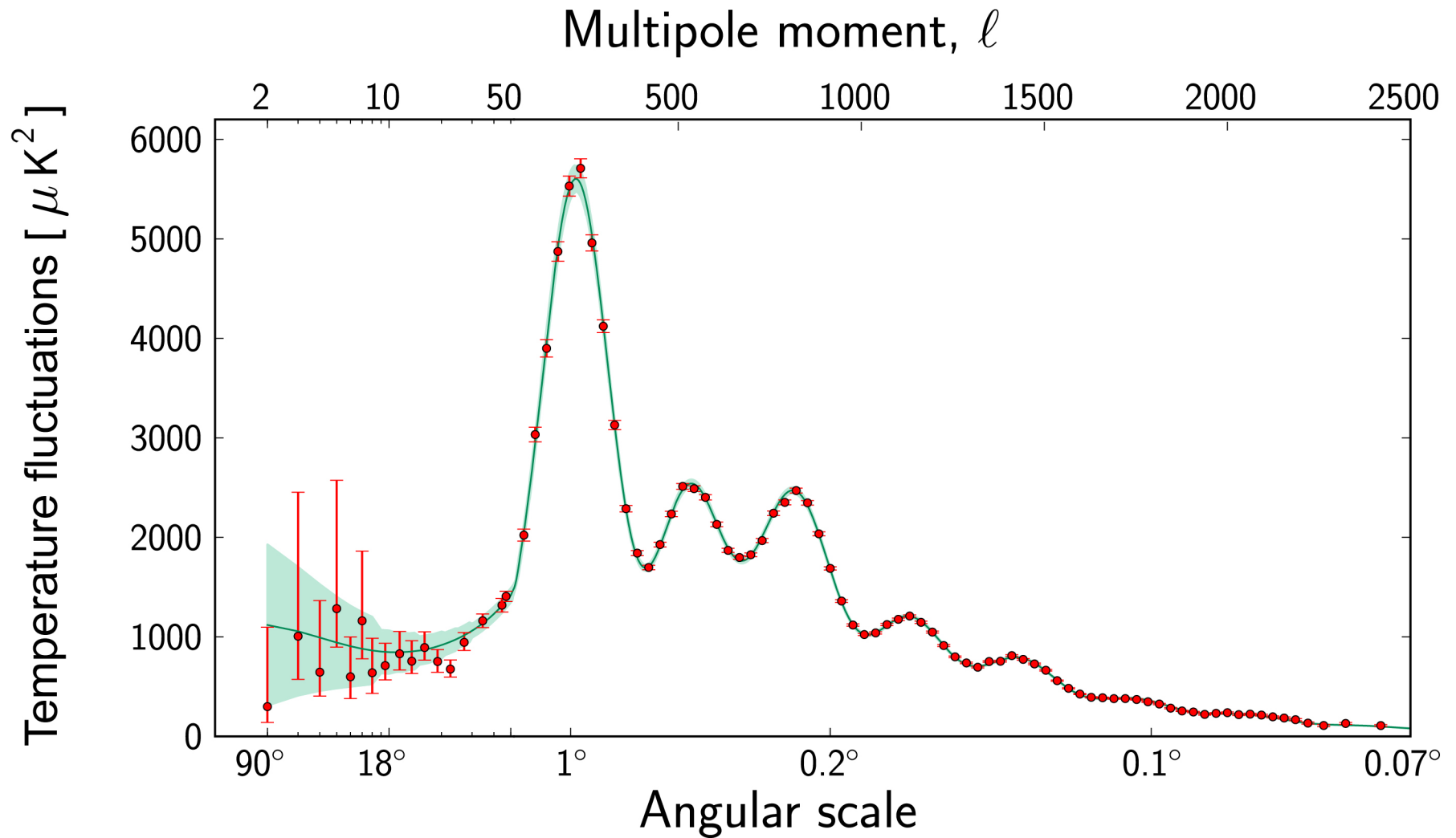
# Anomalies

„Missing Power“

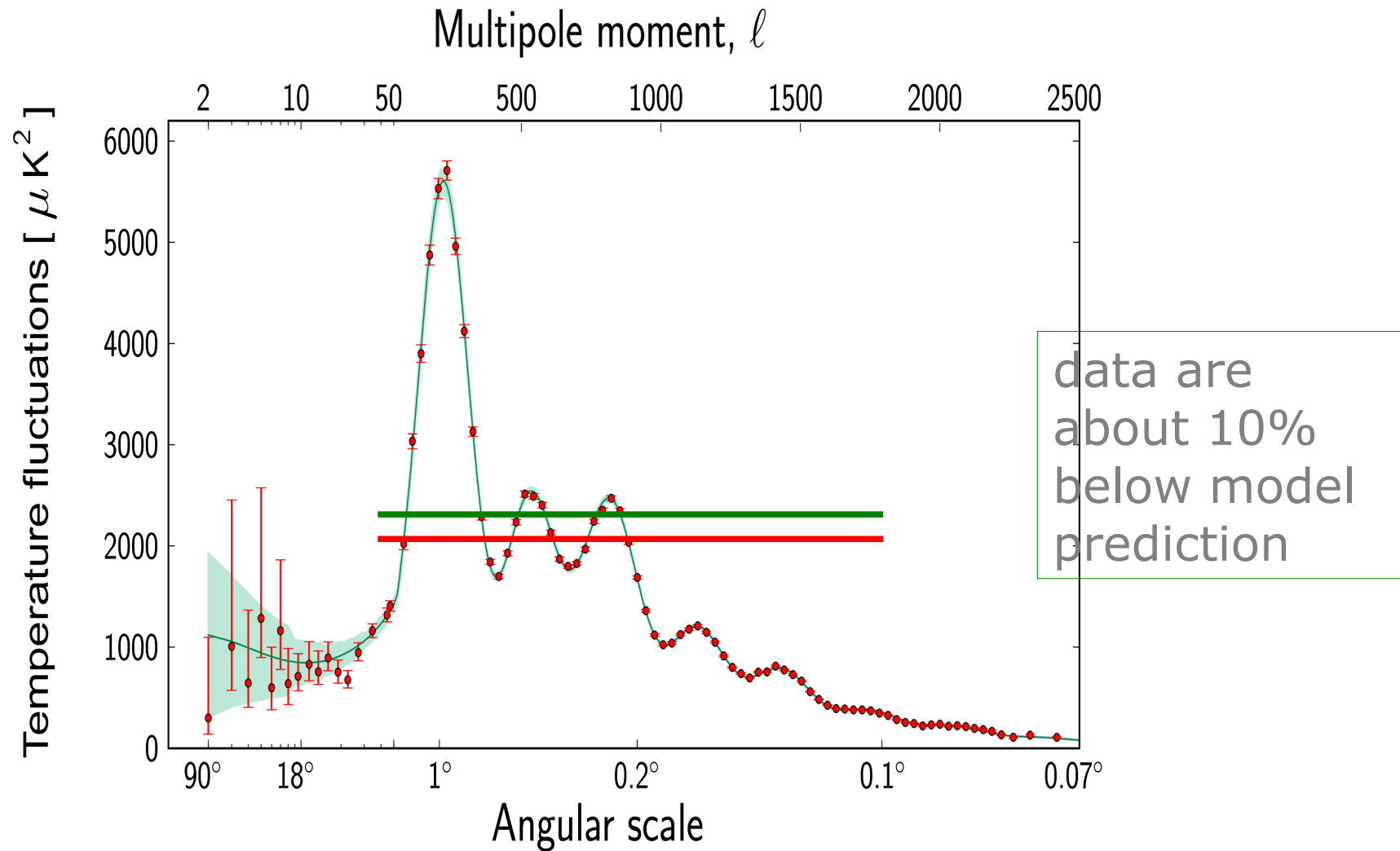
Hemisphere Asymmetry

„Cold Spot“

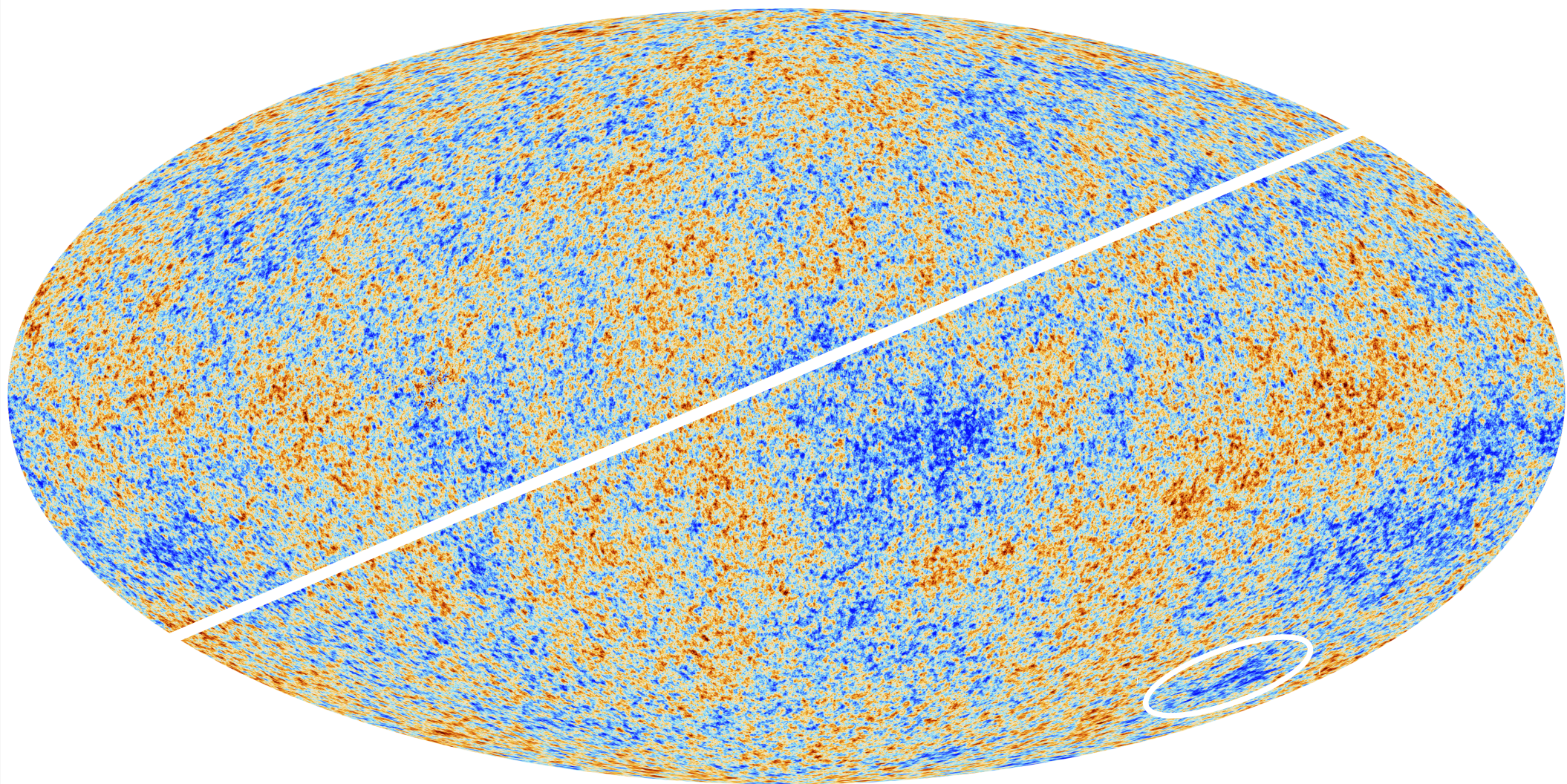
# „Missing Power“



# „Missing Power“



# Hemisphere Asymmetry and „Cold Spot“



# Anomalies

Planck confirms: anomalies previously seen by WMAP are CMB features!

*But what is their meaning?*

Signatures of new physics

(Universe inhomogeneous on largest scales,  
view on beginning of inflation, ...)

or simply statistical fluctuations?

(with the large number of tests applied,  
a few outliers are to be expected)



# Conclusions

Planck provided us with an amazingly consistent cosmological picture

Planck confirmed the quantum fluctuation origin of the cosmic structures

A few discrepancies and anomalies remain, but wouldn't life be boring otherwise?

The scientific results that we present today are the product of the Planck Collaboration, including individuals from more than 50 scientific institutes in Europe, the USA and Canada

